

AR1 (spectral line) commissioning plan

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Engineering verification/commissioning

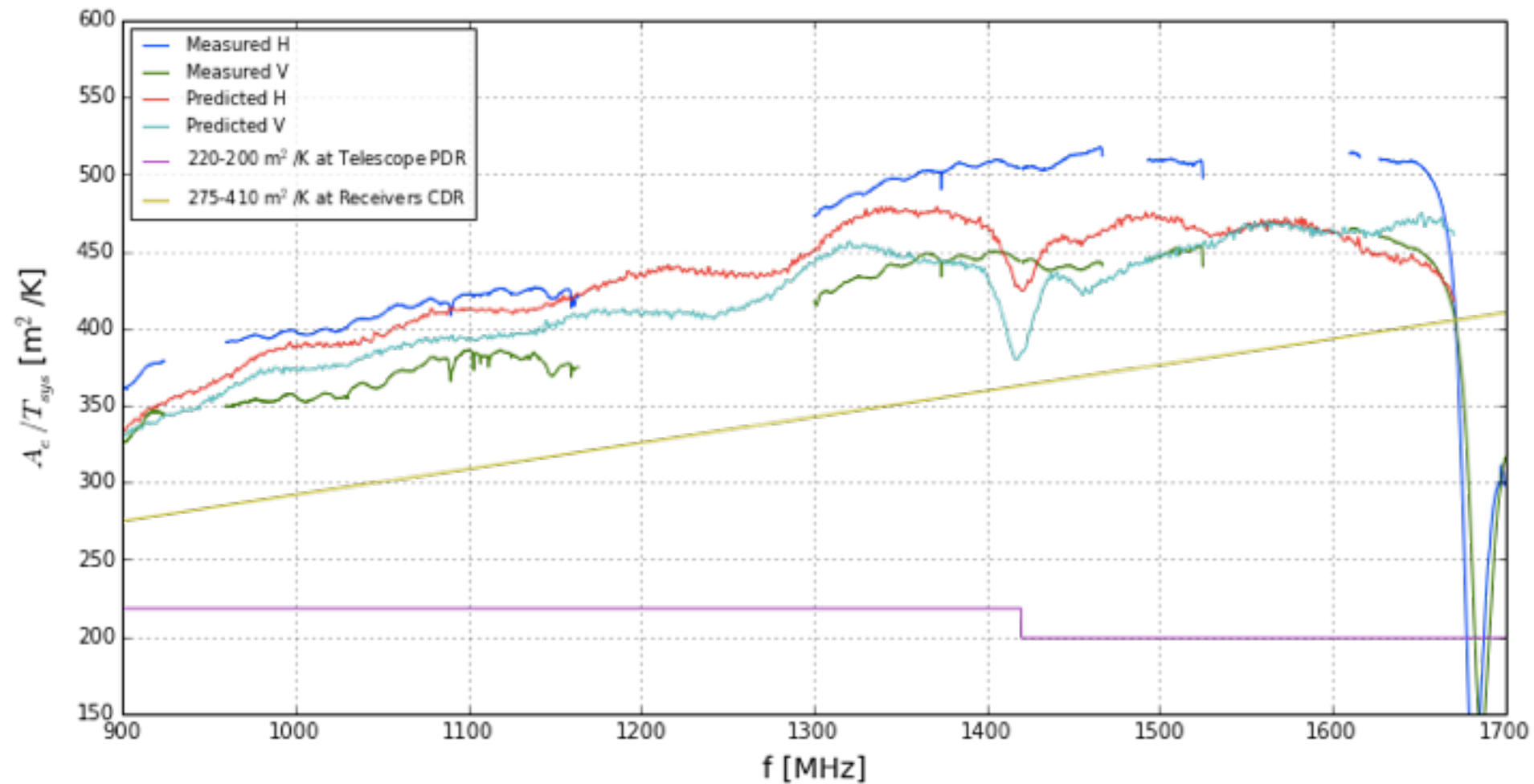


- » Receptor test system (RTS)
 - Acceptance testing of antennas after handover
 - 4 antenna inputs
 - 4K & 32 K channels
 - L band, 2x Ku test receivers
 - Single dish & interferometric tests

