



The Color Glass Condensate

QCD at modern collider facilities

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Outline

- 1 Motivation: gluons form the CGC
 - Background information on the standard model
 - Current and planned collider experiments
 - Enhanced gluon production at high energies
 - CGC: why the name
- 2 JIMWLK evolution: properties of the CGC
 - Gluons in observables
 - The evolution equation
 - The saturation scale
- 3 A sample experiment
 - Geometric scaling @ HERA
- 4 Getting quantitative
 - NLO corrections
 - HERA fits
- 5 Applications and outlook

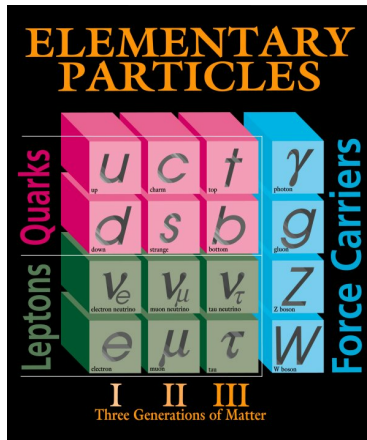
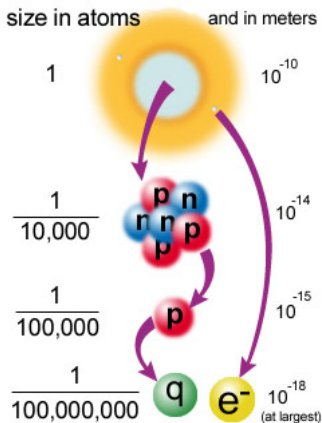


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From atoms to the standard model



+Higgs



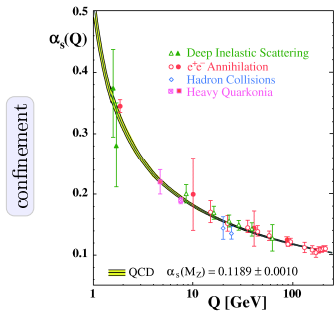
QCD: the strong interaction

■ Focus on QCD:



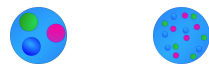
+Higgs

■ Running coupling



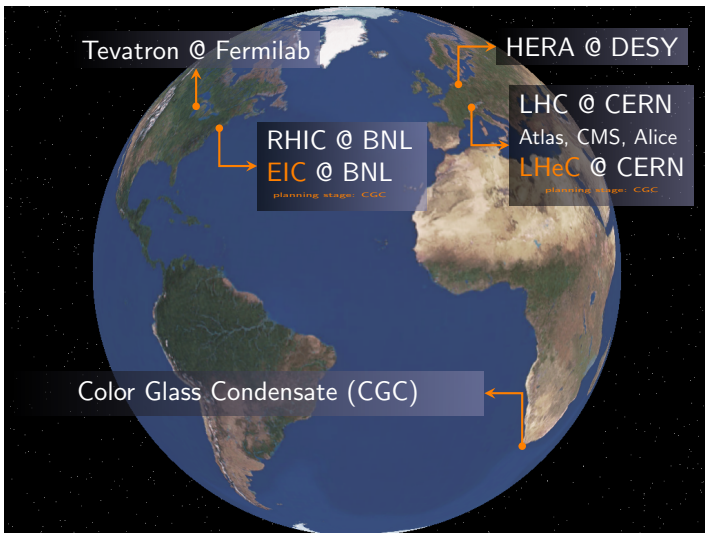
confinement

asymptotic freedom



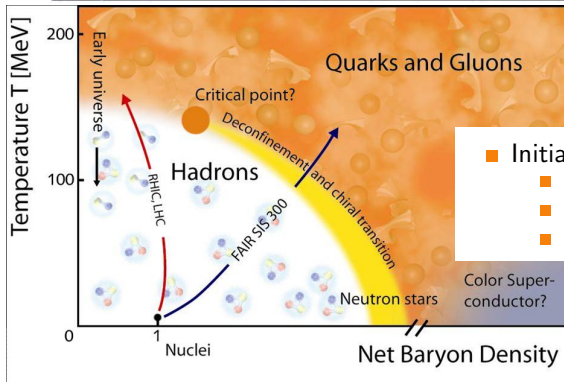
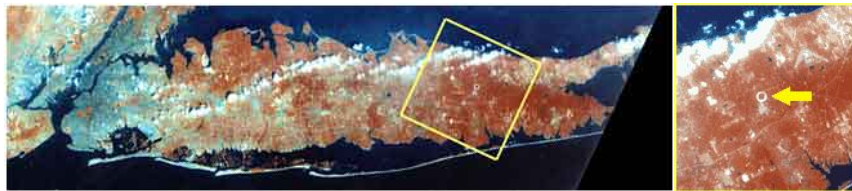


High energy physics viewed from UCT





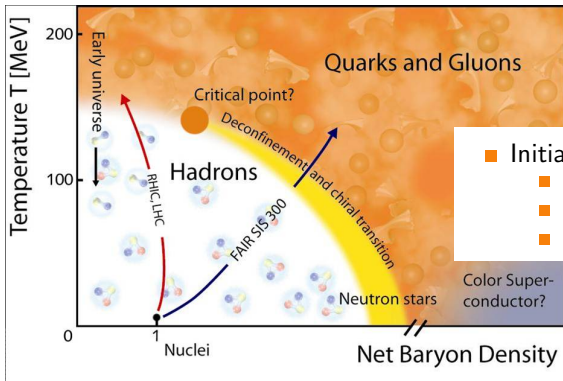
The Quark Gluon Plasma at RHIC and LHC



- Initial to final state via
- thermalization (times!)
- initial hard particle production
- ...



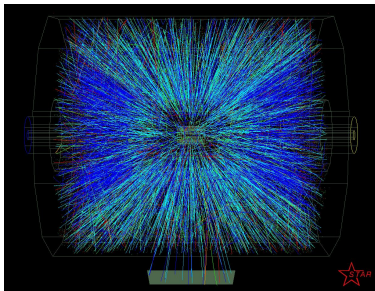
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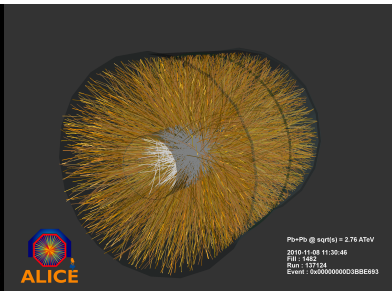
- Initial to final state via
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 - initial hard particle production
 - ...



