

The hottest matter on earth: what its made of,
how its formed, and how it flows

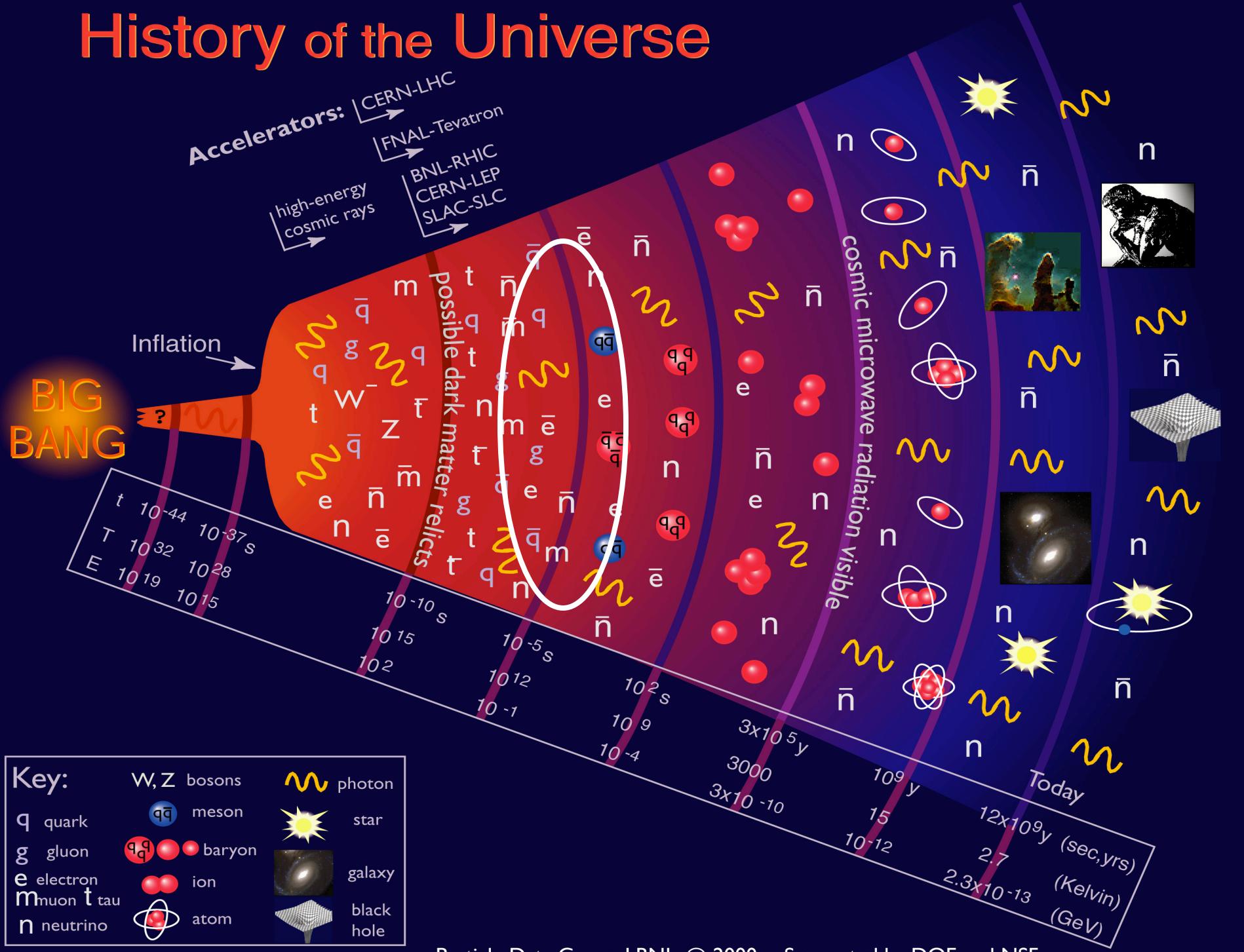
Raju Venugopalan
Brookhaven National Laboratory

UCT lectures, February 2012

Outline of lectures

- ◆ **Lecture I: QCD and the Quark-Gluon Plasma**
- ◆ **Lecture II: Gluon Saturation and the Color Glass Condensate**
- ◆ **Lecture III: Quantum field theory in strong fields. Factorization and the Glasma**
- ◆ **Lecture IV: Quantum field theory in strong fields.
Instabilities and the spectrum of initial quantum fluctuations**
- ◆ **Lecture V: Quantum field theory in strong fields. Decoherence, hydrodynamics, Bose-Einstein Condensation and thermalization**
- ◆ **Lecture VI: Future prospects: RHIC, LHC and the EIC**

History of the Universe



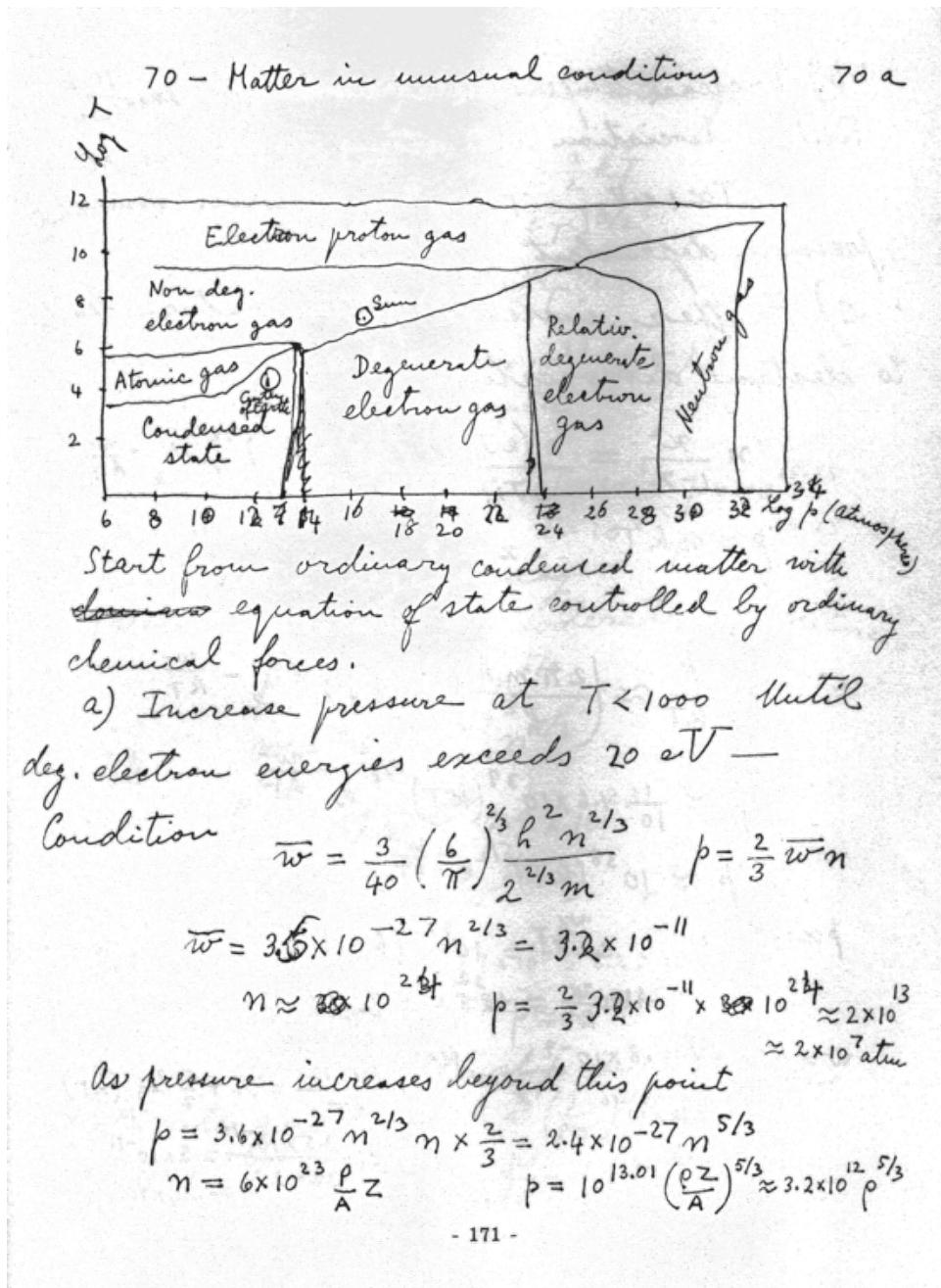


**What happens to matter when it
is heated up to temperatures
100,000 times hotter than the
center of the sun?**

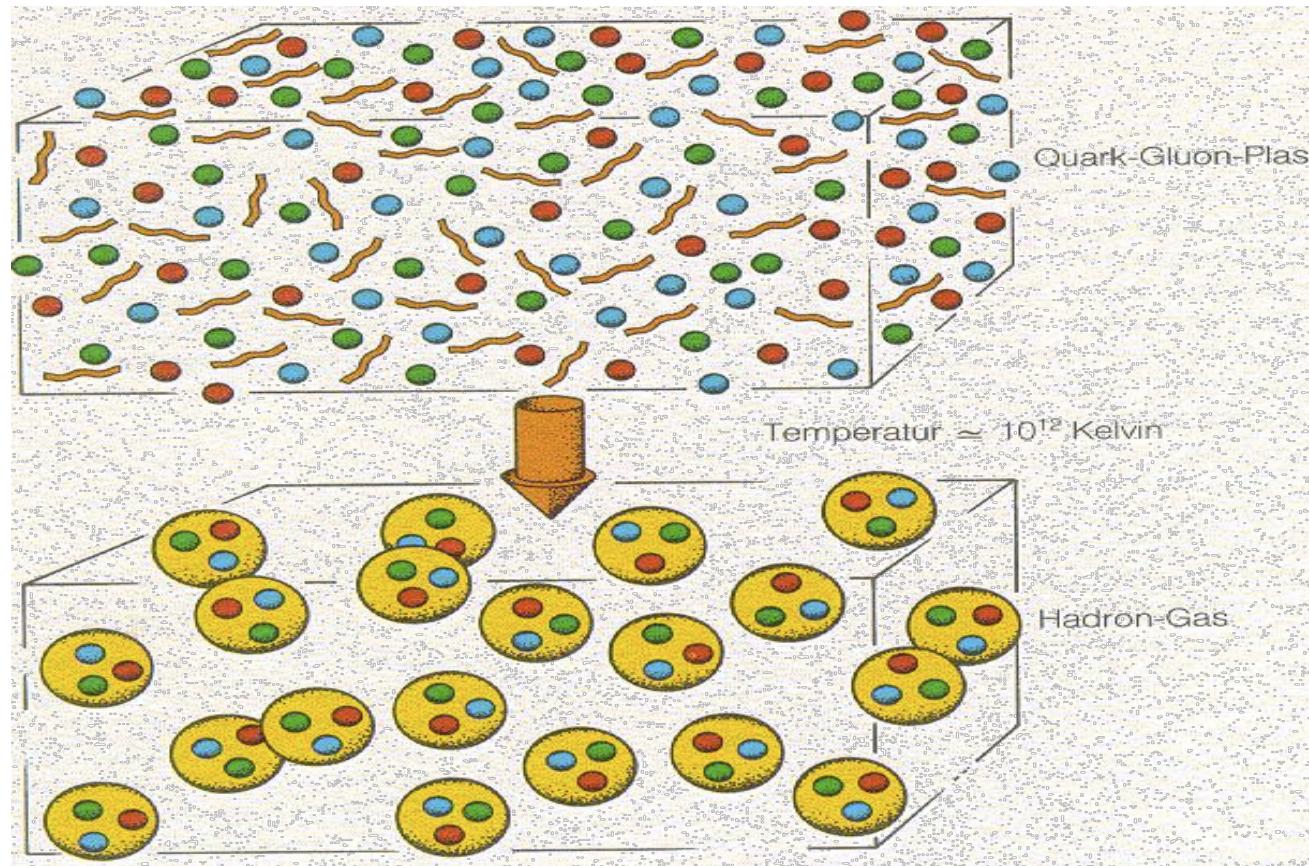
center of sun 15,000,000 Celsius (27,000,000 Fahrenheit)

Matter in unusual conditions

E. Fermi: " Notes on Thermodynamics and Statistics " (1953)

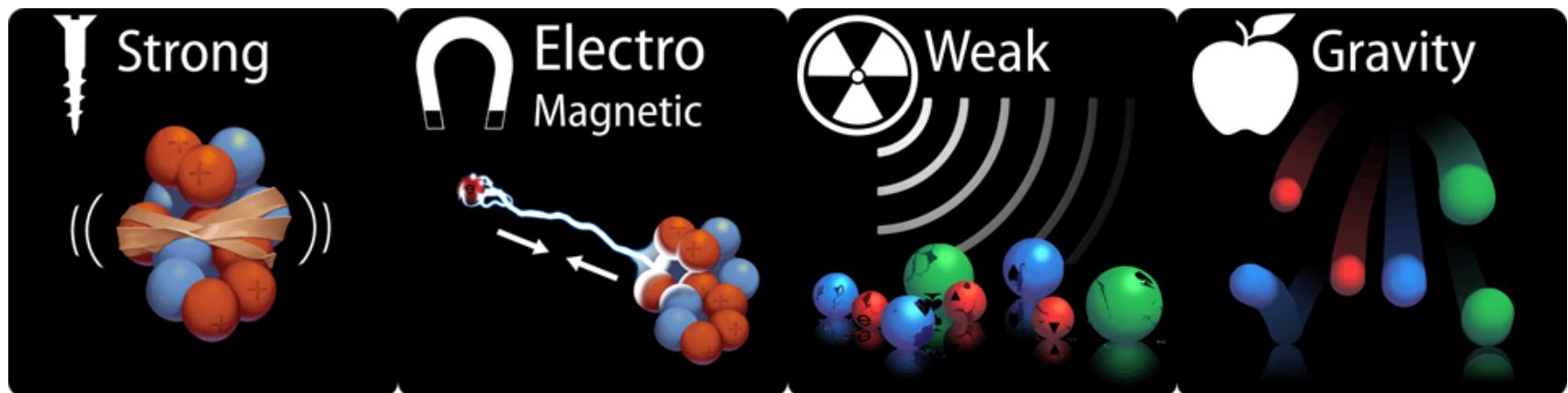


Universe evolved from Quarks & Gluons to Hadrons



How did this happen ?
Is there a phase transition (vapor to liquid) transition ?