- » Antennas then handed over to AR1

- » Currently have M024, M025, M031

Measured sensitivity



Array release 1



- » 16 dishes
 - » 4K (209 kHz) and 32K* (26 kHz) channelisation
 - » Tied array
 - » Basic sub-array function
-
- » Currently have M062, M063 (~2 km baseline)

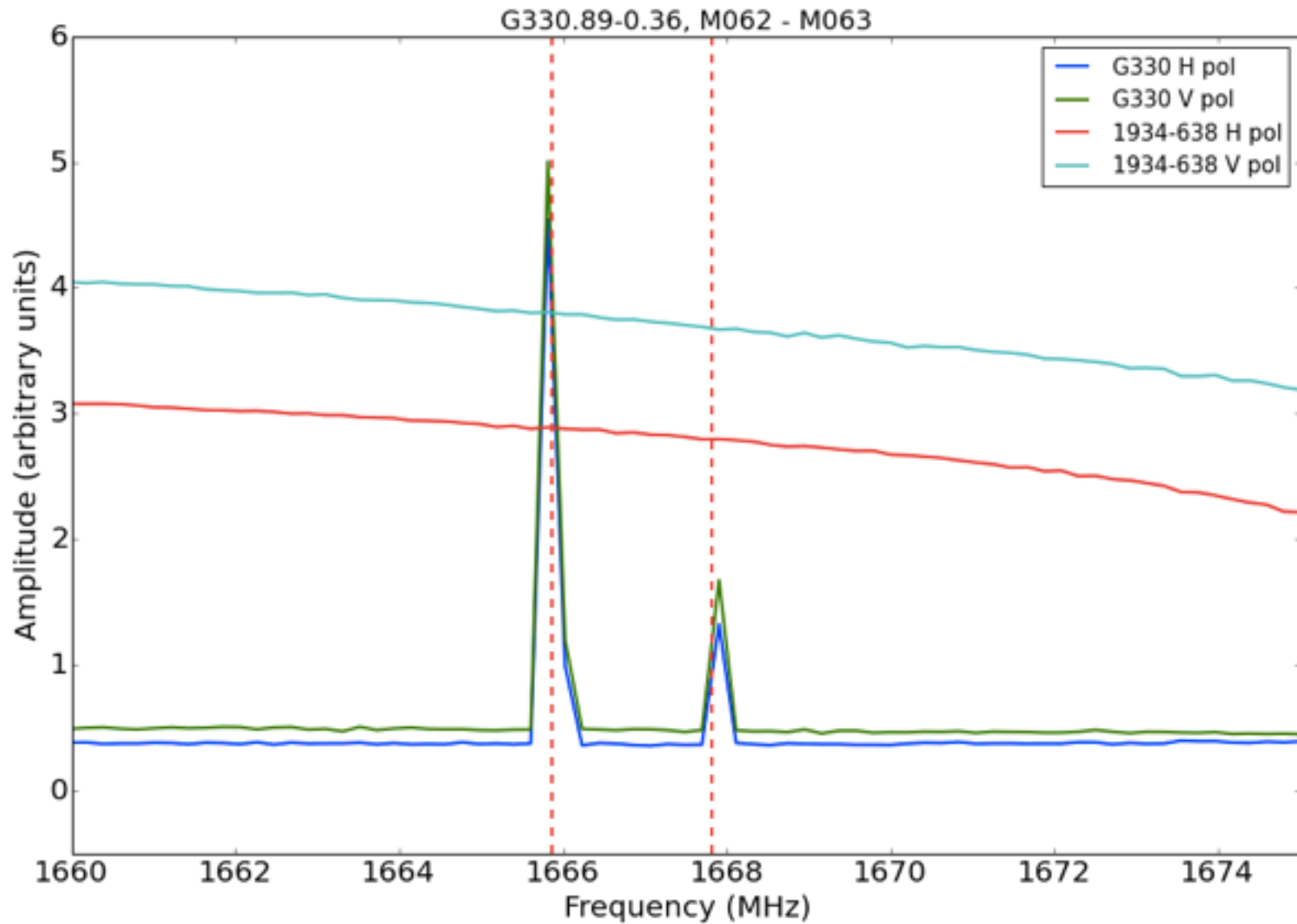
* Not yet available

Basic tests

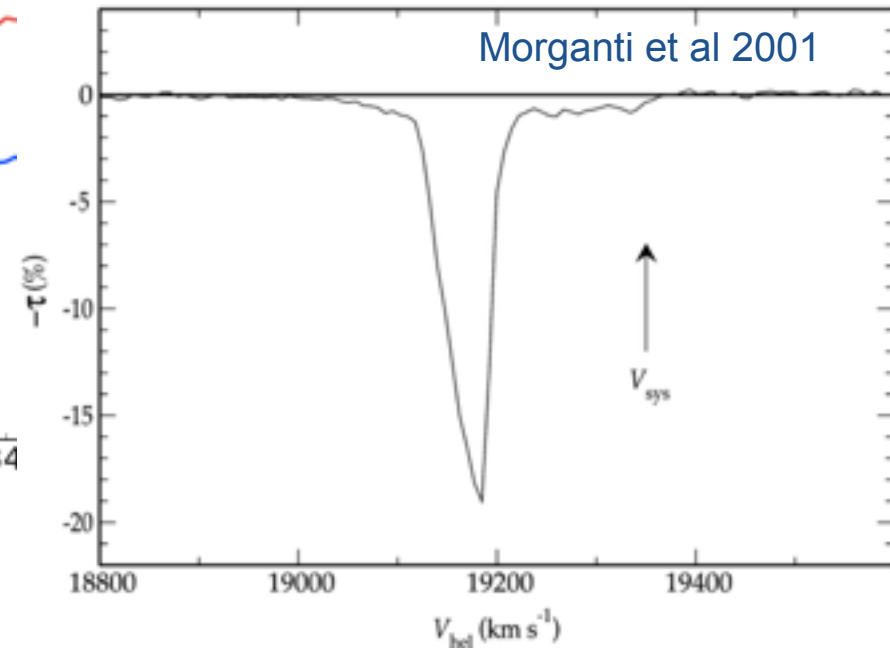
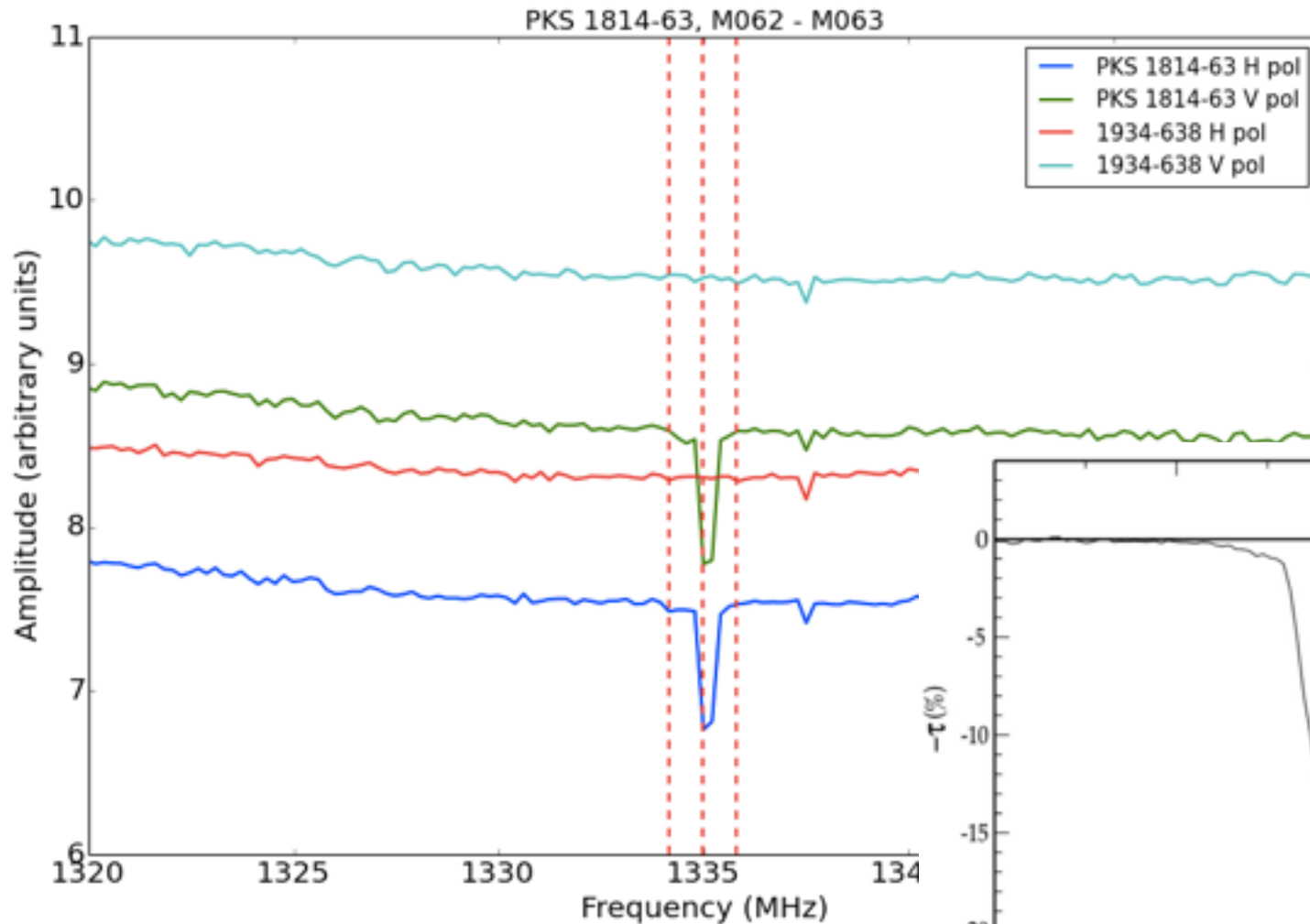