The Quark Gluon Plasma at RHIC and LHC



RHIC event (STAR), side view



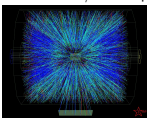
LHC event (ALICE)



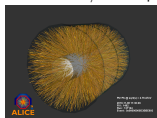
Particle production at modern colliders

- Large amounts of energy available: 200-14000 m_{proton}
- heavy ions @ RHIC & LHC: QGP

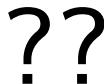
RHIC @ 200GeV/nucleon pair



LHC @ 2.76TeV/nucleon pair



LHC @ 4TeV/nucleon pair



- QCD drives particle production
 - serious backgrounds for particle searches
 - new physics phenomena
 - copious gluon production

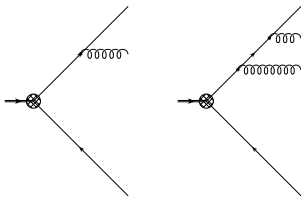


Color Glass Condensate
CGC



Energy dependence: from photons to gluons

■ photon-like contributions



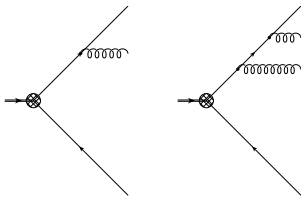
■ enhanced by phase space integrals $\frac{dE}{E} \frac{d\theta}{\theta} \rightarrow \alpha_s \ln E \ln \theta$

■ all orders calculation needed $\sum_{n=0}^{\infty} (\alpha_s \ln E)^n \dots$

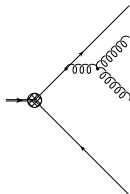


Energy dependence: from photons to gluons

■ photon-like contributions



■ QCD: **charged** gluons



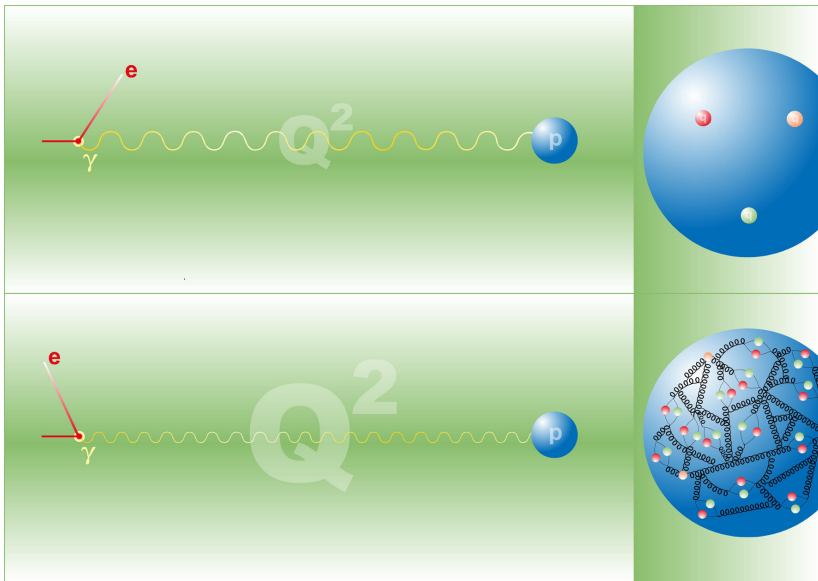
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■ all orders calculation needed $\sum_{n=0}^{\infty} (\alpha_s \ln E)^n \dots$

■ gluons **charged** \rightarrow radiation **nonlinear** in QCD



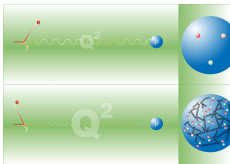
Example: ep at HERA





Example: ep at HERA

- Q^2 determines the resolution



$$Q^2 := -q^2 \gg 0$$

spacelike!

transverse resolution

$$\Delta r \sim \frac{1}{Q}$$

- $\ln E$ comes with many aliases:



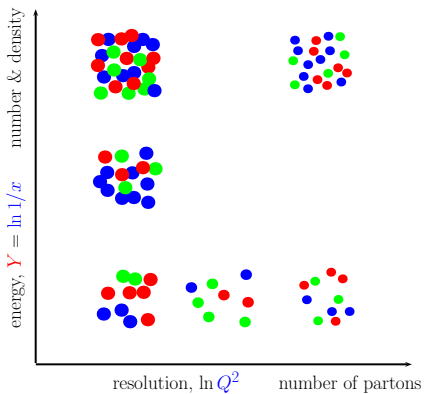
$$Y = \ln \frac{1}{x} = \ln E$$

all used
synonymously

- Y rapidity
- $x = x_{Bj} := \frac{Q^2}{2p \cdot q} = \frac{Q^2}{2m E}$ Bjorken x

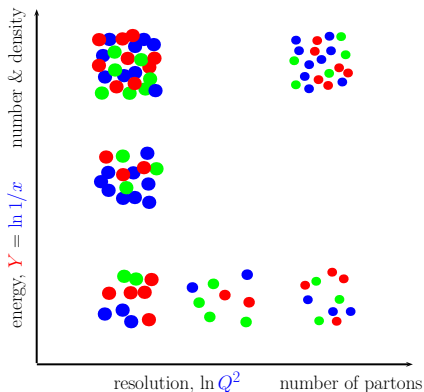


Large energies mean large densities





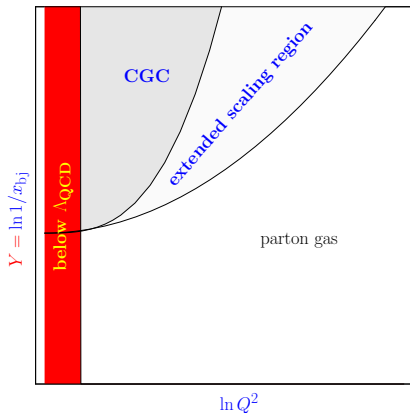
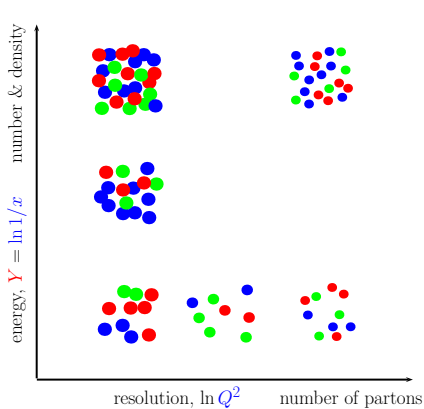
Large energies mean large densities



- density \rightarrow nonlinear effects
- finite correlation length R_s



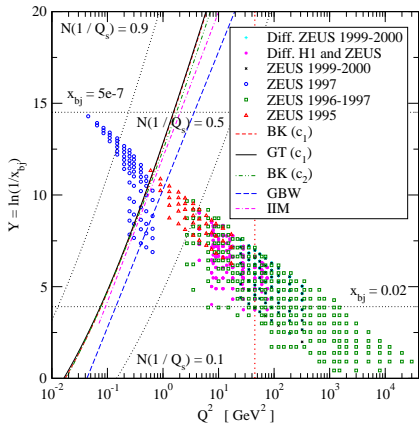
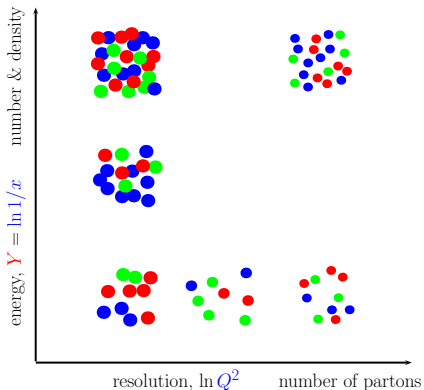
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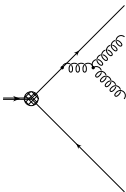
- Real world example: HERA ep



Why the name?