QCD-Theory of the Strong Force



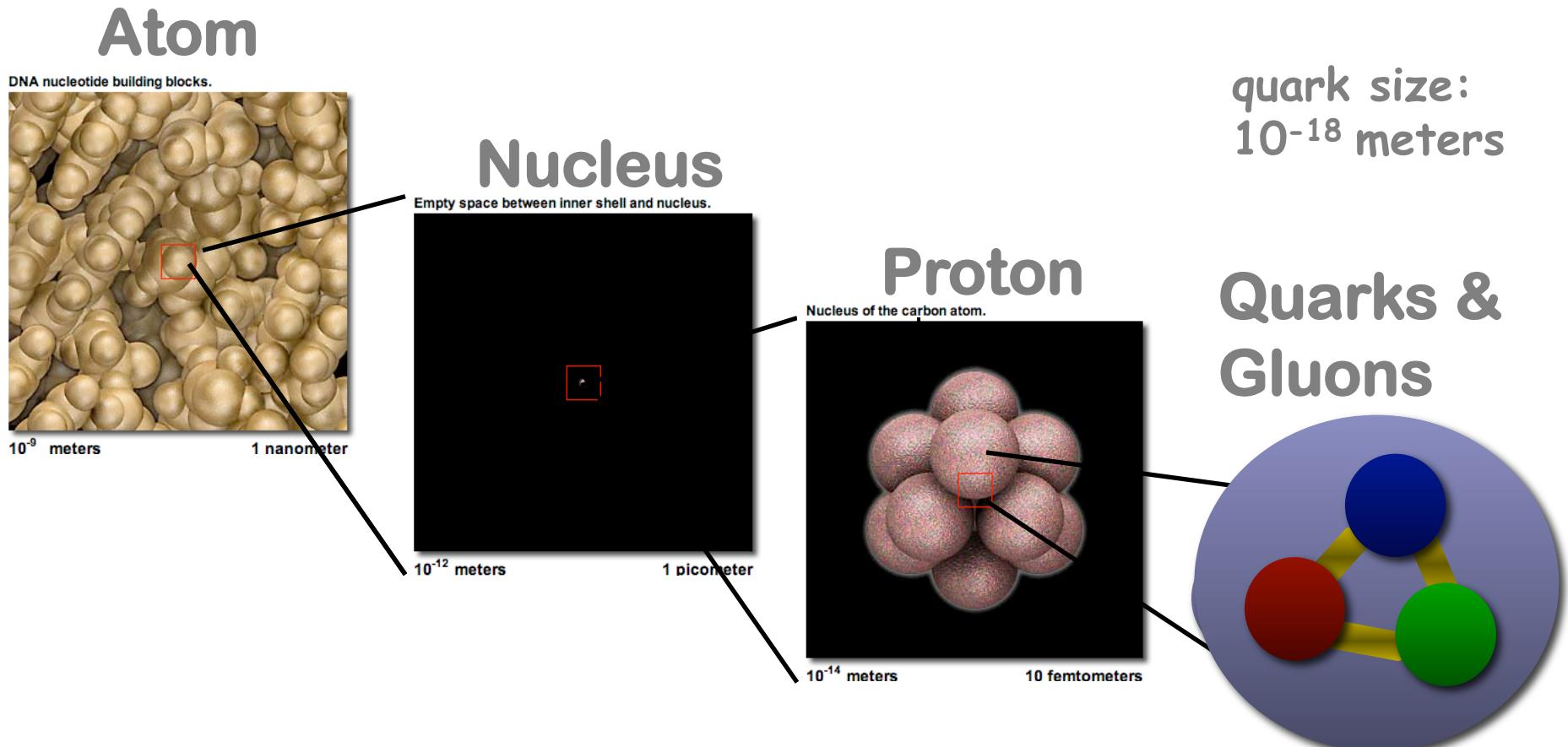
Relative strength	1	10^{-2}	10^{-6}	10^{-38}
Interaction range	10^{-15}	∞	10^{-15}	∞



2004 Nobel Prize in Physics

What does the strong force act on?

Quarks & Gluons: elementary, fundamental constituents of 99 % of visible matter in the universe.



Quantum Chromodynamics (QCD): fundamental theory of the strong force

QCD: theory of the strong interaction

$$\mathcal{L}_{\text{QCD}} = -\frac{1}{4}F_{\mu\nu}^a F^{\mu\nu,a} + \sum_f \bar{q}_f^a (i\gamma^\mu D_\mu - m_f) q_f^a$$

$F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g f^{abc} A_\mu^b A_\nu^c$

$D_\mu = \partial_\mu - ig A_\mu^a \frac{\lambda^a}{2}$

Quark fields

Gluon fields

Symmetries of QCD Lagrangian:

$$SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\text{i}} \times \underbrace{U(1)_A \times U(1)_B}_{\text{iii}}$$

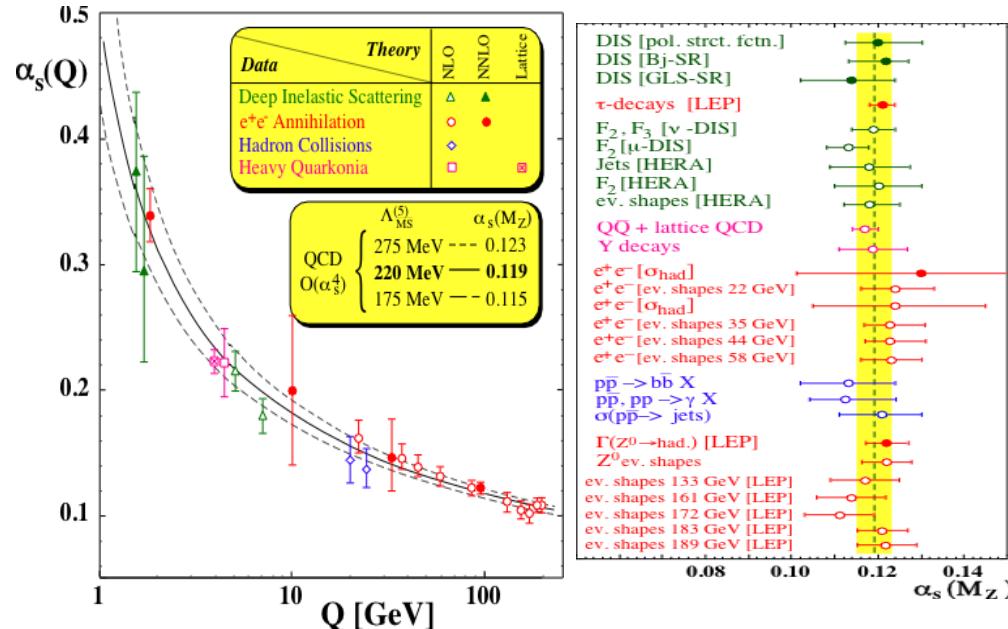
- i) Gauge “color” symmetry: unbroken but confined
- ii) Global “chiral” symmetry: exact for massless quarks
- iii) Baryon number and axial charge ($m=0$) are conserved
- iv) Scale invariance: $\psi(x) \rightarrow \lambda^{3/2} \psi(\lambda x)$; $A_\mu^a(x) \rightarrow \lambda A_\mu^a(\lambda x)$
- v) Discrete C,P & T symmetries

Chiral, Axial, Scale and (in principle) P & T are broken by Vacuum / Quantum effects

QCD: asymptotic freedom



Gross,Wilczek, Politzer



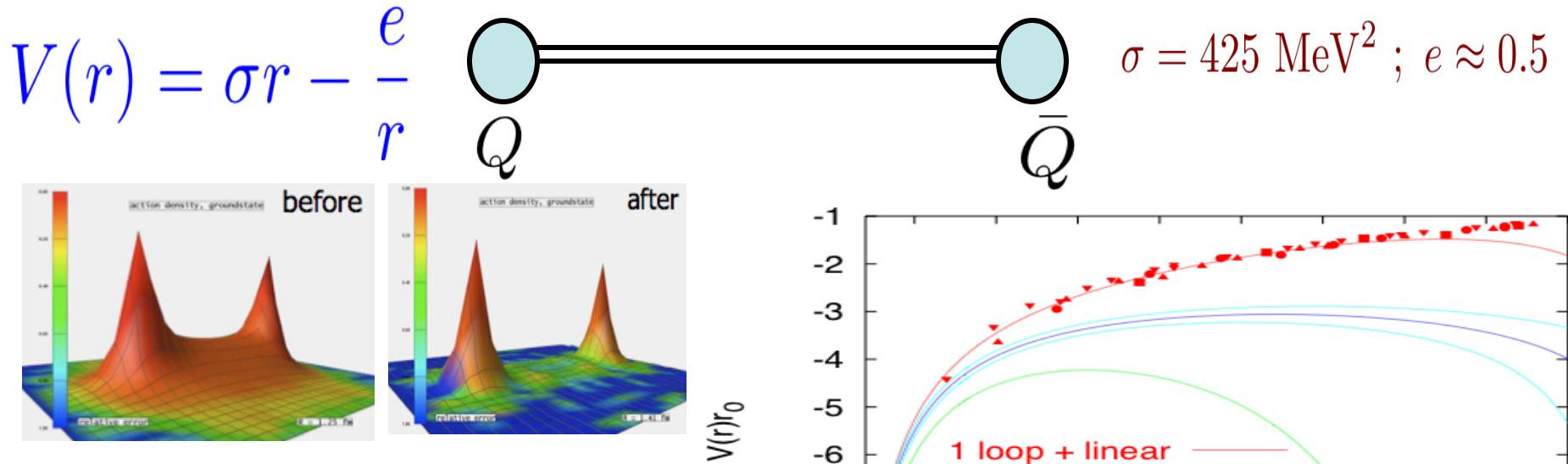
Coupling grows weaker
with increasing momentum
transfer (shorter distances)

$$\alpha_S(Q^2) = \frac{12\pi}{(11N_c - 2n_f) \ln\left(\frac{Q^2}{\Lambda_{\text{QCD}}^2}\right)} \left[1 - \frac{6(153 - 19n_f)}{(11N_c - 2n_f)^2} \frac{\ln(\ln(Q^2/\Lambda_{\text{QCD}}^2))}{\ln(Q^2/\Lambda_{\text{QCD}}^2)} \right] + \dots$$

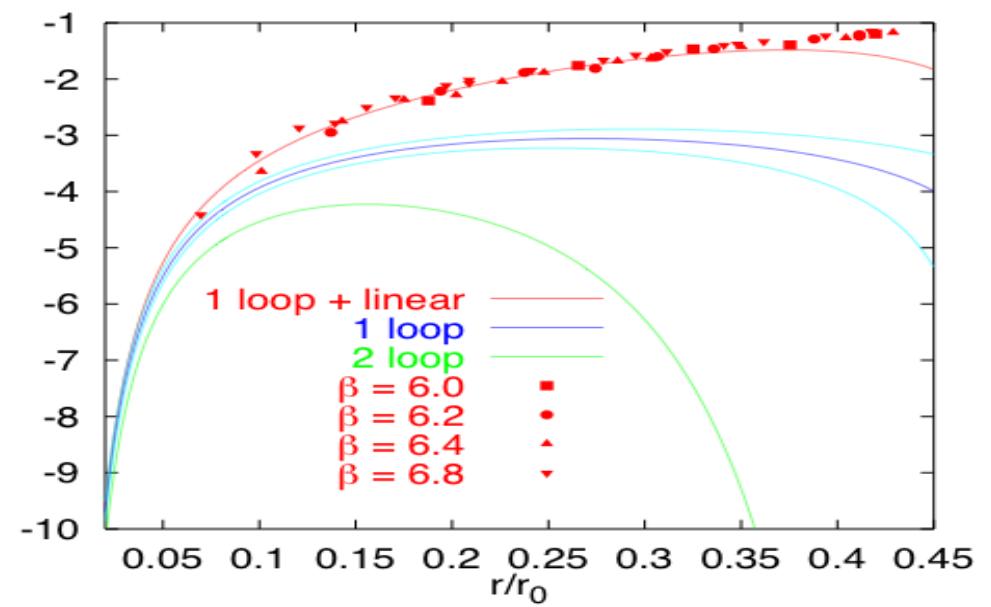
Super-dense and Super-hot QCD matter
is a weakly coupled gas of quarks and gluons
-- analytical computations feasible

Cabibo-Parisi; Collins-Perry

QCD: Infrared Slavery



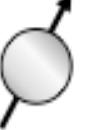
Potential between static
quark-anti-quark pair
grows linearly at large distances -
intuitive picture of confinement



QCD matter is strongly interacting at low Temp. and Density
- static properties computed using lattice methods



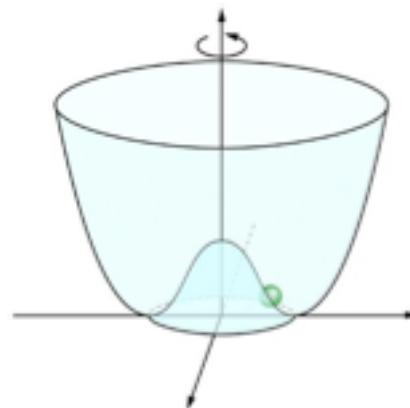
QCD: Chiral symmetry breaking

Chiral basis :  $q_L = \frac{1}{2}(1 - \gamma_5)q$, $q_R = \frac{1}{2}(1 + \gamma_5)q$ 