- » Single dish
 - Pointing
 - Tipping curve
 - Flux calibration
 - Gain stability
 - Spectral baseline
 - RFI
- » Single-baseline
 - » Verify fringes
 - » Verify stopped fringes/ delay tracking
 - » Verify strong spectral lines detected

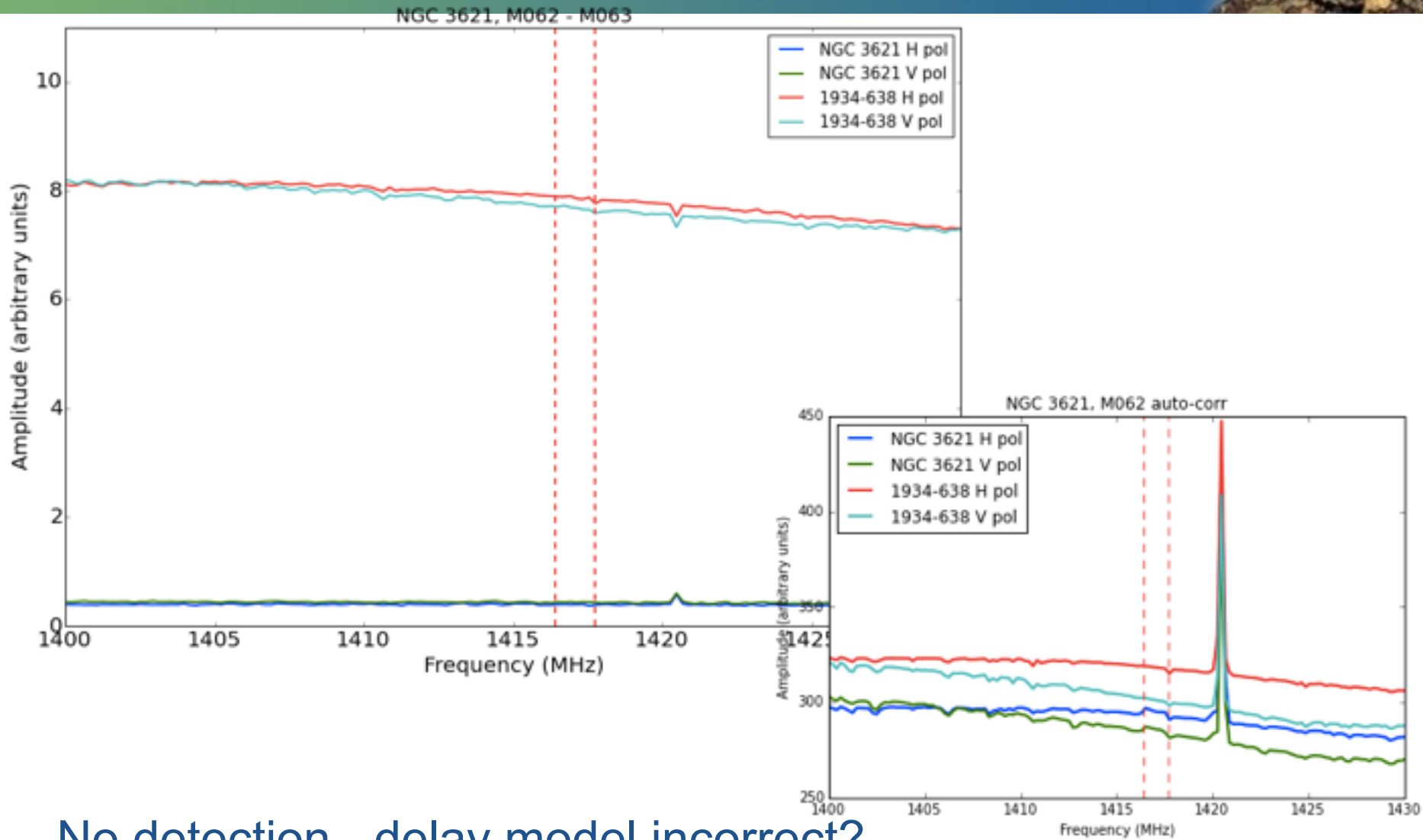
OH maser (M062 - M063)



HI absorption PKS 1814-63 (M062-M063)



HI emission NGC 3621



No detection - delay model incorrect?

(min 4 antennas)

- » Phase & amplitude closure
- » Baseline calibration
 - track strong unresolved continuum sources
 - confirm antenna positions and determine delays
- » Delay tracking
 - track strong unresolved sources
 - verify phase stability and coherence across channels and in different pointing directions



Calibration



- » Bandpass and delay stability
- » Gain stability
- » Flux-scale calibration
- » Polarisation stability
- » Direction dependent effects
 - Field with strong point sources across FOV (J1915-7439)
 - antenna pointing
 - polarisation

Spectral line requirements (full array)



- » The telescope shall achieve limiting sensitivity in Stokes I of $\text{RMS} < 0.8 \text{ mJy}$ in 5 km/s channels after only 2 hours of integration, when observing HI or OH lines across the entire frequency range
- » After (self) calibration, **line-to-continuum** dynamic range of 60 dB or more must be achievable in L-band, across the -3 dB FoV.
- » After (self) calibration, **line-to-line** dynamic range of 40 dB or more must be achievable in L-band, across the -3 dB FoV.
- » On applying natural weighting the L-band sensitivity of the interferometer shall change by no more than 25% for angular scales from 8" to 80" for declinations $< +10^\circ$ and tracks of ≥ 2 hour duration.

Test fields (AR1)



- » Strong continuum sources (overlap with continuum commissioning)
- » PKS 1814-63 (10 Jy continuum, 2.3 Jy absorption)
- » G330.89-0.36 (OH maser, 800 Jy peak, sub-Jy maser sources in field of view, extended continuum, weak absorption)
- » NGC 3621 (very bright HI galaxy)
- » NGC 5236 / M83 (extended HI)
- » IC 1459 (faint extended HI, mosaicing?)
- » Cen A (extended strong continuum + HI emission)