

Where the Action is

Hamilton's principle provides a powerful tool in the classical dynamics of particles and even more so in 'classical' field theory.

In the former one has generalized coordinates $q^i(t)$, the index "i" labeling different components and/or particles; in the latter relativistic considerations require the generalized coordinates to be $\phi^i(x)$, functions of x^μ carrying an index "i"

- to distinguish different components and/or fields. For particles the generalized velocities are $\dot{q}^i = dq^i/dt$; time is the 0-th component of a 4-vector so for relativistic fields the analogue is $\partial_\mu \phi^i$. Particle mechanics says: take a Lagrangian $L(q, \dot{q})$ and integrate it over time to get the action S , a real number. Field theory says: for m
- a Lagrangian density $\mathcal{L}(\phi, \partial\phi)$ which is a Lorentz scalar and integrate it over the 4-volume $d^4x = dt dx dy dz$ to form

$$S[\phi] = \int d^4x \mathcal{L}(\phi, \partial\phi) \quad (4.1)$$

Notationally: S is a 'functional' of ϕ , a mapping from functions to numbers - in other words its value depends on ϕ^i at every space-time point.

Hamilton's principle for particles deals with infinitesimal variations $q'(t) \rightarrow q'(t) + \delta q'(t)$. Field theory allows us to change both the generalized coordinates and the space time ones

$$x^\mu \rightarrow x'^\mu = x^\mu + \delta x^\mu, \quad \phi^i(x) \rightarrow \phi^i(x') = \phi^i(x) + \delta \phi^i(x) \quad (4.2)$$

It is useful to have a 'local variation' which compares things with the same coordinates - compare (2.78) - so

$$\delta \phi^i(x) = \phi^i(x') - \phi^i(x) = \Delta \phi^i(x) + \delta x^\mu \partial_\mu \phi^i(x) \quad (4.3)$$

to first order. The difference of $\delta \phi^i$ and $\Delta \phi^i$ is important for

$$\begin{aligned} \delta (\partial_\mu \phi^i(x)) &\equiv \partial'_\mu \phi^i(x') - \partial_\mu \phi^i(x) \\ &= \partial_\mu \Delta \phi^i(x) + \partial_\mu (\delta x^\nu \partial_\nu \phi^i(x)) - (\partial_\mu \delta x^\nu) \partial_\nu \phi^i(x) \end{aligned} \quad (4.4)$$

Since

$$d^4 x' = \det \left(\frac{\partial x'}{\partial x} \right) d^4 x = (1 + \partial_\mu \delta x^\mu) d^4 x \quad (4.5)$$

one finds

$$\begin{aligned} \delta S[\phi] &= \int d^4 x \left[(\partial_\mu \delta x^\mu) \mathcal{L} + \frac{\partial \mathcal{L}}{\partial \phi^i} (\Delta \phi^i + \delta x^\mu \partial_\mu \phi^i) + \right. \\ &\quad \left. \frac{\partial \mathcal{L}}{\partial (\phi^i_{,\mu})} (\partial_\mu \Delta \phi^i + \delta x^\nu \partial_\mu \partial_\nu \phi^i) \right] \\ &= \int d^4 x \left[\left(\frac{\partial \mathcal{L}}{\partial \phi^i} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\phi^i_{,\mu})} \right) \Delta \phi^i - \partial_\mu J^\mu \right], \end{aligned} \quad (4.6)$$

$$J^\mu \equiv - \frac{\partial \mathcal{L}}{\partial (\phi'^i{}_{,\mu})} \Delta \phi^i - \delta X^\mu \mathcal{L} \quad (4.7)$$

Thus $\delta S = 0$ provides us with the Euler - Lagrange field equations

$$\partial_\mu \frac{\partial \mathcal{L}}{\partial (\phi'^i{}_{,\mu})} - \frac{\partial \mathcal{L}}{\partial \phi^i} = 0 \quad (4.8)$$

and a conserved current (Noether current)

$$\partial_\mu J^\mu = 0 \quad (4.9)$$

plus a conserved (Noether) "charge"

$$Q = \int d^3x J^0, \quad \frac{d}{dt} Q = 0 \quad (4.10)$$

if the fields fall off sufficiently fast at spatial infinity.