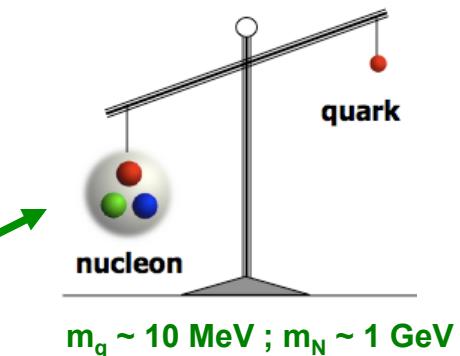
QCD Lagrangian : $\mathcal{L}_{\text{cl}} = \mathcal{L}_{\text{cl}}(q_L, A) + \mathcal{L}_{\text{cl}}(q_R, A)$

Quantum QCD vacuum:

Chiral condensate: spontaneously generate mass via
Nambu-Goldstone mechanism



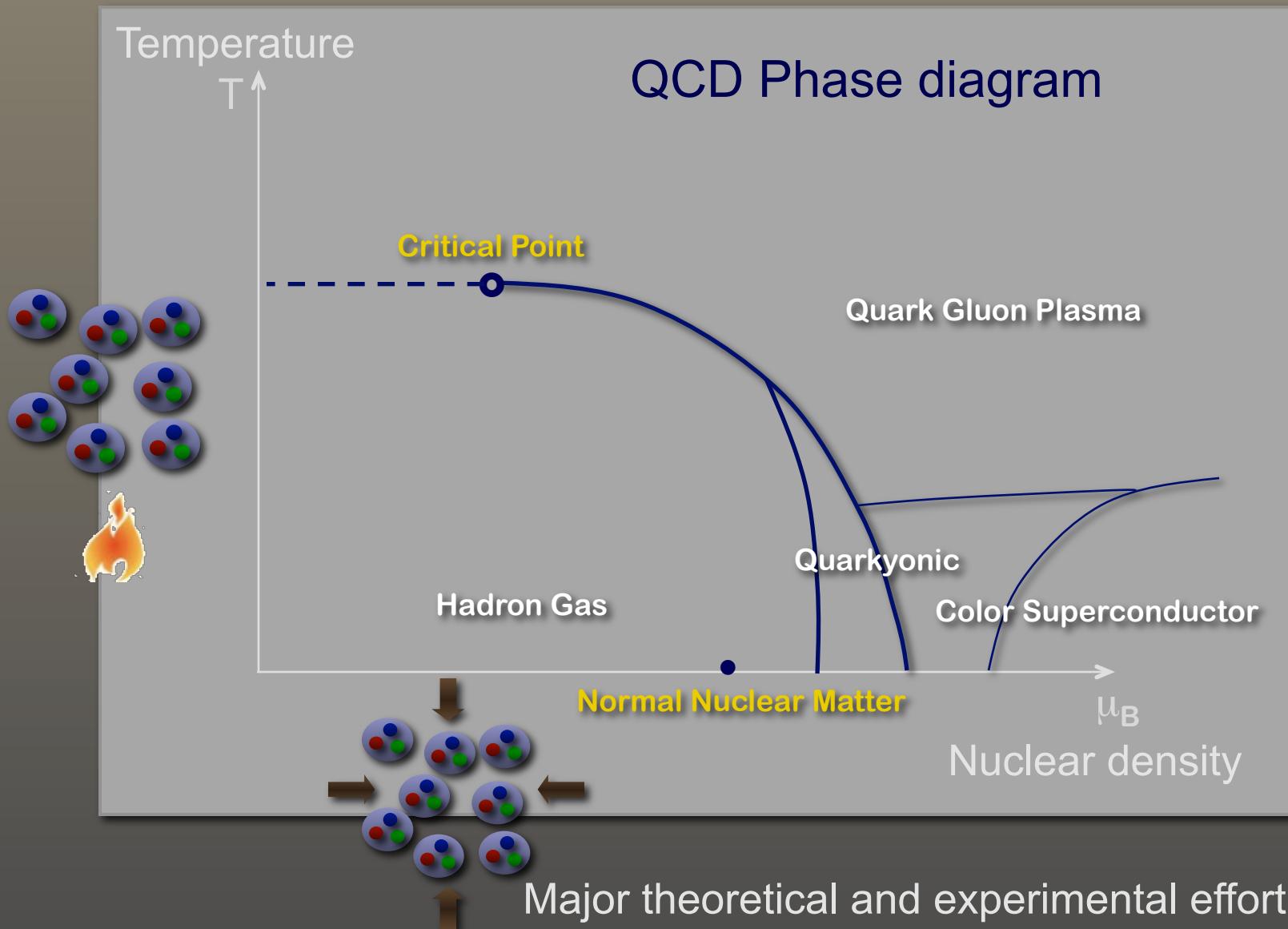
$$\rightarrow \langle \bar{q}_R q_L \rangle \neq 0$$



Quantum violation of U(1)_A $\partial_\mu J_A^\mu = -2n_f \frac{\alpha_S}{8\pi} F^{\mu\nu,a} \tilde{F}_{\mu\nu}^a$

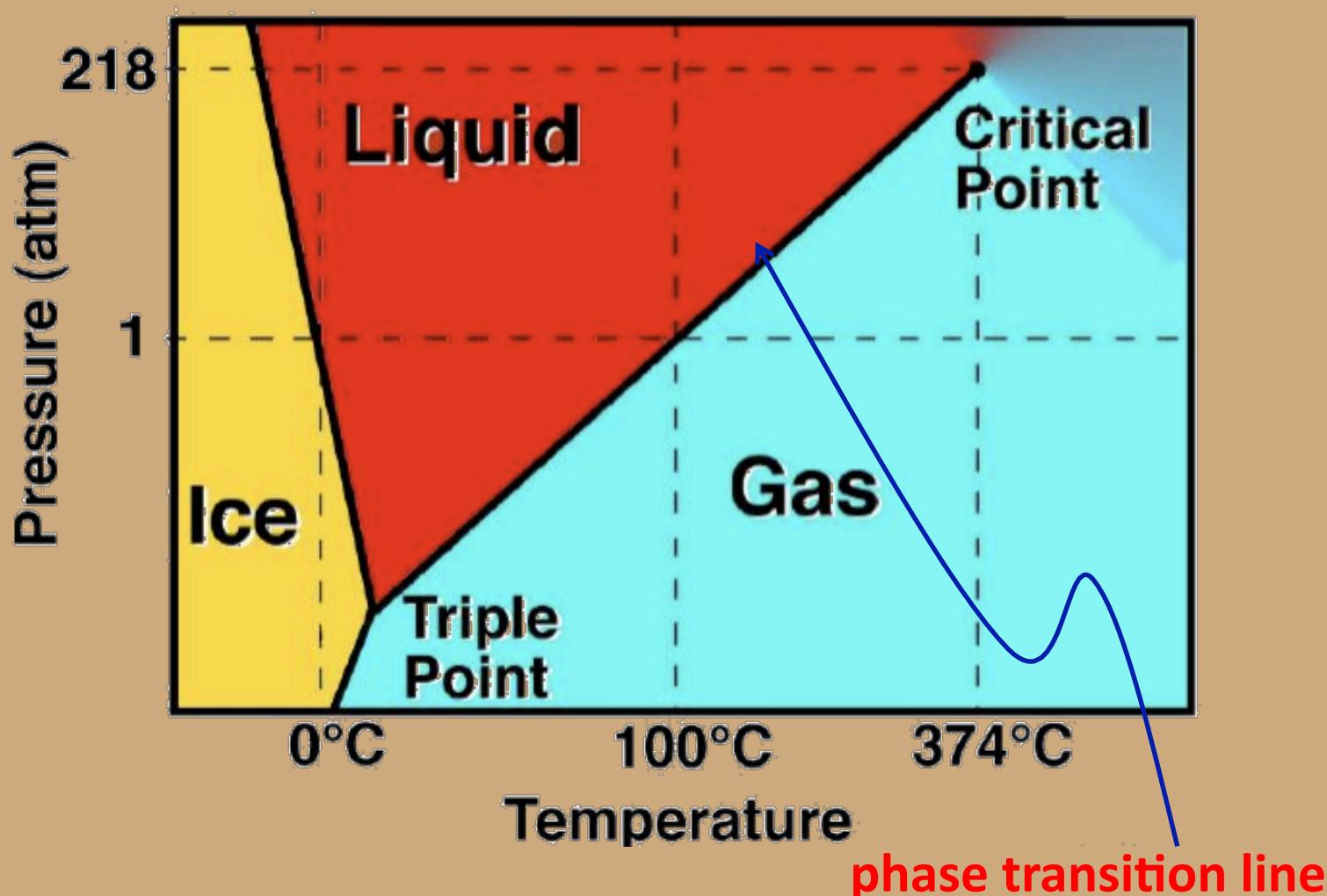
Chiral symmetry restoration at large T: $\langle \bar{q}q \rangle \longrightarrow 0$

phase diagram of nuclear matter



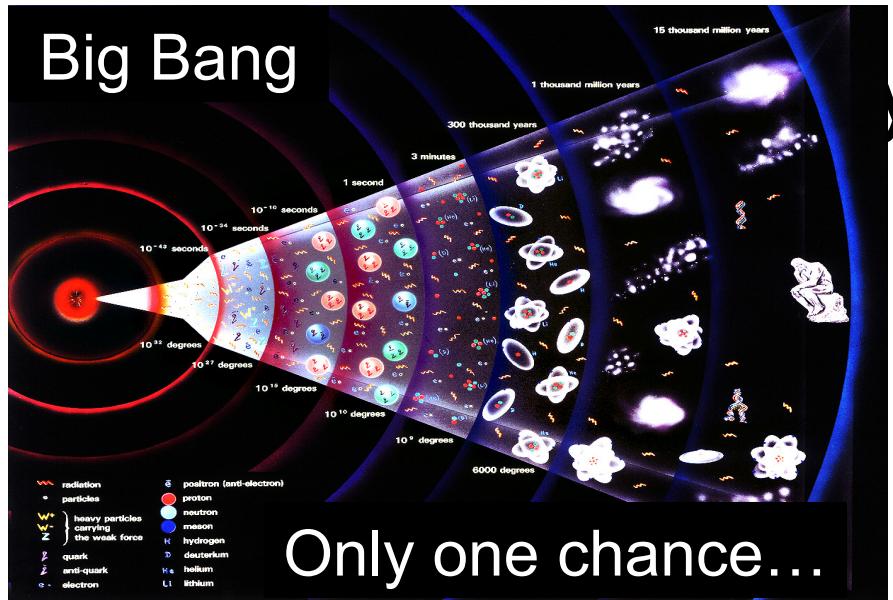
Phase Diagram of Water

Physicists' way of mapping the phase that will exist for a given pressure and temperature



Where to study QCD matter ?

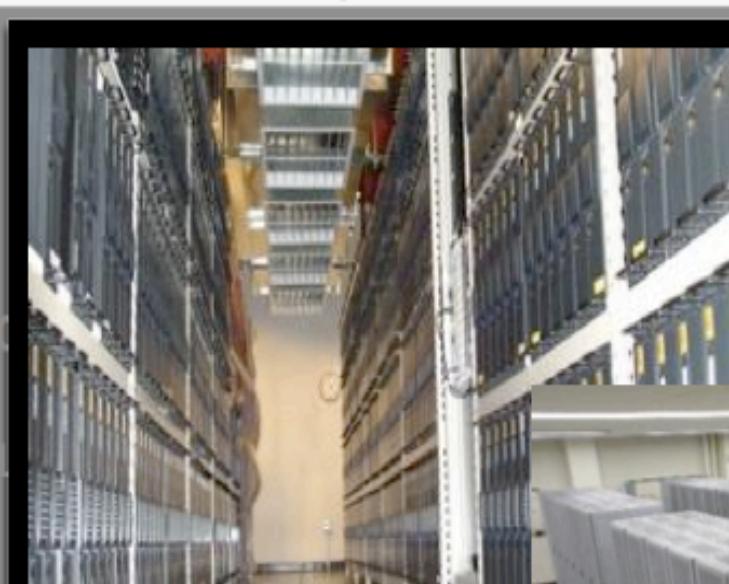
Peter Steinberg



Lattice QCD



1
1
1
(Flops)



QCD Cluster @ FNAL



BlueGene/L @ KEK



ApeNEXT @ Rome



PACS-CS @ Tsukuba

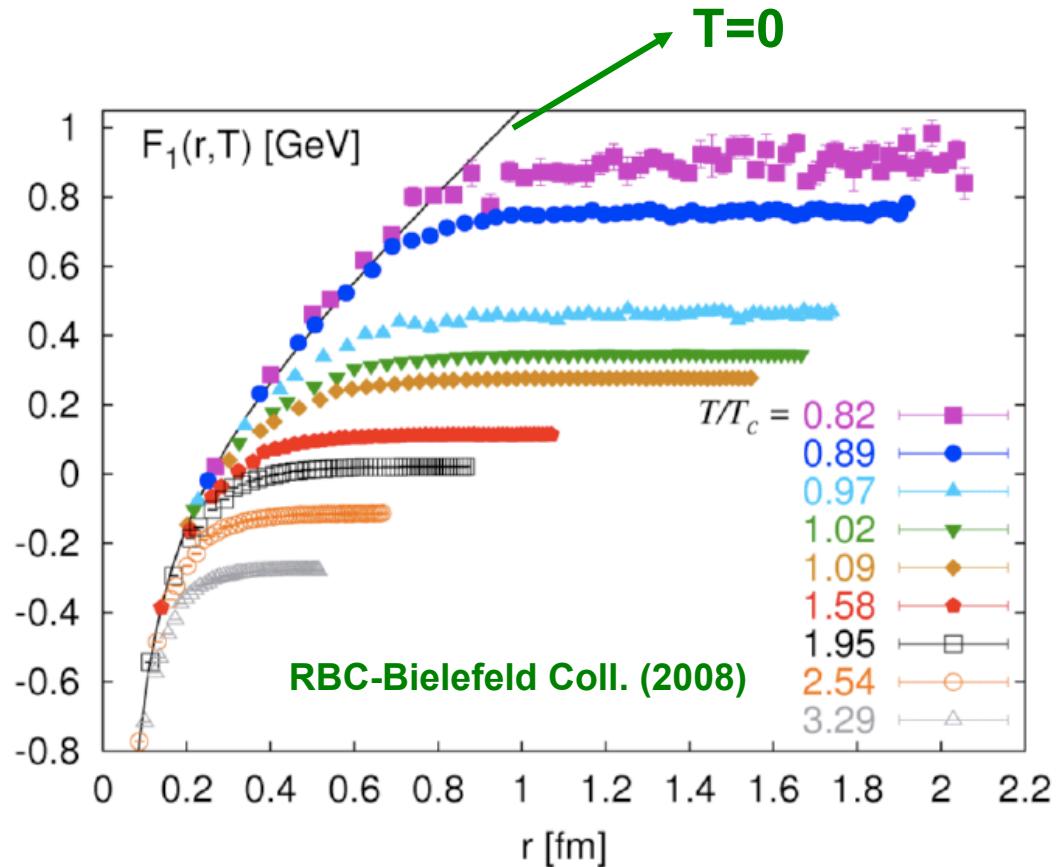


QCDOC @ RBRC-Columbia

(year)