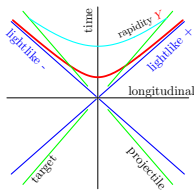
Color

- QCD
- and QCD only!
quarks and gluons



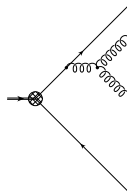
Glass

- fields evolve slowly compared to natural time scales
- time scales
energy, time dilation



Condensate

- phase space density
 $\sim \frac{1}{\alpha_s}$ & saturates
- density
energy, gluons charged





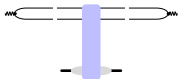
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Total cross section (zeroth order in $\alpha^m (\alpha_s \ln(1/x))^n$)

$$\sigma_{\text{DIS}}(Y, Q^2) = 2lm$$





Total cross section (zeroth order in $\alpha^m (\alpha_s \ln(1/x))^n$)

$$\sigma_{\text{DIS}}(Y, Q^2) = 2\text{Im} \int d^2r |\psi^2|(r^2, Q^2)$$

photon wave functions/impact factor



Total cross section (zeroth order in $\alpha^m (\alpha_s \ln(1/x))^n$)

$$\sigma_{\text{DIS}}(Y, Q^2) = 2\text{Im} \int d^2r |\psi^2|(r^2, Q^2) 2 \int d^2b \left\langle \frac{\text{tr}(1 - U_{\mathbf{x}} U_{\mathbf{y}}^\dagger)}{N_c} \right\rangle(Y)$$

photon wave functions/impact factor \rightarrow $|\psi^2|(r^2, Q^2)$
 $\sigma_{\text{dipole}} \rightarrow$ $\left\langle \frac{\text{tr}(1 - U_{\mathbf{x}} U_{\mathbf{y}}^\dagger)}{N_c} \right\rangle(Y)$



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photon wave functions/impact factor \rightarrow $|\psi^2|(r^2, Q^2)$
 $\sigma_{\text{dipole}} \rightarrow$ $2 \int d^2b \left\langle \frac{\text{tr}(1 - U_x U_y^\dagger)}{N_c} \right\rangle(Y)$

- σ_{dipole} contains U_x



Total cross section (zeroth order in $\alpha^m (\alpha_s \ln(1/x))^n$)

$\sigma_{\text{DIS}}(Y, Q^2) = 2 \text{Im}$

photon wave functions/impact factor

$$= \int d^2 r |\psi^2|(r^2 Q^2) \underbrace{2 \int d^2 b \left\langle \frac{\text{tr}(1 - U_x U_y^\dagger)}{N_c} \right\rangle(Y)}_{\sigma_{\text{dipole}}}$$

σ_{dipole}

- σ_{dipole} contains U_x

$\langle \dots \rangle(Y)$ difficult:

- target wavefunction is non-perturbative



Total cross section (zeroth order in $\alpha^m (\alpha_s \ln(1/x))^n$)

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$$\frac{d}{dY} \langle \dots \rangle(Y)$$



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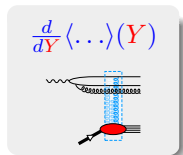
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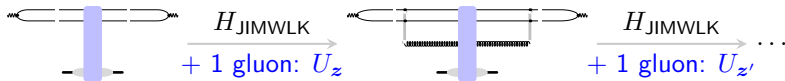
- target wavefunction is non-perturbative



- Bookkeeping device: $\langle \dots \rangle(Y) = \int \hat{D}[U] \dots \hat{Z}_Y[U]$

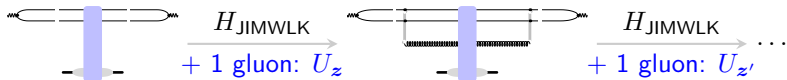


The JIMWLK evolution equation





The JIMWLK evolution equation



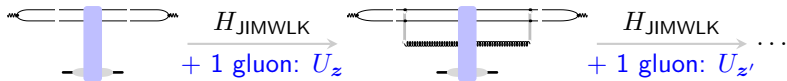
$$\bullet \frac{d}{dY} Z_Y[U] = -H_{\text{JIMWLK}}[U] Z_Y[U]$$

Heribert Weigert Nucl. Phys. A703, 2002, 823

▶ explicit form



The JIMWLK evolution equation



$$\blacksquare \frac{d}{dY} Z_Y[U] = -H_{\text{JIMWLK}}[U] Z_Y[U]$$

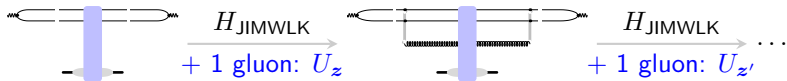
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▶ explicit form

- resums all $\sim [\alpha_s \ln(1/x)]^n$ (at LO)



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▶ explicit form

■ resums all $\sim [\alpha_s \ln(1/x)]^n$ (at LO)

■ **→** energy dependence of $\langle \dots \rangle(Y)$



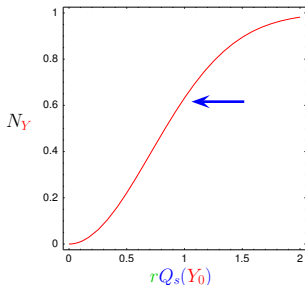
Saturation scale and cross section

$$\blacksquare \langle \dots \rangle(Y) \quad \longrightarrow \quad \left\langle \frac{\text{tr}(1 - U_{\mathbf{x}} U_{\mathbf{y}}^\dagger)}{N_c} \right\rangle(Y) =: N_Y(\mathbf{r})$$



Saturation scale and cross section

- $\langle \dots \rangle(Y) \rightarrow \left\langle \frac{\text{tr}(1 - U_{\mathbf{x}} U_{\mathbf{y}}^\dagger)}{N_c} \right\rangle(Y) =: N_Y(\mathbf{r})$
- qualitative expectation:



$$R_s(Y) \sim \frac{1}{Q_s(Y)}$$

$R_s(Y) \equiv$ correlation length

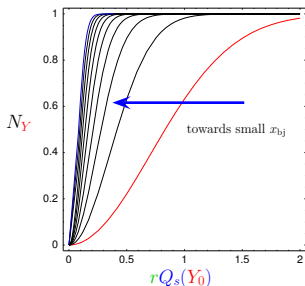
$Q_s(Y) \equiv$ saturation scale



Saturation scale and cross section

- $\langle \dots \rangle(Y) \xrightarrow{\text{red arrow}} \left\langle \frac{\text{tr}(1 - U_{\mathbf{x}} U_{\mathbf{y}}^\dagger)}{N_c} \right\rangle(Y) =: N_Y(\mathbf{r})$
- qualitative expectation:

correlation length shrinks:



$$R_s(Y) \sim \frac{1}{Q_s(Y)}$$

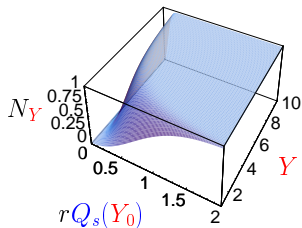
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JIMWLK: IR safety and scaling