Let's see how this works for a pertinent case, that of a complex (Lorentz) scalar field Φ . One could work with the real and imaginary parts, $\Phi = (\phi' + i\phi'')/\sqrt{2}$ but we'll stick with Φ and let "i" include the independent complex conjugate Φ^* . We want $\mathcal{L}$ real and so take

$$\begin{aligned} \mathcal{L} &= g^{\mu\nu} (\partial_\mu \Phi^*) (\partial_\nu \Phi) - V(\Phi^* \Phi) = (\partial_\mu \Phi^*) (\partial^\mu \Phi) - V(|\Phi|^2) \\ &= |\partial \Phi|^2 - V \end{aligned} \quad (4.11)$$

indicating the range of notations. Then

$$\frac{\partial \mathcal{L}}{\partial \Phi} = -V' \Phi, \quad \frac{\partial \mathcal{L}}{\partial \Phi^*} = -\Phi^* V', \quad (4.12)$$

$$\frac{\partial \mathcal{L}}{\partial (\Phi_{,\mu})} = \partial^\mu \Phi, \quad \frac{\partial \mathcal{L}}{\partial \Phi_{,\mu}^*} = \partial^\mu \Phi^*,$$

$V' = V'(|\Phi|^2) = \partial V(|\Phi|^2) / \partial |\Phi|^2$, and we get two field equations

$$\square \Phi + V' \Phi = 0, \quad \square \Phi^* + \Phi^* V' = 0 \quad (4.13)$$

If $V = m^2 |\Phi|^2$ (4.13) are just Klein-Gordon equations. Under space-time translations $\delta X^\mu = \xi^\mu$, a constant 4-vector $\delta \Phi$ vanishes but

$$\Delta \Phi = -\xi^\mu \partial_\mu \Phi \quad (4.14)$$

does not. The conserved current reads

$$\begin{aligned} J^\mu &= (\partial^\mu \Phi^*) \xi^\nu (\partial_\nu \Phi) + (\xi^\nu \partial_\nu \Phi^*) (\partial^\mu \Phi) - \xi^\mu \mathcal{L} \\ &= T^\mu{}_\nu \xi^\nu, \end{aligned} \quad (4.15)$$

whence four associated charges written

$$P^\mu = \int d^3x T^{0\mu} \quad (4.16)$$

Examining P^0 ,

$$\begin{aligned} P^0 &= \int d^3x [|\partial_0 \Phi|^2 + |\nabla \Phi|^2 + V(|\Phi|^2)] \\ &= \int d^3x \mathcal{H}, \end{aligned} \quad (4.17)$$

we see the analogue of the energy or hamiltonian in particle mechanics - indeed P^i are the 'three momenta' of the field(s). Note the 'stress energy tensor' $T_{\mu\nu}$ is symmetric.

Lorentz invariance of (4.11) gives six more generalized charges; we leave this as an exercise. There is yet another symmetry of $\mathcal{L}$: it is unchanged by 'global phase transformations' $\Phi \rightarrow e^{i\alpha} \Phi$, $\Phi^* \rightarrow \Phi^* e^{-i\alpha}$ with α constant. In that case

$$\delta \Phi = \Delta \Phi = i\alpha \Phi, \quad (4.18)$$

$\delta X^\mu = 0$ and the conserved current is

$$J^\mu = i\alpha(\Phi^* \partial^\mu \Phi - (\partial^\mu \Phi^*) \Phi) \equiv i\alpha \Phi^* \overleftrightarrow{\partial}^\mu \Phi. \quad (4.19)$$

Now, it must be possible to design an $\mathcal{L}$ whose Euler-Lagrange field equations are those of electrodynamics*. The basic building block is $F^{\mu\nu}$ in (2.19) and Lorentz invariance says we need $F_{\mu\nu} F^{\mu\nu}$ and $A^\mu J_\mu$. Thus try

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - A^\mu J_\mu \quad (4.20)$$

with the μ on A^μ like the i on Φ^i . Since

$$\frac{\partial \mathcal{L}}{\partial A_{\nu,\mu}} = -\frac{1}{2} F^{\mu\nu} + \frac{1}{2} F^{\nu\mu} = -F^{\mu\nu}, \quad \frac{\partial \mathcal{L}}{\partial A_\nu} = -J^\nu \quad (4.21)$$

* If not you wouldn't be reading this!

the relative sign is correct but what about the absolute sign? Under translations $\Delta A^\mu = -\xi^\nu \partial_\nu A^\mu$ and our recipe gives the 'canonical' stress-energy tensor ($J^\mu = 0$)

$$T_{can}^{\mu\nu} = F^{\mu\alpha} \partial^\nu A_\alpha - g^{\mu\nu} \mathcal{L} \quad (4.22)$$

This is a bit of an embarrassment: $T_{can}^{\mu\nu}$ isn't symmetric or even gauge invariant! Still, the field equations are unaltered if we modify $\mathcal{L}$ by a 4-divergence but the Noether current is changed; by a judicious choice we can arrange this so

$$T^{\mu\nu} = F^{\mu\alpha} F^\nu{}_\alpha - g^{\mu\nu} \mathcal{L} \quad (4.23)$$

Using the result of Problem 2.1, $T^{00} = (\vec{E}^2 + \vec{B}^2)/2 \geq 0$ so the overall sign is right.