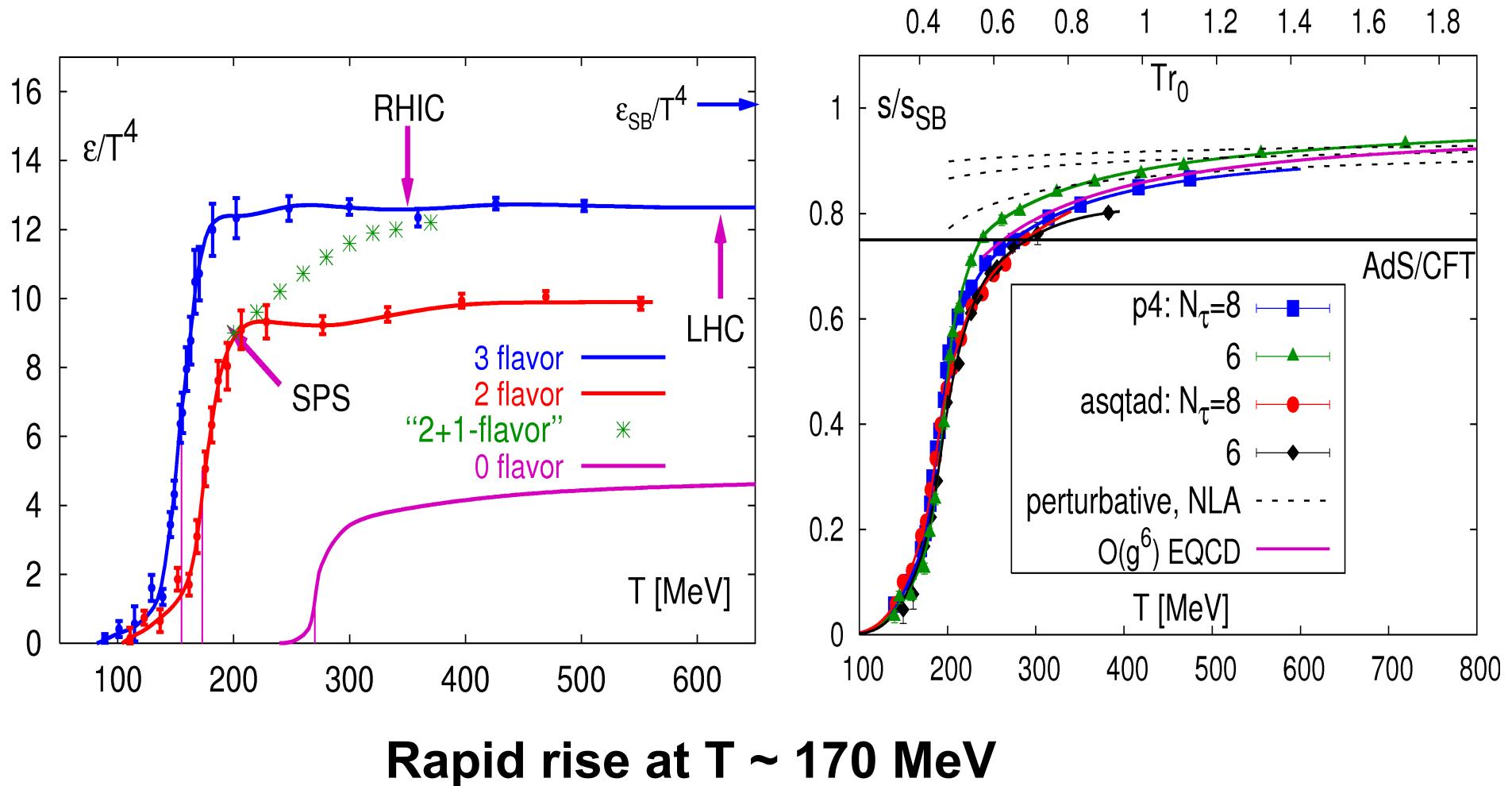
Hatsuda, PANIC 08

Lattice QCD at finite T: Plasma screens potential between heavy quarks



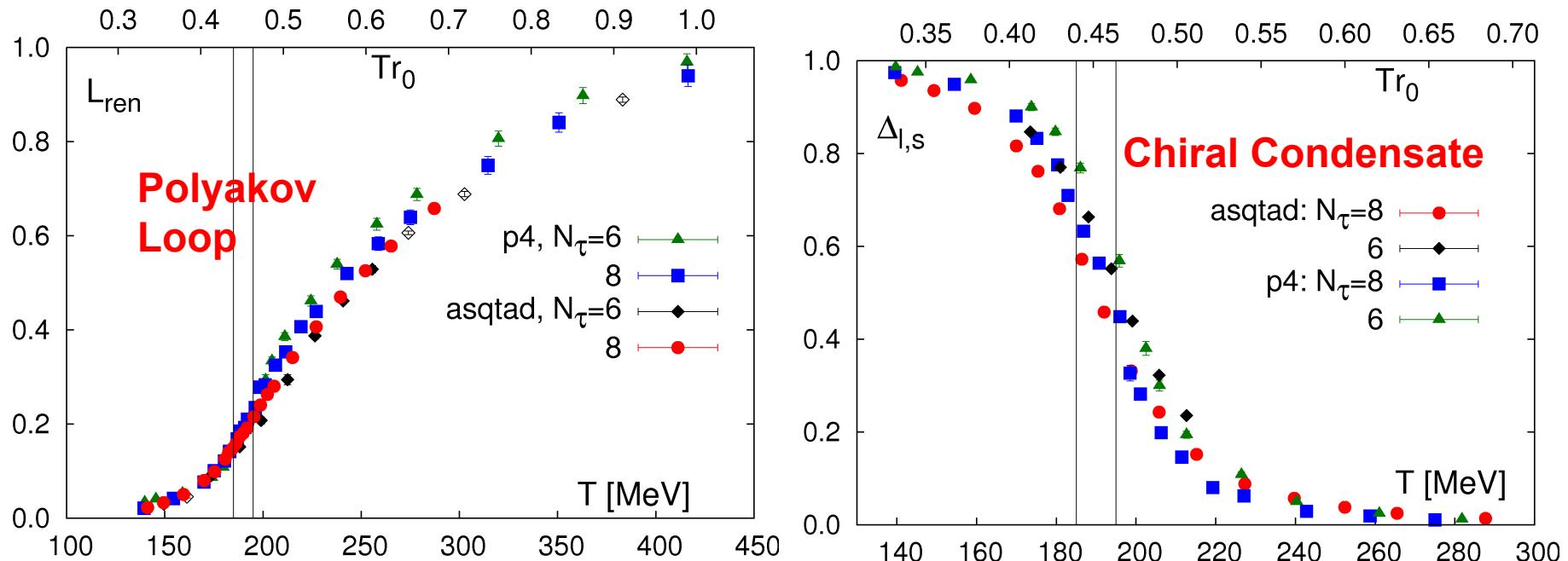
Suppression of heavy quarks suggested as possible signature of
Quark-Gluon formation - Matsui, Satz 1986

QCD at finite T: Lattice results for Energy & Entropy Density



Rapid rise at $T \sim 170$ MeV

QCD at finite T: Lattice results for Confinement and Chiral Symmetry



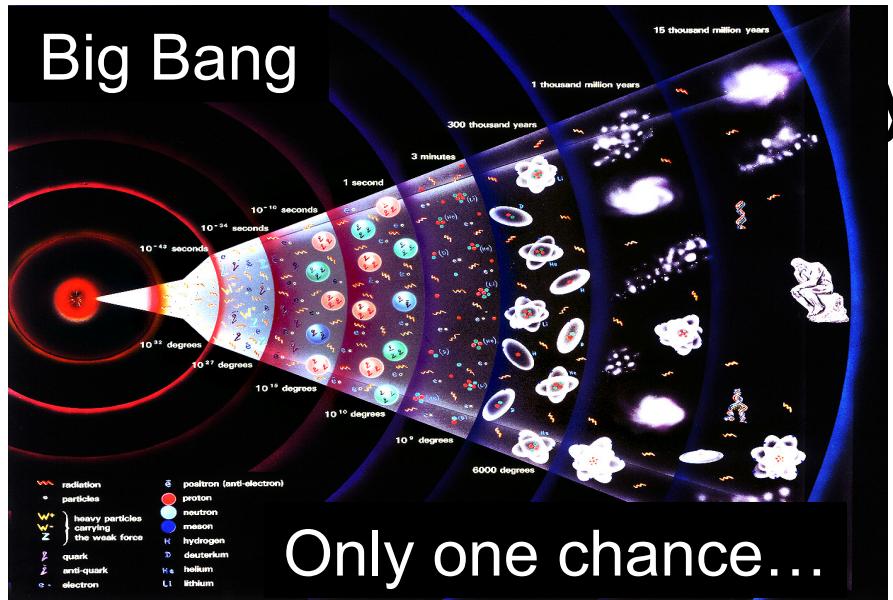
**Rapid change in same T interval
corresponding to rapid change in
Energy and Entropy densities**

Bazavov et al, (HotQCD Coll.) arXiv:0903.4379

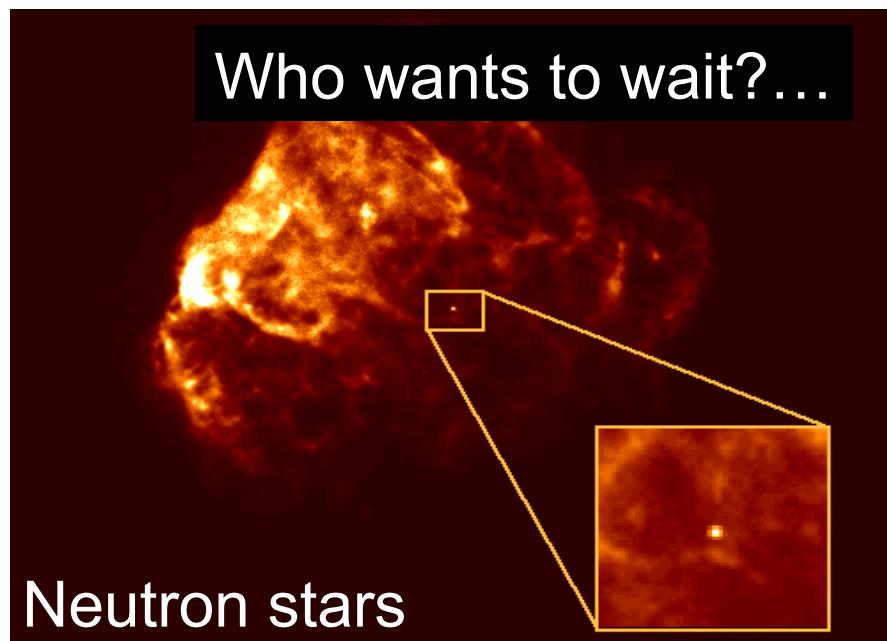
**Finite T transition is a
crossover** Aoki et al, Nature 443 (06) 675

Where to study QCD matter ?

Peter Steinberg



Lattice QCD



Heavy Ion Experiments

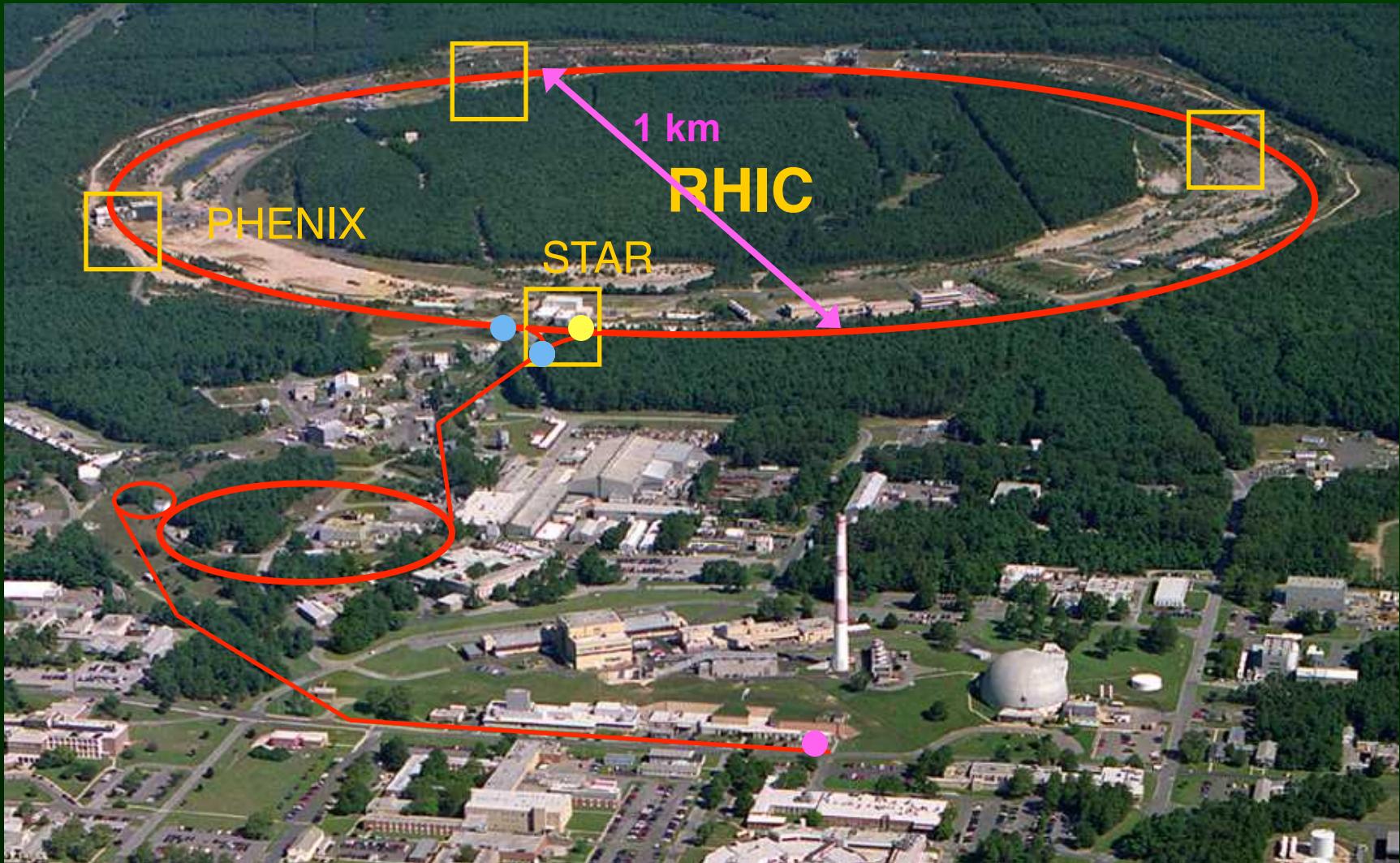
Facility	Location	System	Energy (CMS)
AGS	BNL, New York	Au+Au	2.6-4.3 GeV
SPS	CERN, Geneva	Pb+Pb	8.6-17.2 GeV
RHIC	BNL, New York	Au+Au	200 GeV
LHC	CERN, Geneva	Pb+Pb	5.5 TeV