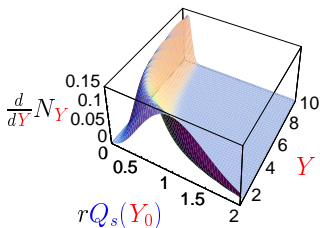
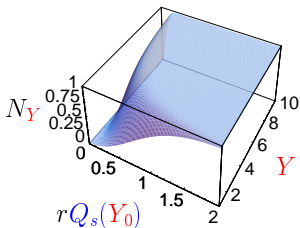
- Activity (new gluon production) near $Q_s(Y)$





JIMWLK: IR safety and scaling

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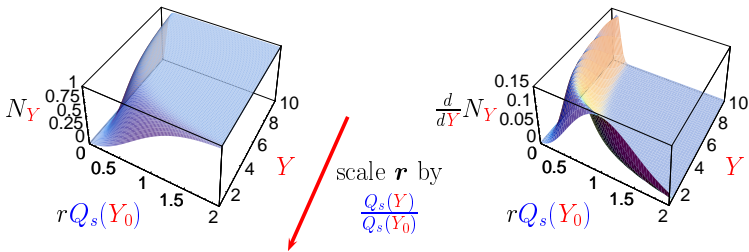


- activity follows $Q_s(Y)$
- IR safety **perturbative** ✓



JIMWLK: IR safety and scaling

- Activity (new gluon production) near $Q_s(Y)$



- activity follows $Q_s(Y)$
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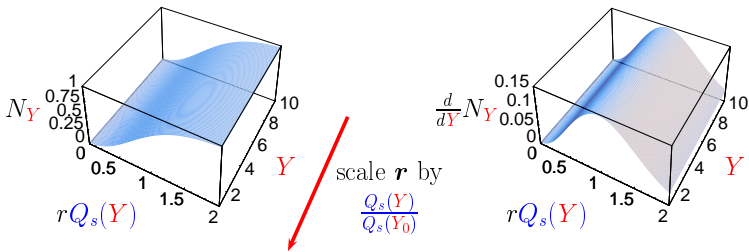
Detailed analysis:

scaling with $Q_s(Y)$
[persists approximately @ NLO]



JIMWLK: IR safety and scaling

- Activity (new gluon production) near $Q_s(Y)$



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Detailed analysis:

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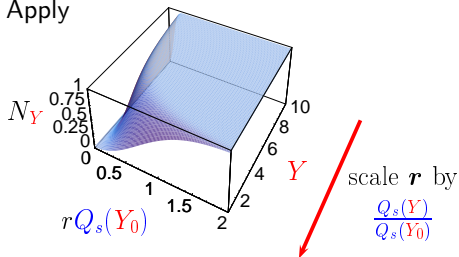
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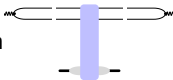
Geometric scaling @ HERA

Apply



to Hera

$$\sigma_{\text{DIS}}(Y, Q^2) = 2\ln$$

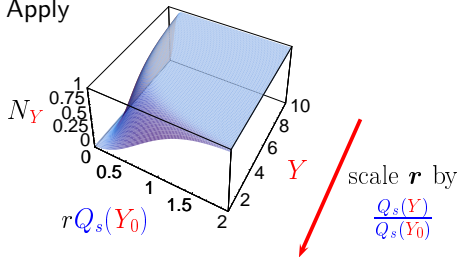


Golec-Biernat, Wüsthoff; PRD 60 (1999) 114023 [hep-ph/9903358]



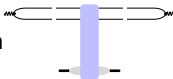
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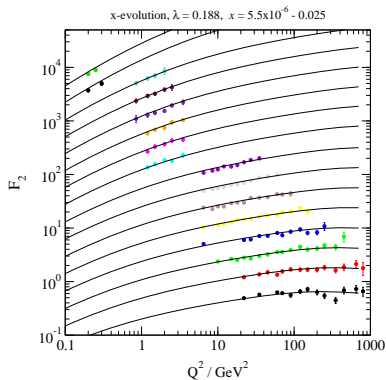
to Hera

$$\sigma_{\text{DIS}}(Y, Q^2) = 2\text{Im}$$



■ scaling fit to HERA:

$$\sigma(Y, Q^2) \sim F_2(Y, Q^2) \cdot Q^2$$

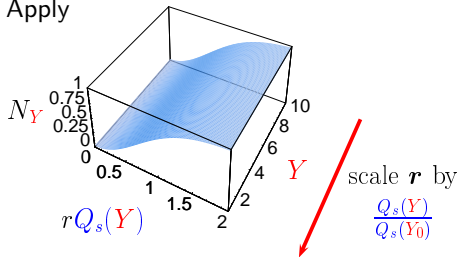


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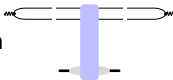
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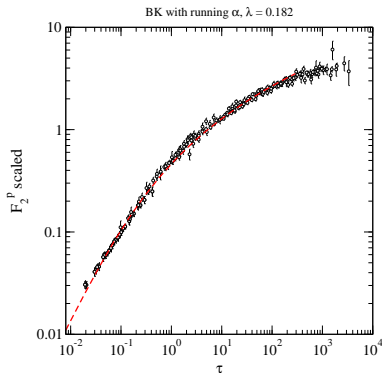
to Hera

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■ scaling fit to HERA:

$$\sigma(Y, Q^2) = \sigma(Y_0, \tau = Q^2 \frac{Q_s^2(Y_0)}{Q_s^2(Y)})$$

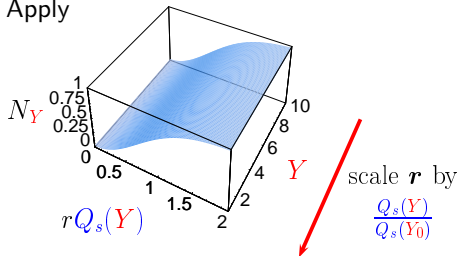


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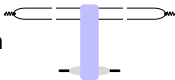
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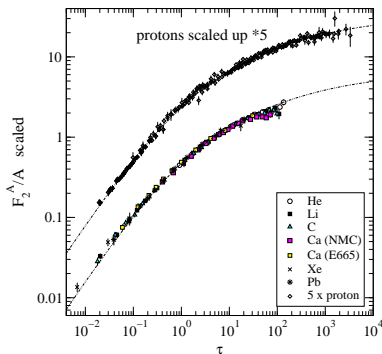
to Hera

$$\sigma_{\text{DIS}}(Y, Q^2) = 2lm$$



■ ... & with nuclei:

$$\sigma(Y, Q^2) = \sigma(Y_0, \tau = Q^2 \frac{(Q_s^A(Y_0))^2}{(Q_s^A(Y))^2})$$



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NLO-corrections

LO: $[\alpha_s \ln(1/x)]^n$; NLO: $[\alpha_s]^n [\ln(1/x)]^{n-1}$

- Corrections to evolution:

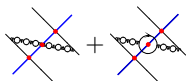
- Corrections to wave functions/impact factors



NLO-corrections

$$\text{LO: } [\alpha_s \ln(1/x)]^n; \quad \text{NLO: } [\alpha_s]^n [\ln(1/x)]^{n-1}$$

- Corrections to evolution:
 - running coupling



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov
Balitsky

- Corrections to wave functions/impact factors

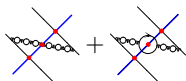


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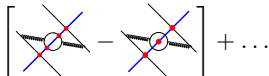
- Corrections to evolution:

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Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov
Balitsky

- new channels: quark/gluon-pair production (“conformal”)



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov

- Corrections to wave functions/impact factors

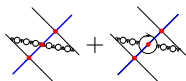


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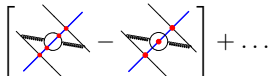
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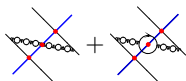


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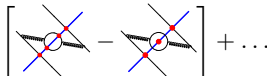
- Corrections to evolution:

- running coupling



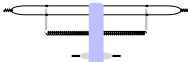
Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov
Balitsky

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Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov

- Corrections to wave functions/impact factors



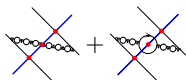


NLO-corrections

$$\text{LO: } [\alpha_s \ln(1/x)]^n; \quad \text{NLO: } [\alpha_s]^n [\ln(1/x)]^{n-1}$$

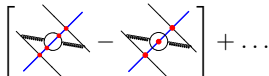
■ Corrections to evolution:

■ running coupling



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov
Balitsky

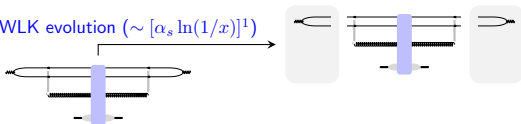
■ new channels: quark/gluon-pair production (“conformal”)



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov

■ Corrections to wave functions/impact factors

LO JIMWLK evolution ($\sim [\alpha_s \ln(1/x)]^1$)



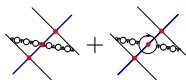


NLO-corrections

$$\text{LO: } [\alpha_s \ln(1/x)]^n; \quad \text{NLO: } [\alpha_s]^n [\ln(1/x)]^{n-1}$$

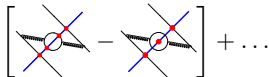
■ Corrections to evolution:

■ running coupling



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov
Balitsky

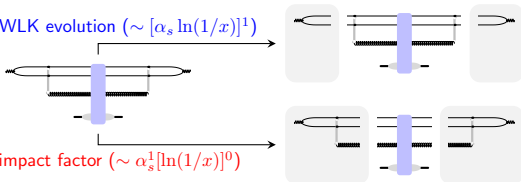
■ new channels: quark/gluon-pair production (“conformal”)



Gardi, Kuokkanen, Rummukainen, Weigert
Weigert, Kovchegov

■ Corrections to wave functions/impact factors

LO JIMWLK evolution ($\sim [\alpha_s \ln(1/x)]^1$)



NLO impact factor ($\sim \alpha_s^1 [\ln(1/x)]^0$)



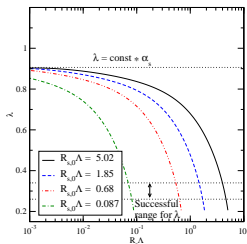
Balitsky, Chirilli
(2011)
not yet included



Effects of NLO-corrections

- NLO evolution: speed reduced

$$\lambda(Y) := \frac{d}{dY} \ln Q_s^2(Y)$$



LO JIMWLK

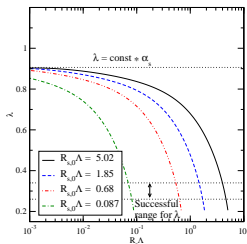
- too fast



Effects of NLO-corrections

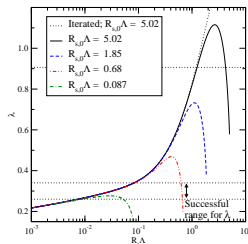
- NLO evolution: speed reduced

$$\lambda(Y) := \frac{d}{dY} \ln Q_s^2(Y)$$



LO JIMWLK

- too fast



+ running coupling

- remarkable slowdown
- fits become possible

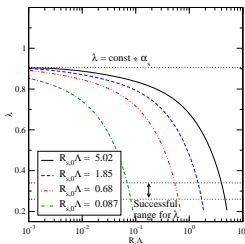
▶ large effect expected



Effects of NLO-corrections

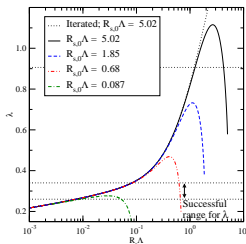
- NLO evolution: speed reduced

$$\lambda(Y) := \frac{d}{dY} \ln Q_s^2(Y)$$



LO JIMWLK

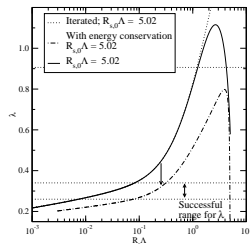
- too fast



+ running coupling

- remarkable slowdown
- fits become possible

▶ large effect expected



+ energy cons. corr.

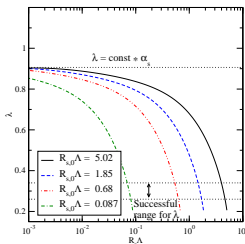
- asymptotic fits preferred



Effects of NLO-corrections

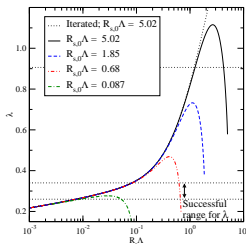
- NLO evolution: speed reduced

$$\lambda(Y) := \frac{d}{dY} \ln Q_s^2(Y)$$



LO JIMWLK

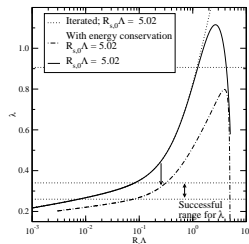
- too fast



+ running coupling

- remarkable slowdown
- fits become possible

▶ large effect expected



+ energy cons. corr.