Spinor fields can be treated similarly:

$$\mathcal{L} = \bar{\Psi} (i\not{D} - m) \Psi, \quad (4.24)$$

$$\not{D} \equiv \gamma^\mu \partial_\mu \quad (4.25)$$

yields up the free Dirac equation (3.12) as $\partial \mathcal{L} / \partial \bar{\Psi}_{,\mu} = 0$. Integration by parts gives $\partial \mathcal{L} / \partial \Psi_{,\mu} = 0$ and the conjugate

$$\bar{\Psi} (i\not{D} + m) = i\partial_\mu \bar{\Psi} \gamma^\mu + \bar{\Psi} m = 0 \quad (4.26)$$

Taking

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\Psi} (i\partial - eA - m) \Psi \quad (4.27)$$

produces the Dirac equation in an electromagnetic field, (3.33), and the electrodynamics field equations with the source current

$$J^\mu = e \bar{\Psi} \gamma^\mu \Psi. \quad (4.28)$$

This raises an important point about the Dirac field: because the charges are determining the field in which they move in (4.27) one expects $\mathcal{L}$ to be the same: if we replace all the fields by their charge conjugates, including $A^{\mu c} = -A^\mu$

$$\mathcal{L}^c = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\Psi}^c (i\partial + eA - m) \Psi^c \quad (4.29)$$

but of course $J^{\mu c} = -J^\mu$. Note that in the Weyl representation

$$\begin{aligned} \gamma^0 &= \gamma^{0T} = \gamma^{0*}, \quad \gamma^1 = -\gamma^{1T} = \gamma^{1*}, \\ \gamma^2 &= \gamma^{2T} = -\gamma^{2*}, \quad \gamma^3 = -\gamma^{3T} = \gamma^{3*} \end{aligned} \quad (4.30)$$

and (3.38) can be written

$$\Psi^c = -i\gamma^2\gamma^0\bar{\Psi}^T \quad (4.31)$$

so

$$\bar{\Psi}^c = \Psi i\gamma^2\gamma^0 = \Psi^T i\gamma^0\gamma^2 \quad (4.32)$$

Since

$$(i\gamma^0\gamma^2)(\gamma^\mu, 1)(-i\gamma^2\gamma^0) = (-1, \gamma^{\mu T}) \quad (4.33)$$

in fact

$$\mathcal{L}_c = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \psi^\dagger \gamma^\mu (i\partial_\mu + eA_\mu) \bar{\psi} + m \psi^\dagger \bar{\psi}, \quad (4.34)$$

$$j^\mu = e \psi^\dagger \gamma^\mu \bar{\psi} \quad (4.34b)$$

The resolution of this paradox is that the Dirac field does not commute like an ordinary c-number, even classically, but anticommutes; eg putting in the spinor indices

$$j^\mu = e \psi_i (\gamma^\mu)_{ij} \bar{\psi}_j = -e \bar{\psi}_j (\gamma^\mu)_{ij} \psi_i \quad (4.35)$$

Thus the Dirac fields are 'anticommuting c-numbers' or elements of a 'Grassmann algebra'. This is the origin of Fermi-Dirac statistics and the exclusion principle!

References

Classical field theory is discussed in eg H. Goldstein "Classical Mechanics", Addison-Wesley, Reading, Massachusetts, 1980; beware though the ict (pronounced yuch) metric. Our treatment follows sections 6, 4, 5 of Cushing, *ibid.*. The peculiarities of spinor fields is discussed in Ramond, *ibid.*

Problems

- 4.1. Verify eqs (4.4) and (4.5).
- 4.2. Show (4.10) follows from (4.9).
- 4.3. For the model (4.11) obtain the Noether current associated to infinitesimal Lorentz transformations (2.7) as

$$J^{\mu} = \frac{1}{2} M^{\alpha\beta} \omega_{\alpha\beta}$$

Show the space-space components of

$$Q^{\alpha\beta} = \int d^3x M^{0\alpha\beta}$$

give the angular momentum of the field

- 4.4. Determine the 4-divergence of f^{μ} which must be added to (4.20) to yield (4.23).