The place to look for quark-gluon plasma



the Collider at
Brookhaven Lab
visible from space

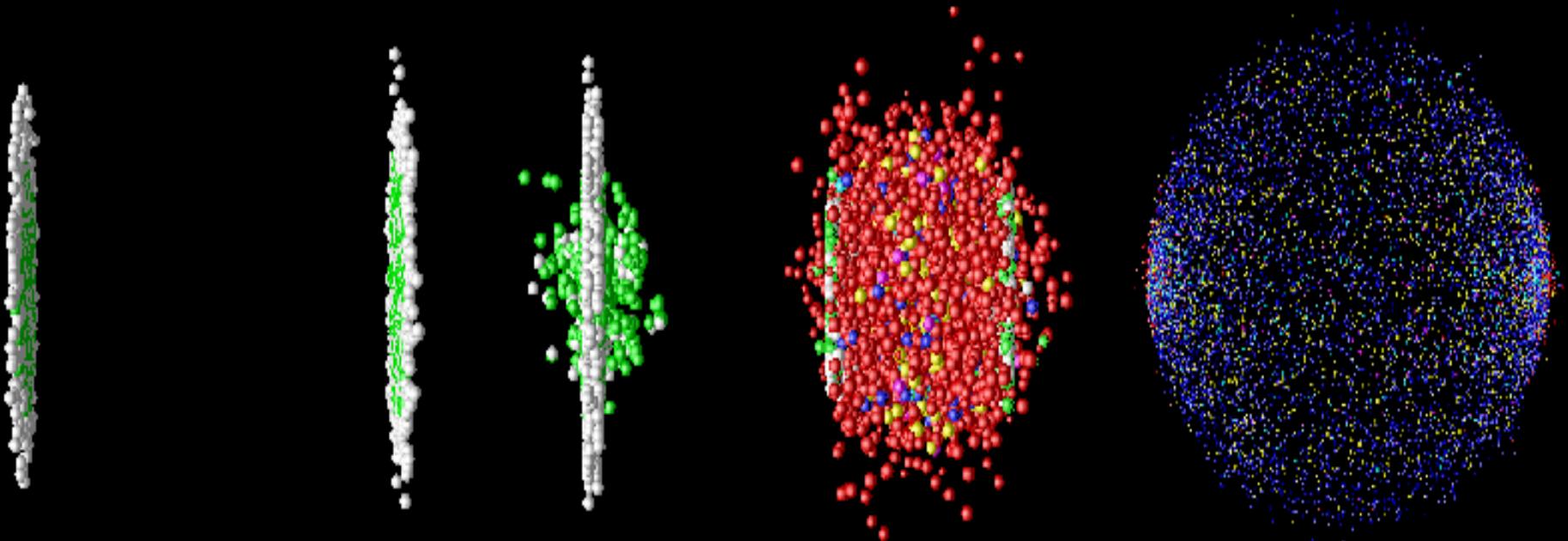
Relativistic Heavy Ion Collider RHIC @ BNL



3 km ring

animation by Mike Lisa

major international enterprise: thousands of scientists and engineers



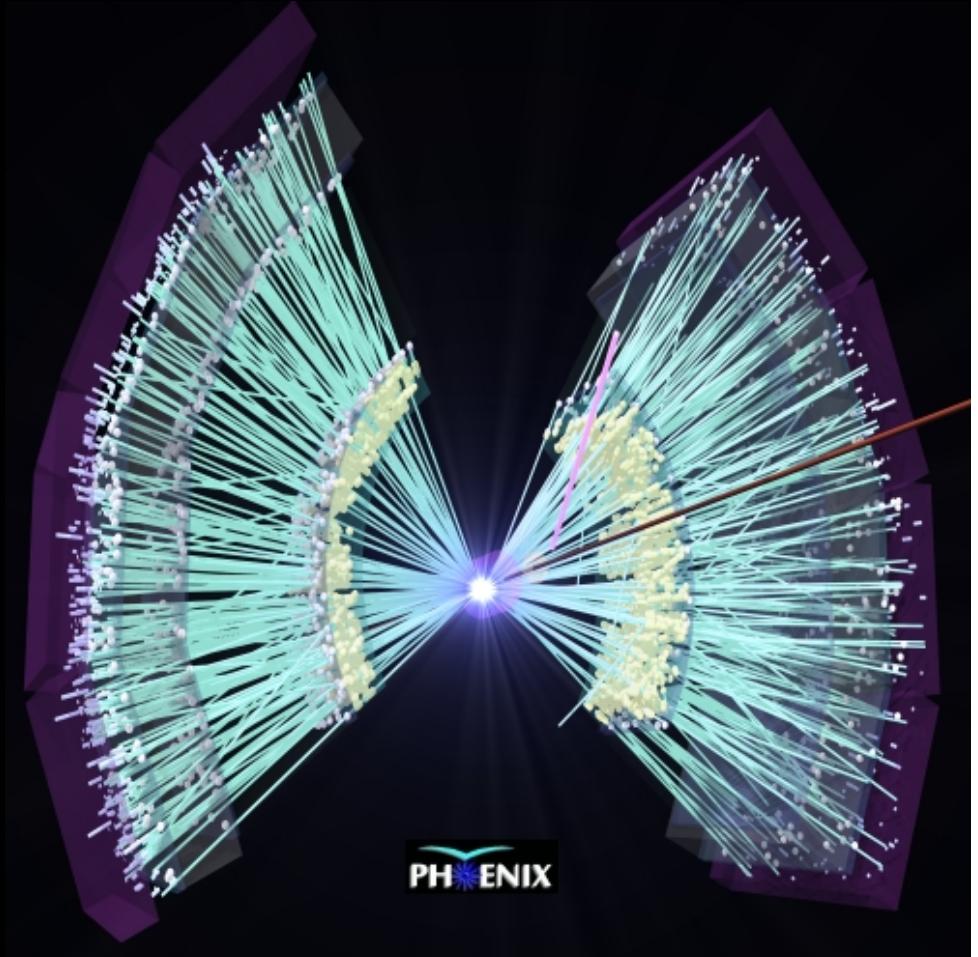
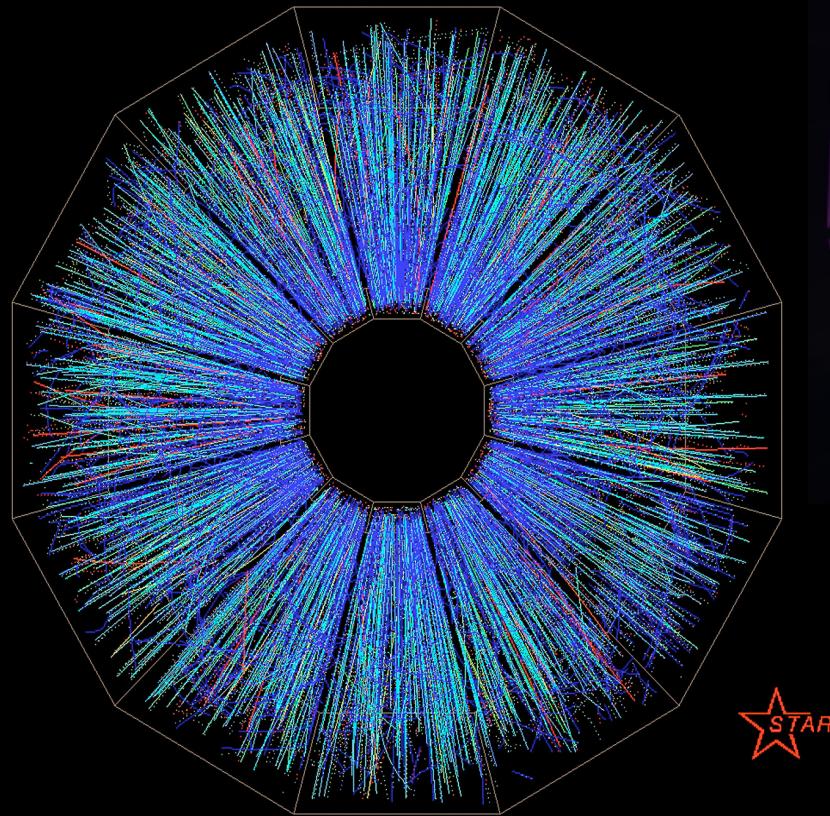
Heavy "gold" nucleus at
0.9997 the speed
of light

Two such nuclei collide

Quark Gluon Plasma
formed

Decays into 1000s of
sub-atomic particles

We cannot determine if a QGP was produced by observing it directly - its lifetime is too short ($\sim 10^{-23}$ seconds...)



but studying the detailed structure behind these images, RHIC scientists conclude the QGP was a nearly perfect fluid

BBC NEWS | Science/Nature | Early Universe was 'liquid-like'

<http://news.bbc.co.uk/2/hi/science/nature/4462209.stm>

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BBC NEWS

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Early Universe was 'liquid-like'

Physicists say they have created a new state of hot, dense matter by crashing together the nuclei of gold atoms.

The high-energy collisions prised open the nuclei to reveal their most basic particles, known as quarks and gluons.

The researchers, at the US Brookhaven National Laboratory, say these particles were seen to behave as an almost perfect "liquid".

The work is expected to help scientists explain the conditions that existed just milliseconds after the Big Bang.

The details, presented to the American Physical Society in Florida, will be published across a number of papers in the journal Nuclear Physics A.

They summarise the work of four collaborative experiments - dubbed Brahms, Phenix, Phobos and Star - which have been running on Brookhaven's

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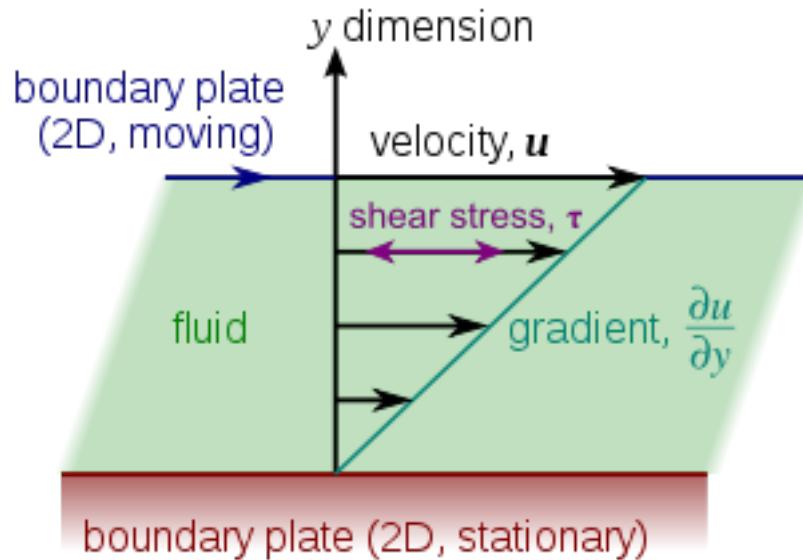
TOP SCIENCE/NATURE STORIES

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Viscosity

- Introduced by Claude Navier in 1822 (Navier Stokes equation)



Pitch tar: viscosity 230 billion times
that of water

$$\frac{F}{A} = \eta \frac{\partial u_x}{\partial y}$$

A (nearly) perfect fluid

“Bjorken Hydrodynamics” for relativistic fluids

Viscous term smaller than ideal term for

$$\frac{d\varepsilon}{d\tau} = - \frac{(\varepsilon + P - \frac{4}{3} \frac{\eta}{\tau})}{\tau}$$
$$\frac{\eta}{\varepsilon + P} \frac{1}{\tau} \equiv \boxed{\frac{\eta}{s} \frac{1}{\tau T}} \ll 1$$