- asymptotic fits preferred

- Effect of NLO impact factors?

yet unknown



Fit to HERA data

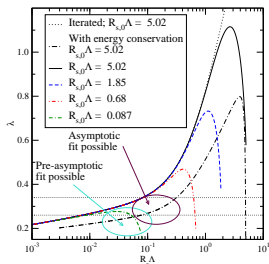
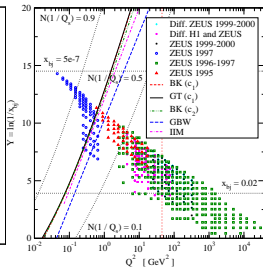
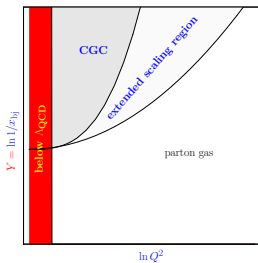
- Total cross section:

- Rapidity gap events (diffractive events):



Fit to HERA data

■ Total cross section:



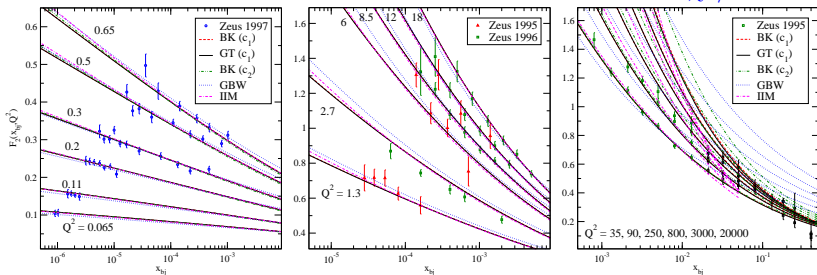
■ Rapidity gap events (diffractive events):



Fit to HERA data

■ Total cross section:

$\chi^2/\text{dof} \sim .8$

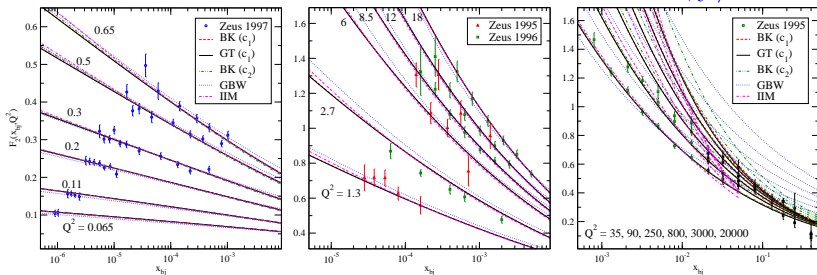


■ Rapidity gap events (diffractive events):



Fit to HERA data

■ Total cross section:

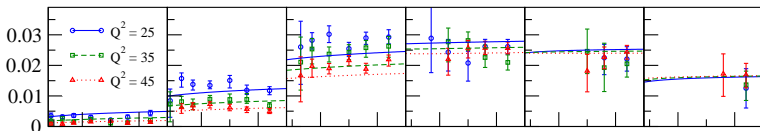
 $\chi^2/\text{dof} \sim .8$


■ Rapidity gaps events (diffractive events):

 $\chi^2/\text{dof} \sim 1.3$

ratios diffractive/total cross sections (sample only):

$0.28 \leq M_x \leq 2$ $2 \leq M_x \leq 4$ $4 \leq M_x \leq 8$ $8 \leq M_x \leq 15$ $15 \leq M_x \leq 25$ $25 \leq M_x \leq 35$



■ Lack of NLO impact factors: predictive power down!



Outline

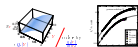
- 1 Motivation: gluons form the CGC
 - Background information on the standard model
 - Current and planned collider experiments
 - Enhanced gluon production at high energies
 - CGC: why the name
- 2 JIMWLK evolution: properties of the CGC
 - Gluons in observables
 - The evolution equation
 - The saturation scale
- 3 A sample experiment
 - Geometric scaling @ HERA
- 4 Getting quantitative
 - NLO corrections
 - HERA fits
- 5 Applications and outlook



Applications

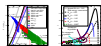
■ Geometric scaling in γ^*p & γ^*A

Golec-Biernat Wüsthoff, Kwiczinski; Kuokkanen, Rummukainen, Weigert; Albacete, Salgado, Wiedemann



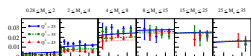
■ Precision fits to HERA σ_{total}

Kuokkanen, Rummukainen, Weigert



■ almost precision fits for HERA $\sigma_{\text{diffractive}}$

Kuokkanen, Rummukainen, Weigert



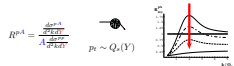
■ Sets the initial conditions of heavy ion collisions

Kovner, Weigert, McLerran; Venugopalan, Krasnitz; Lappi



■ qualitative only: BRAHMS (RHIC)

Suppression of the Cronin peak at forward rapidities (large Y) Kovchegov; Kovner, Wiedemann



■ Fit multiplicities at RHIC, predict LHC

Albacete, partial NLO



■ Monojets at RHIC

Albacete, Marquet, partial NLO



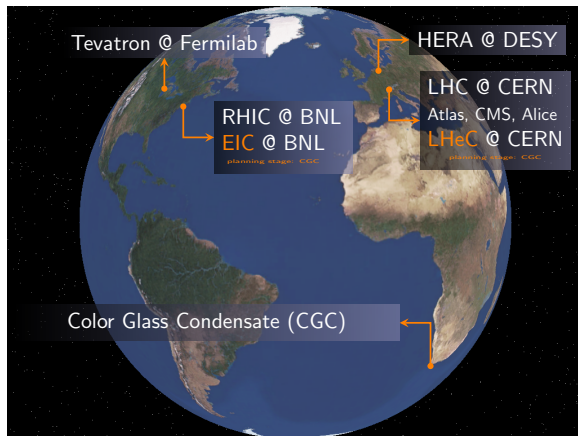
■ Forward (exclusive) particle production at RHIC (predict LHC?)

Albacete, Marquet, partial NLO





The Color Glass Condensate, a birds eye view



CGC in experiments @:

- RHIC, HERA
- Tevatron (new!)
- LHC
- EIC & LHeC (dedicated!)