From kinetic theory

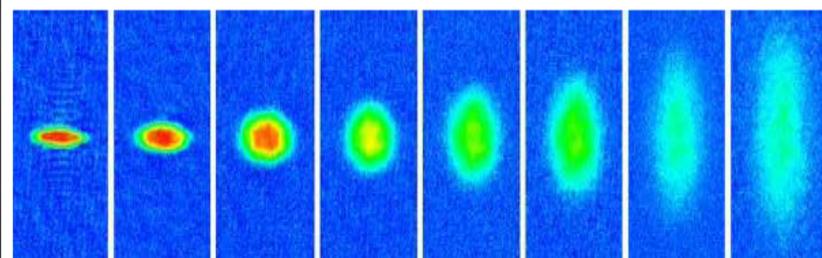
$$\boxed{\frac{\eta}{s} \sim \frac{\hbar}{k_B} \frac{\tau_{\text{relax.}}}{\tau_{\text{quant.}}}}$$



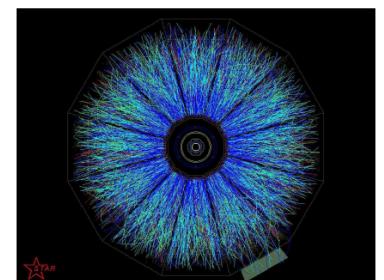
H_2O



${}^4\text{He}$



${}^6\text{Li}$



sQGP

A (nearly) perfect fluid

“Bjorken Hydrodynamics”

$$\frac{d\varepsilon}{d\tau} = - \frac{\left(\varepsilon + P - \frac{4}{3}\frac{\eta}{\tau}\right)}{\tau}$$

$$\frac{\eta}{\varepsilon + P} \frac{1}{\tau} \equiv \boxed{\frac{\eta}{s}} \frac{1}{\tau T} \ll 1$$

Viscous term smaller than ideal term for

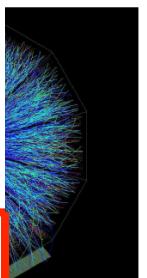
AdS/CFT:

$$\frac{\eta}{s} \geq \frac{\hbar}{k_B} \frac{1}{4\pi}$$



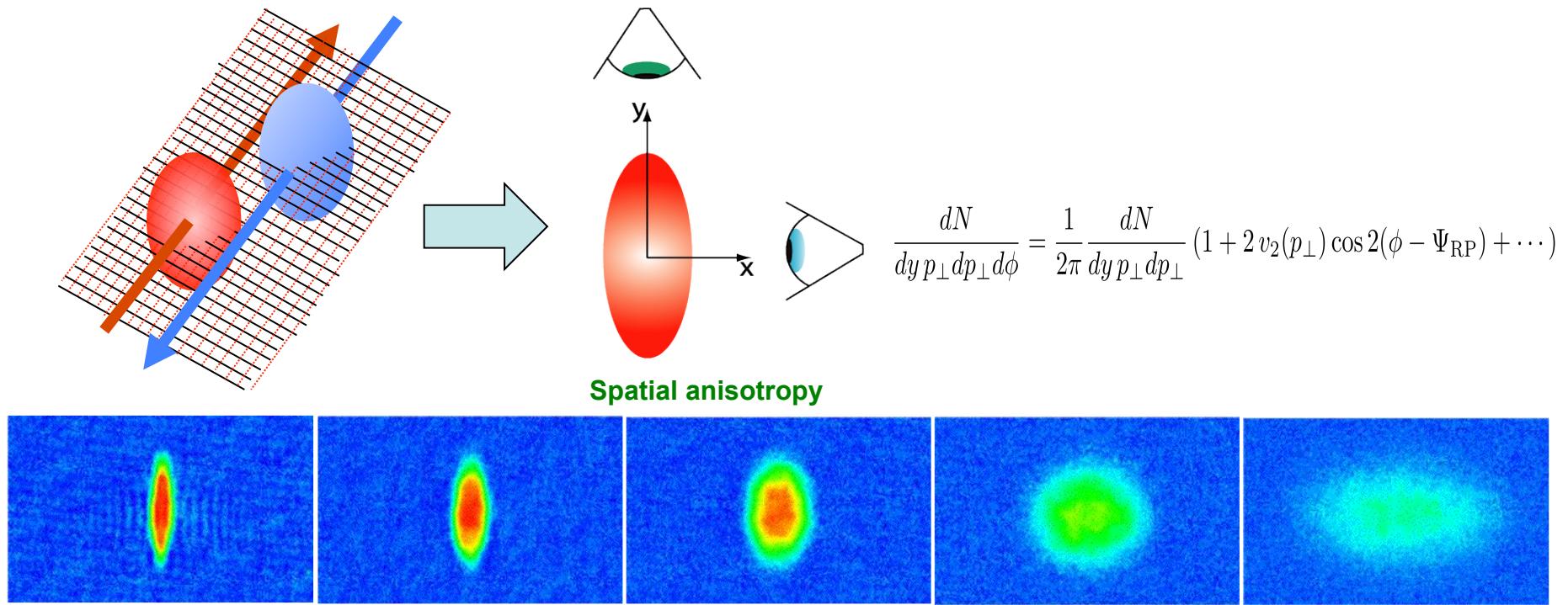
H_2O

Fluid	$T [K]$	$\eta [Pa \cdot s]$	$\eta/n [\hbar]$	$\eta/s [\hbar/k_B]$
H_2O	370	2.9×10^{-4}	85	8.2
4He	2	1.2×10^{-6}	0.5	1.9
6Li ($ a_s \simeq \infty$)	23×10^{-6}	$\leq 1.7 \times 10^{-15}$	≤ 1	≤ 0.5
QGP	2×10^{12}	$\leq 5 \times 10^{11}$	-	≤ 0.4



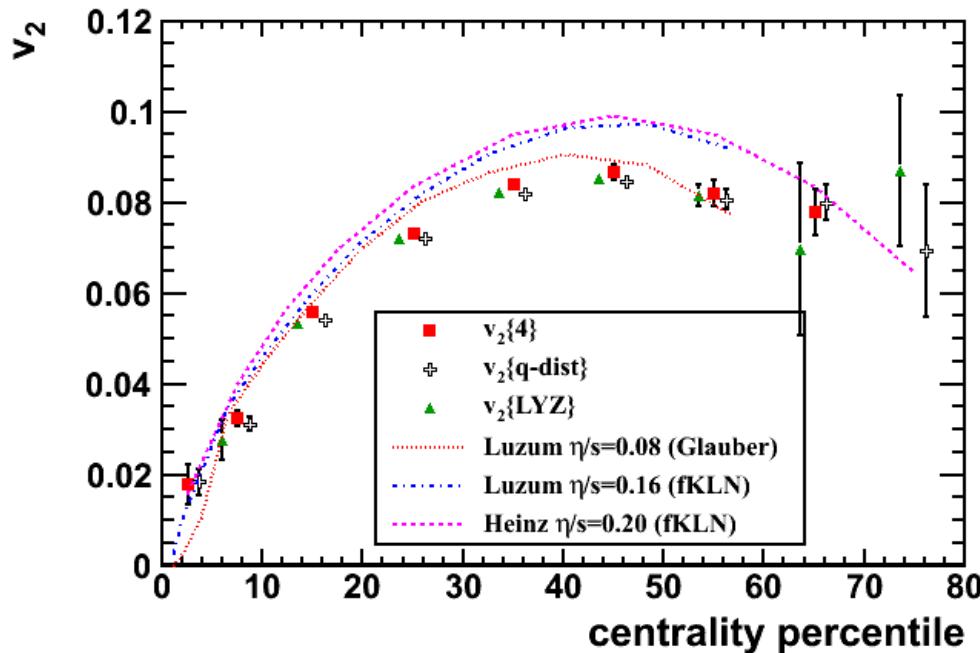
GP

Strong flow = (nearly) ideal hydrodynamics

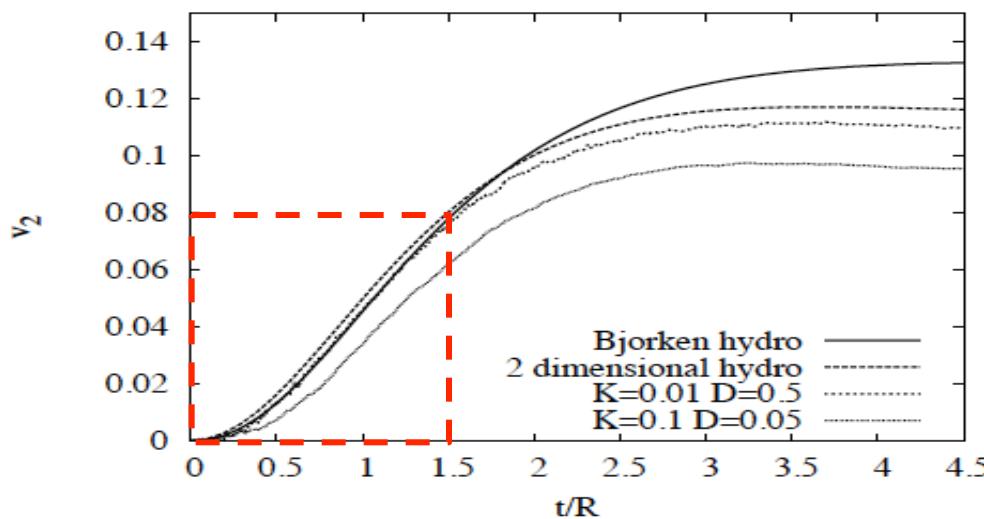


v_2 measures how efficiently hot matter converts spatial anisotropies to momentum anisotropy
– most efficient way is hydrodynamics

Strong flow = (nearly) ideal hydrodynamics

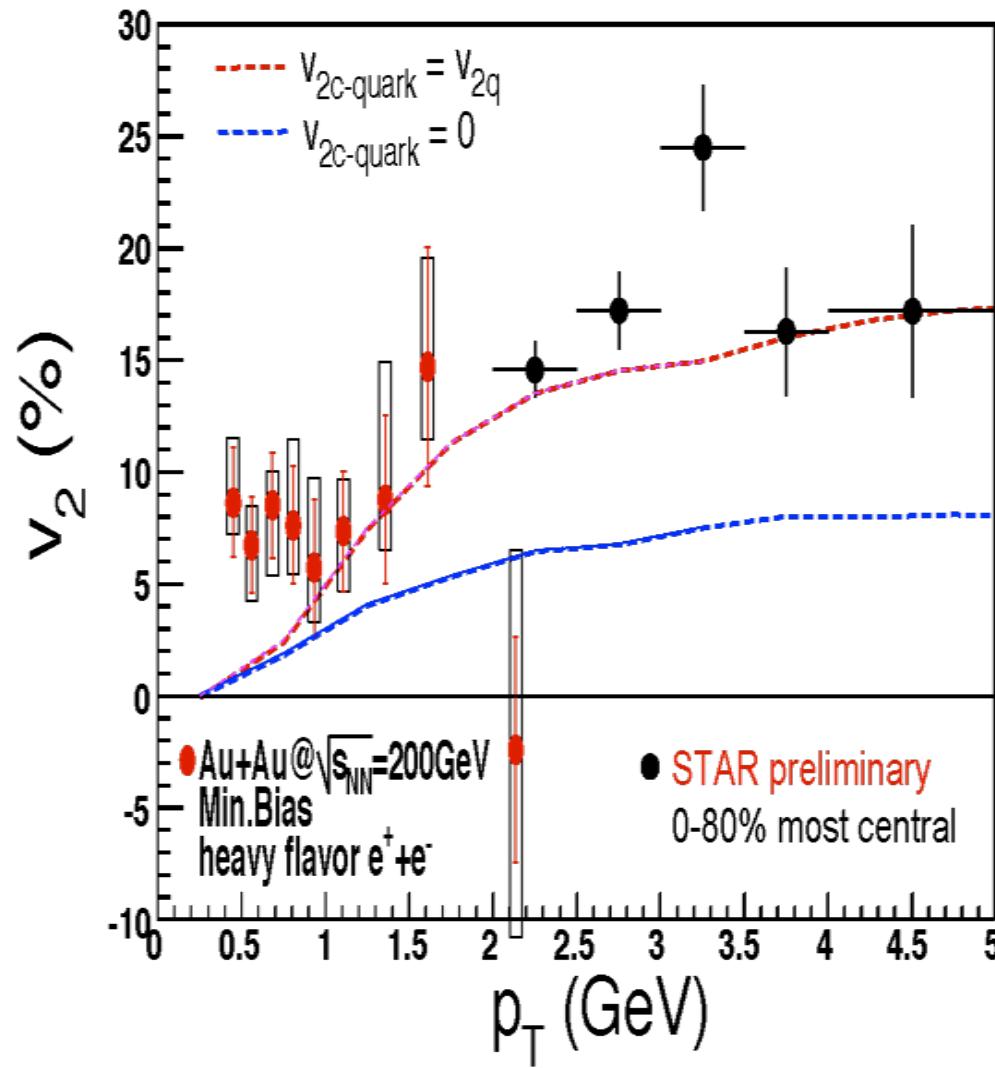


v_2 at LHC in agreement
with (slightly) viscous
relativistic hydrodynamics



Takes a long time $\sim R/c_s$
to build up v_2
Flow must set in very
early (≤ 1 fm)

Even heavy quarks (charm & beauty flow)



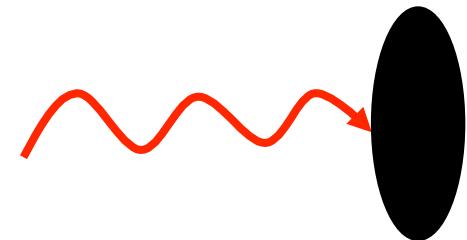
AdS/CFT and the KSS bound

AdS/CFT conjecture: Maldacena

"Holographic Duality between strongly coupled N=4 supersymmetric Yang-Mills theory at large coupling and N_c & classical 10 dimensional gravity in the background of D3 branes

KSS bound: Kovtun,Son,Starinets

Classical absorption cross-section of a graviton with energy ω on a black brane



$$\sigma(\omega) = \frac{8\pi G}{\omega} \int dt d\mathbf{x} e^{i\omega t} \langle [T_{xy}(t, \mathbf{x}), T_{xy}(0, 0)] \rangle$$

From Kubo: $\eta = \frac{\sigma(0)}{16\pi G}$

From Beckenstein: $S = \frac{a}{4G}$

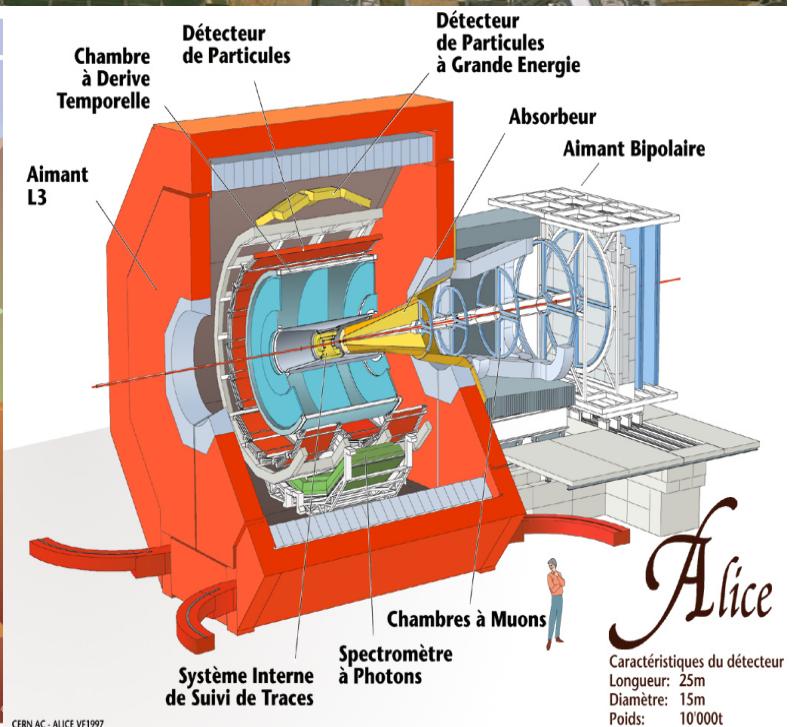
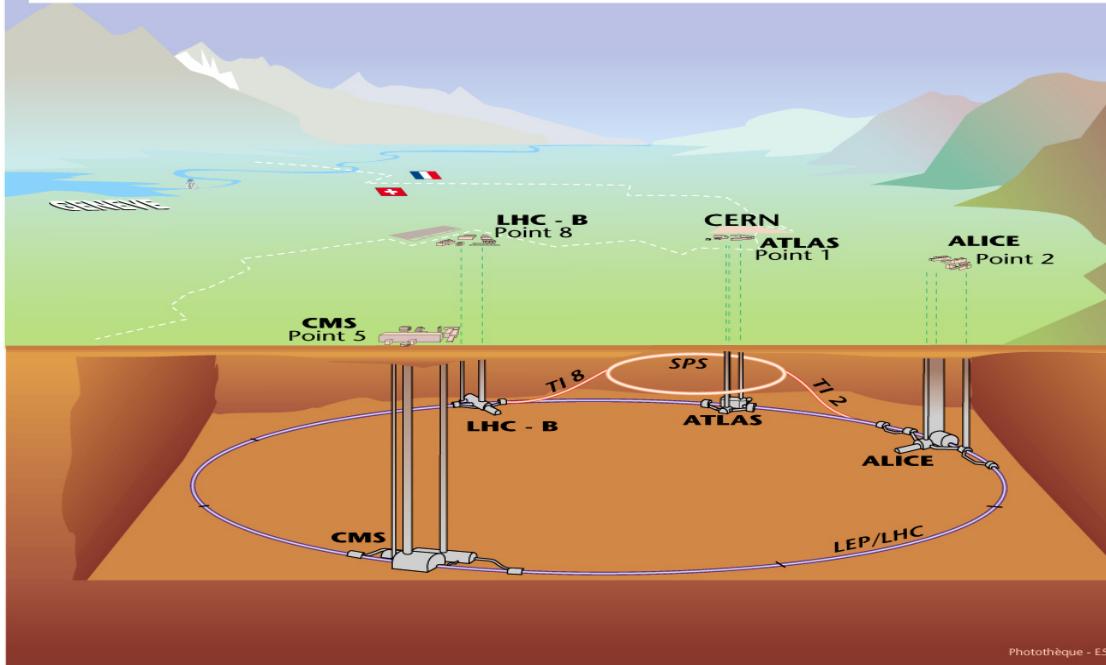
Theorem: $\sigma(0) = a$

where a = area of black brane

Putting it all together: $\frac{\eta}{S} \geq \frac{\hbar}{k_B} \frac{1}{4\pi}$

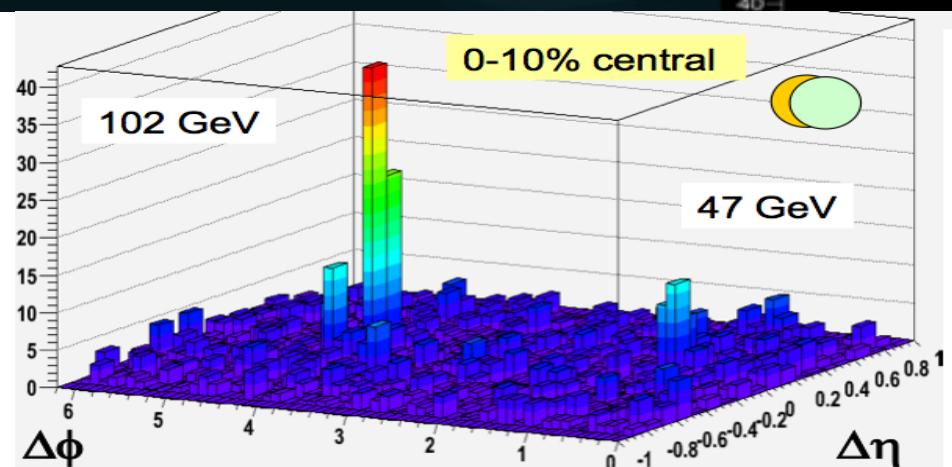
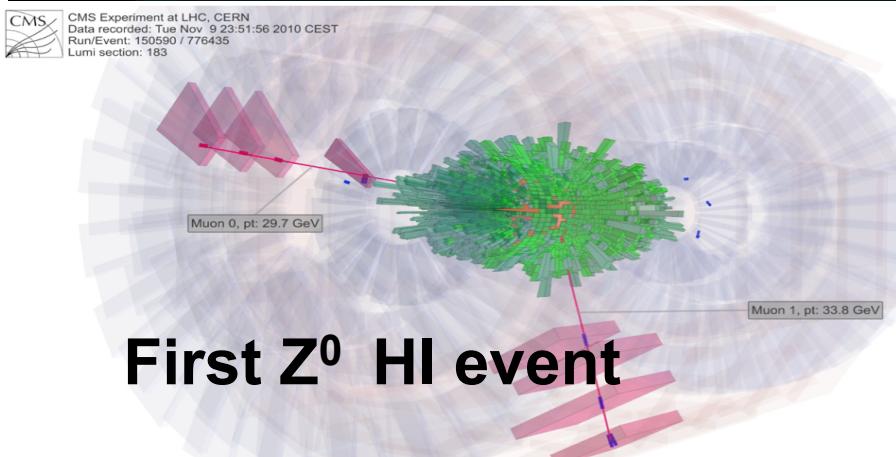
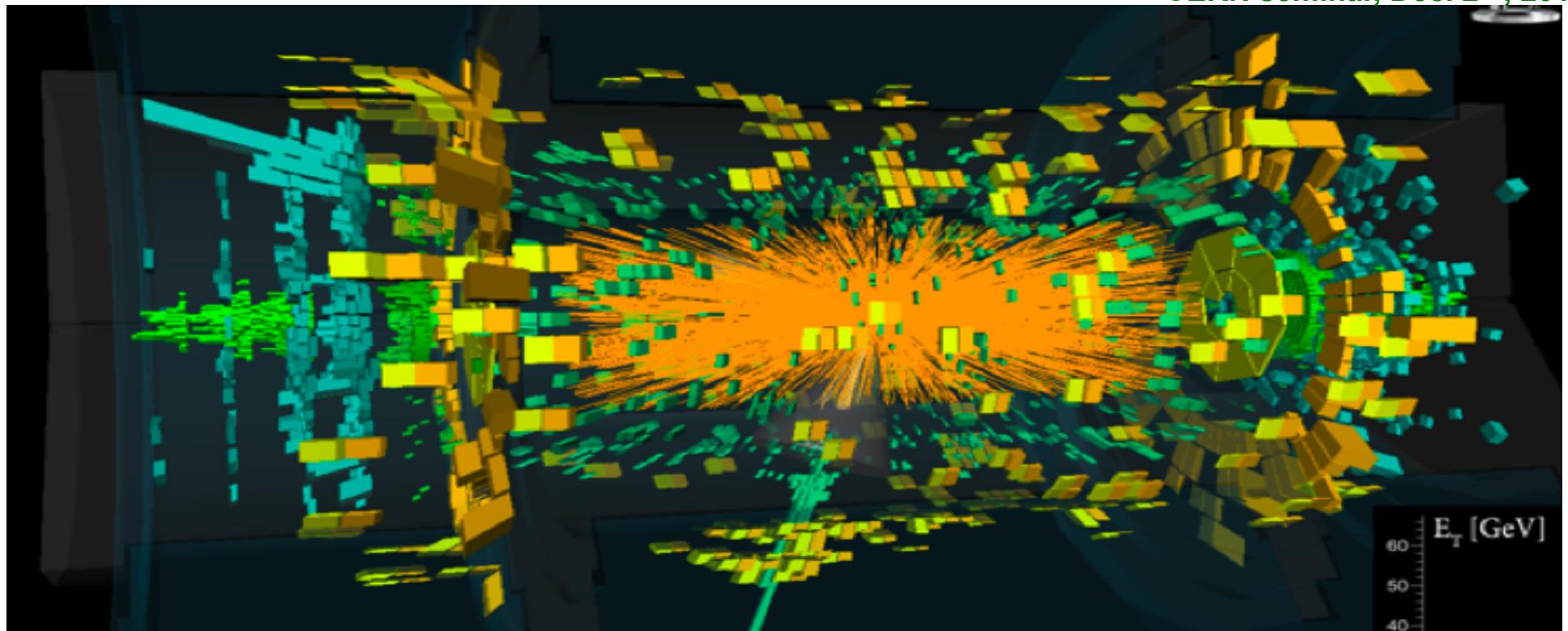


Vue d'ensemble des expériences LHC.