Main characteristic:

- correlation length
 $R_s(Y) \sim \frac{1}{Q_s(Y)}$

Q_s -scaling: Y dependence

- via $Q_s(Y)$



Outline

- 6 The JIMWLK Hamiltonian
- 7 Running coupling
- 8 Experiments
 - From CGC to QGP
 - Cronin effect BRAHMS
 - Multiplicities
 - Monojets RHIC
 - Forward particle production RHIC



The JIMWLK Hamiltonian

[← back](#)

$$H_{\text{JIMWLK}} = -\frac{1}{2} \frac{\alpha_s}{\pi^2} \mathcal{K}_{\mathbf{xzy}} \left[i \nabla_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^a + i \bar{\nabla}_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^a + \tilde{U}_z^{ab} (i \bar{\nabla}_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^b + i \nabla_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^b) \right]$$

$$\mathcal{K}_{\mathbf{xzy}} = \frac{(\mathbf{x} - \mathbf{z}) \cdot (\mathbf{z} - \mathbf{y})}{(\mathbf{x} - \mathbf{z})^2 (\mathbf{z} - \mathbf{y})^2}$$

[integration convention for x, z, y]

$i \nabla_{\mathbf{x}}^a$ and $i \bar{\nabla}_{\mathbf{x}}^a$ are functional derivatives:

$$i \nabla_{\mathbf{x}}^a := -[U_{\mathbf{x}} t^a]_{ji} \frac{\delta}{\delta U_{\mathbf{x},ij}} \qquad i \bar{\nabla}_{\mathbf{x}}^a := [t^a U_{\mathbf{x}}]_{ji} \frac{\delta}{\delta U_{\mathbf{x},ij}}$$



The JIMWLK Hamiltonian

[← back](#)

$$H_{\text{JIMWLK}} = -\frac{1}{2} \frac{\alpha_s}{\pi^2} \mathcal{K}_{\mathbf{xzy}} \left[i\nabla_{\mathbf{x}}^a i\nabla_{\mathbf{y}}^a + i\bar{\nabla}_{\mathbf{x}}^a i\bar{\nabla}_{\mathbf{y}}^a + \tilde{U}_z^{ab} (i\bar{\nabla}_{\mathbf{x}}^a i\nabla_{\mathbf{y}}^b + i\nabla_{\mathbf{x}}^a i\bar{\nabla}_{\mathbf{y}}^b) \right]$$

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$$i\nabla_{\mathbf{x}}^a := -[U_{\mathbf{x}} t^a]_{ji} \frac{\delta}{\delta U_{\mathbf{x},ij}} \quad i\bar{\nabla}_{\mathbf{x}}^a := [t^a U_{\mathbf{x}}]_{ji} \frac{\delta}{\delta U_{\mathbf{x},ij}}$$

generate l. & r. inv vector fields, r & l rotations:

$$e^{-i\omega^a (i\nabla^a)} U = U e^{i\omega^a t^a} \quad e^{-i\omega^a (i\bar{\nabla}^a)} U = e^{-i\omega^a t^a} U$$

reps of the algebras:

$$[i\nabla^a, i\nabla^b] = i f^{abc} i\nabla^c \quad [i\bar{\nabla}^a, i\bar{\nabla}^b] = i f^{abc} i\bar{\nabla}^c \quad [i\bar{\nabla}^a, i\nabla^b] = 0$$



The JIMWLK Hamiltonian

[◀ back](#)

$$H_{\text{JIMWLK}} = -\frac{1}{2} \frac{\alpha_s}{\pi^2} \mathcal{K}_{\mathbf{xzy}} \left[i \nabla_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^a + i \bar{\nabla}_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^a + \tilde{U}_z^{ab} (i \bar{\nabla}_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^b + i \nabla_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^b) \right]$$

$$\mathcal{K}_{\mathbf{xzy}} = \frac{(\mathbf{x} - \mathbf{z}) \cdot (\mathbf{z} - \mathbf{y})}{(\mathbf{x} - \mathbf{z})^2 (\mathbf{z} - \mathbf{y})^2}$$

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$$i \bar{\nabla}_{\mathbf{x}}^a := [t^a U_{\mathbf{x}}]_{ji} \frac{\delta}{\delta U_{\mathbf{x}, ij}}$$

physics content:

- $\tilde{U}_z^{ab} (i \bar{\nabla}_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^b + i \nabla_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^b)$ real emission
- $i \nabla_{\mathbf{x}}^a i \nabla_{\mathbf{y}}^a + i \bar{\nabla}_{\mathbf{x}}^a i \bar{\nabla}_{\mathbf{y}}^a$ virt. correction
- real emission term nonlinear evolution



Outline

6 The JIMWLK Hamiltonian

7 Running coupling

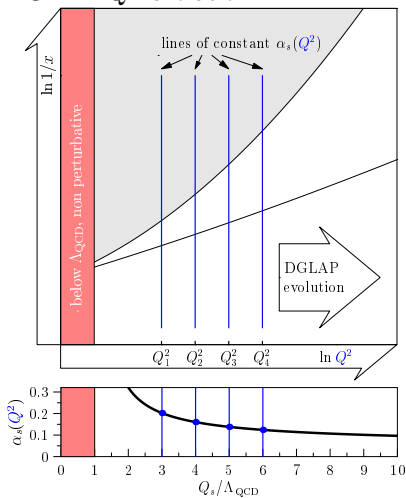
8 Experiments

- From CGC to QGP
- Cronin effect BRAHMS
- Multiplicities
- Monojets RHIC
- Forward particle production RHIC



Running coupling is essential: Q^2 vs small x

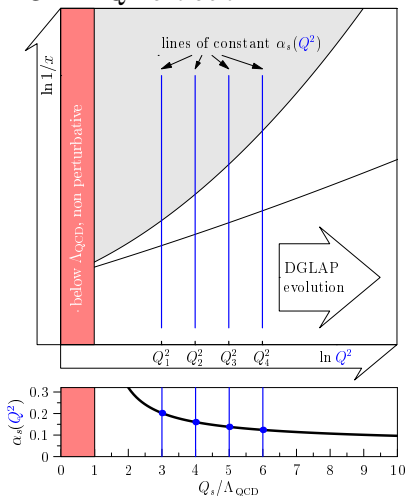
■ DGLAP: Q^2 evolution





Running coupling is essential: Q^2 vs small x

■ DGLAP: Q^2 evolution

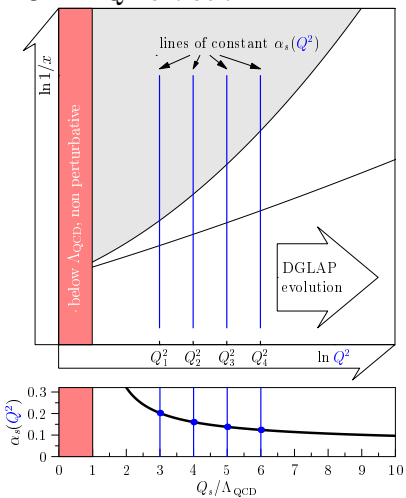


■ no qualitative changes



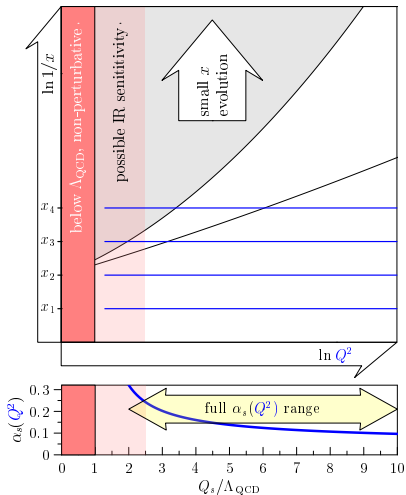
Running coupling is essential: Q^2 vs small x

DGLAP: Q^2 evolution



■ no qualitative changes

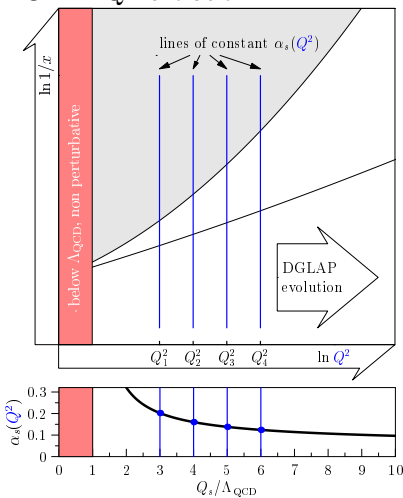
JIMWLK, BK: small x evolution





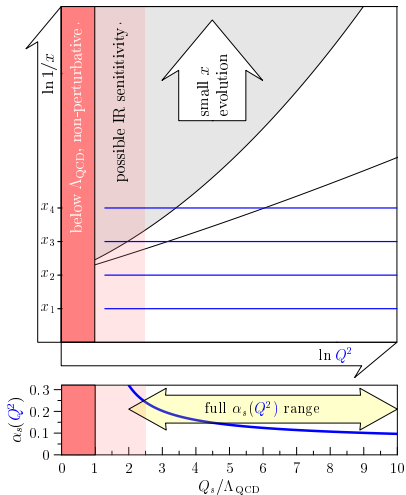
Running coupling is essential: Q^2 vs small x

DGLAP: Q^2 evolution



■ no qualitative changes

JIMWLK, BK: small x evolution



■ qualitative changes



Outline

6 The JIMWLK Hamiltonian

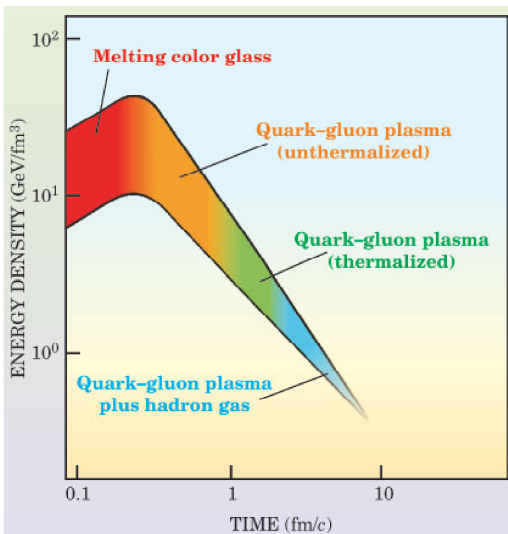
7 Running coupling

8 Experiments

- From CGC to QGP
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From Colored Glass to Quark Gluon Plasma



McLerran, Ludlam
Physics Today
Oct 2003

← core of neutron star

← nuclear matter



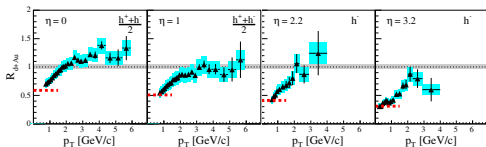
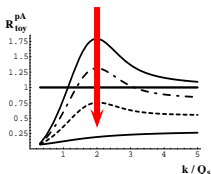
Erasing the Cronin effect on the parton level

[BRAHMS]

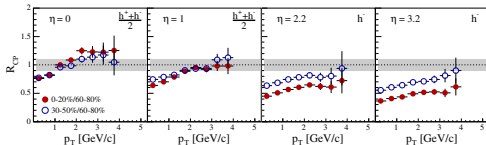
$$R^{pA} = \frac{d\sigma^{pA}}{d^2kdY} \bigg/ A \frac{d\sigma^{pp}}{d^2kdY}$$



$$p_t \sim Q_s(Y)$$



disappears at forward rapidities



disappears faster centrally

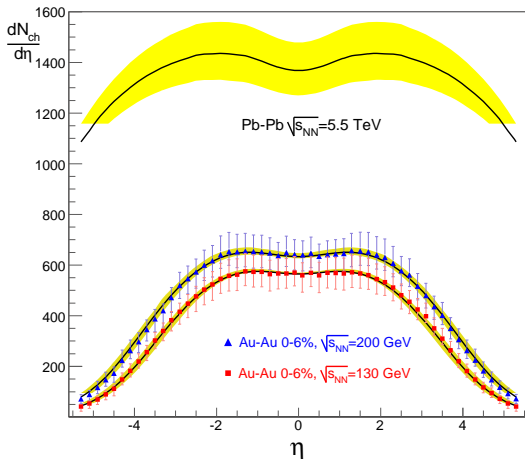
- qualitative only: no hadronization of partons, simulation only LO





Multiplicities at RHIC and LHC(?)

- Fit multiplicities at RHIC, predict LHC

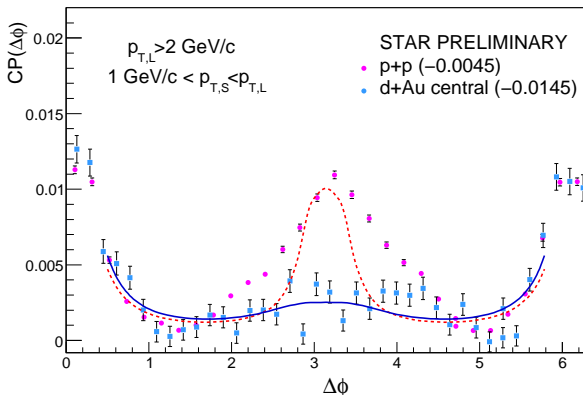


J. L. Albacete, Phys. Rev. Lett. **99** (2007) 262301 [arXiv:0707.2545 [hep-ph]]



Monojets at RHIC

- light nuclei: back to back jets
not quantitative: energy too low centrally, see Cronin
- heavy nuclei: Monojets; back to back correlation is broken

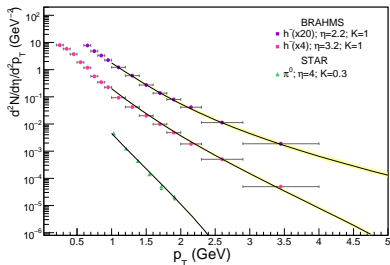


J. L. Albacete and C. Marquet, arXiv:1005.4065 [hep-ph].

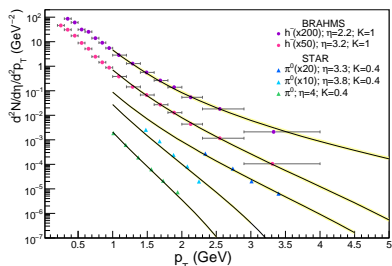
- partial NLO: running coupling only!



■ d Au



■ p p



J. L. Albacete and C. Marquet, Phys. Lett. B **687** (2010) 174 [arXiv:1001.1378 [hep-ph]]

■ partial NLO: running coupling only!