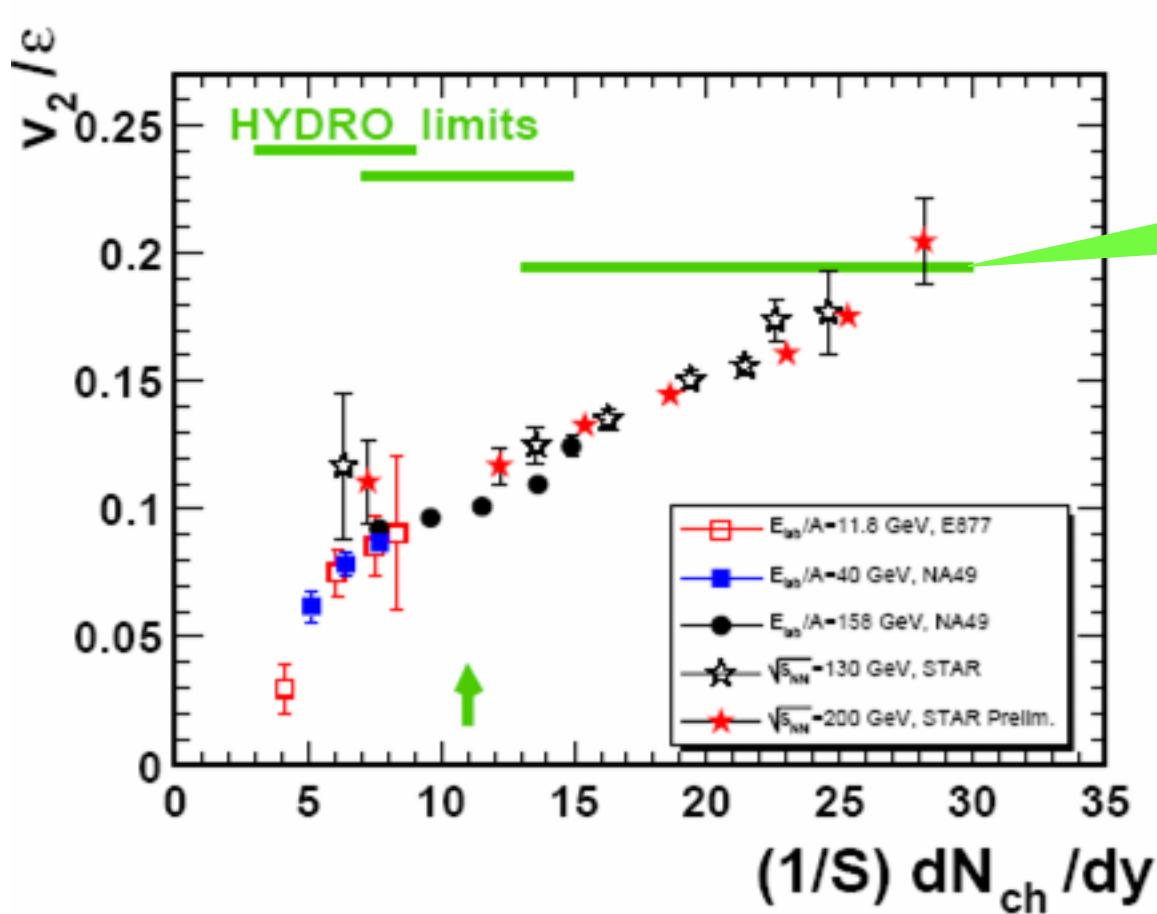


A contemporary view

CERN seminar, Dec. 2nd, 2010



Strong flow = (nearly) ideal hydrodynamics

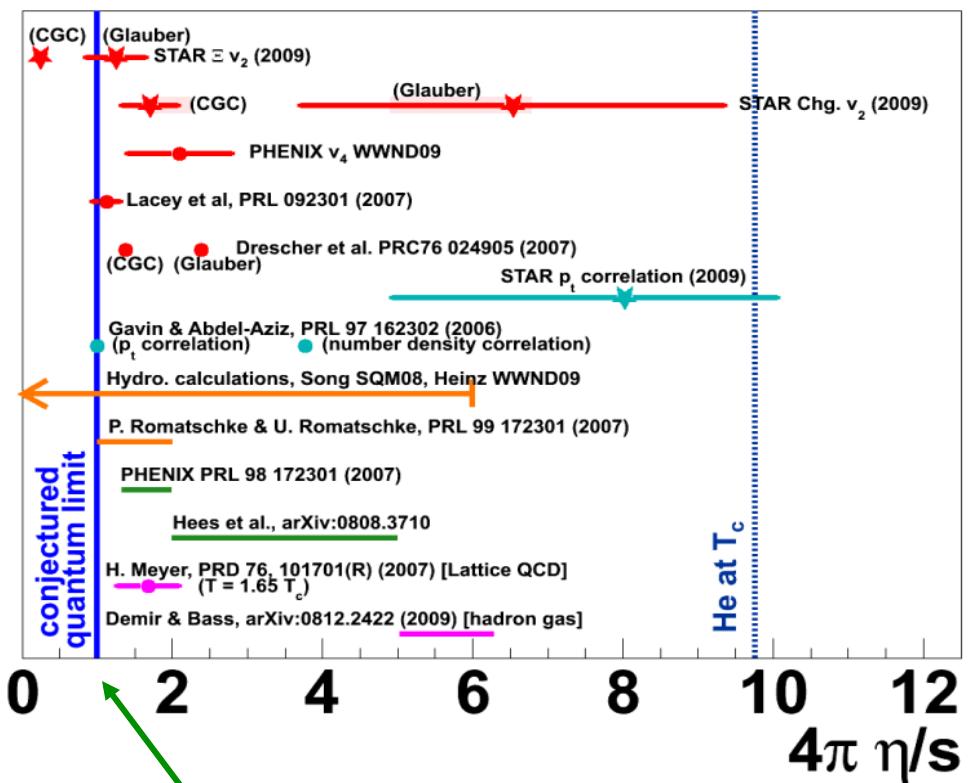
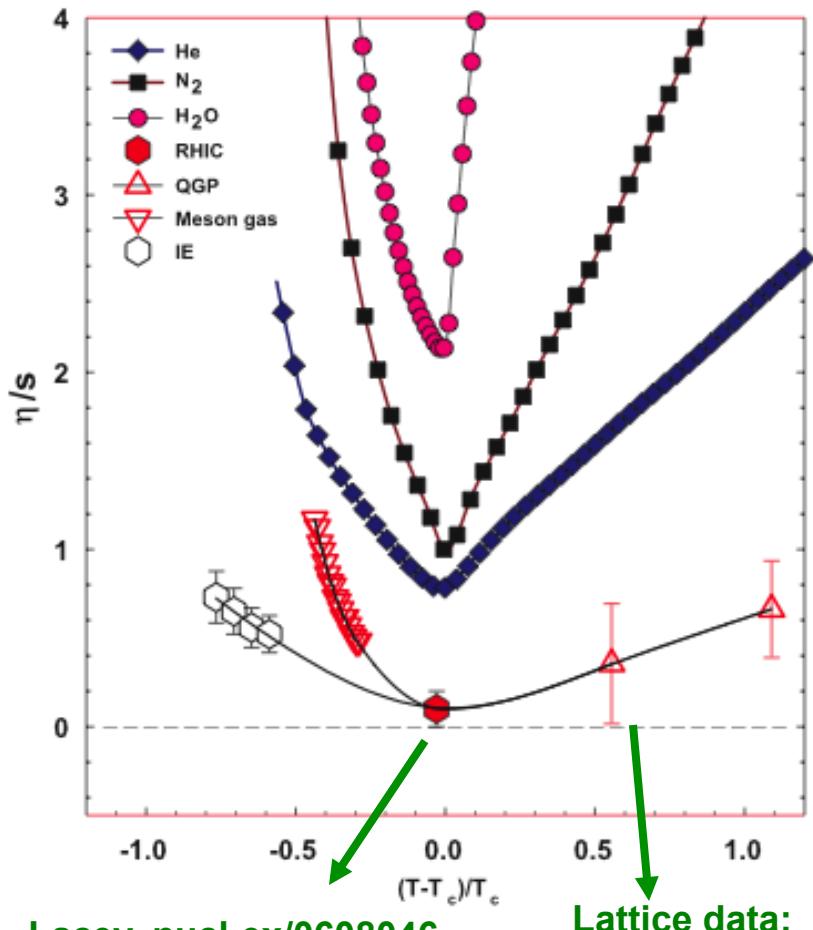


LHC !

CERN Press release, November 26, 2010:
'confirms that the much hotter plasma produced at the LHC behaves as a very low viscosity liquid (a perfect fluid)..'

v_2 at RHIC and the LHC is large

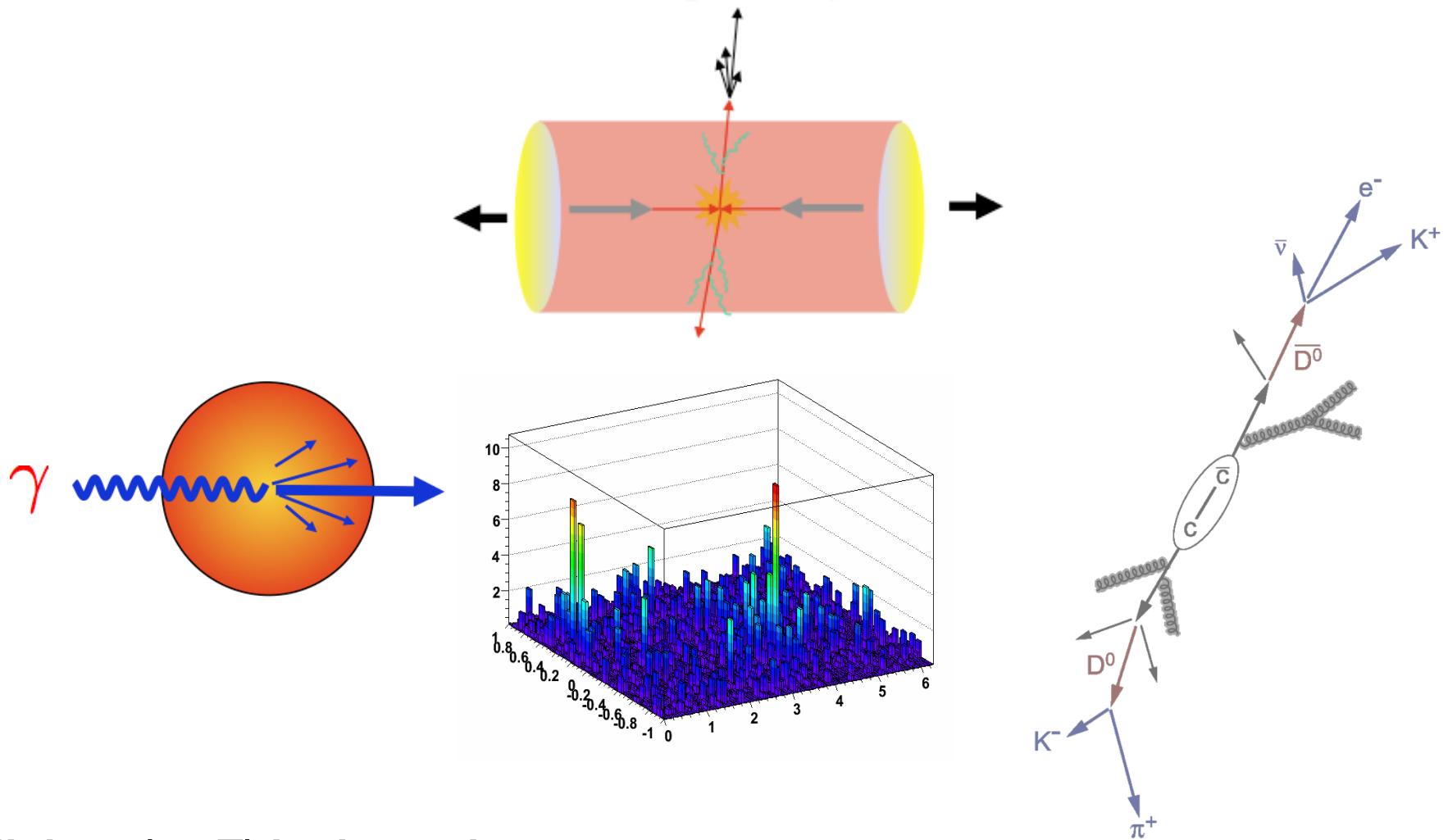
Dissipation in the QGP relative to other fluids



RHIC estimates close to conjectured universal lower bound of $\eta/s = 1/4\pi$ (based on the AdS/CFT correspondence)

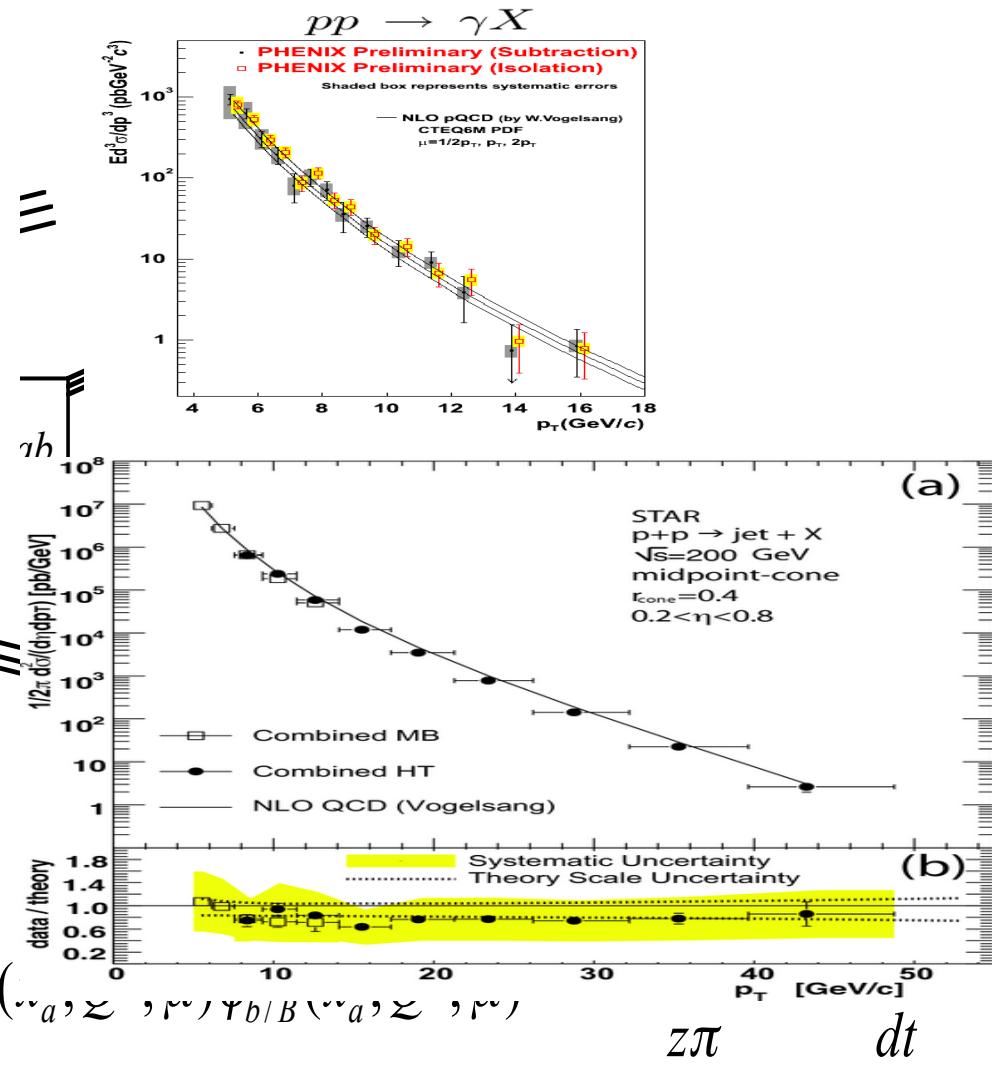
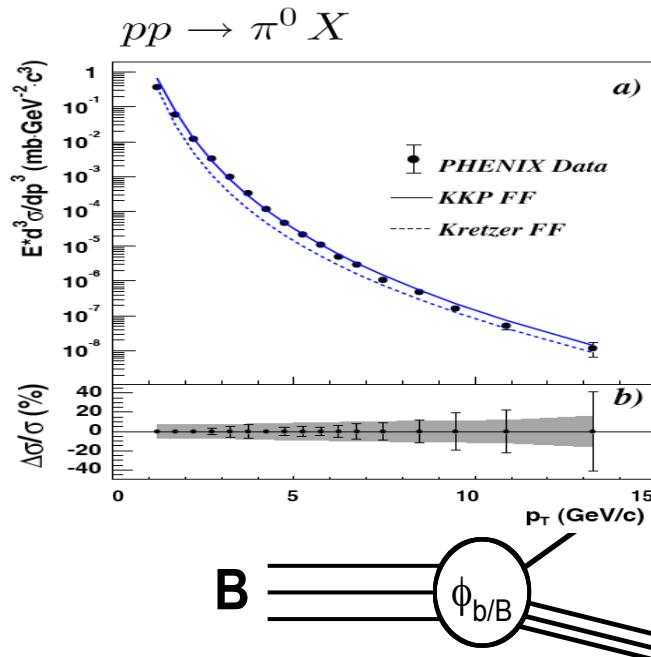
Kovtun, Son, Starinets

Hard Probes of the sQGP



**High p_T ($>> T$) hadrons, heavy mesons
(charm and beauty), J/ψ and Upsilon, hard photons and JETS...**

The pp Benchmark for Heavy Ion Collisions

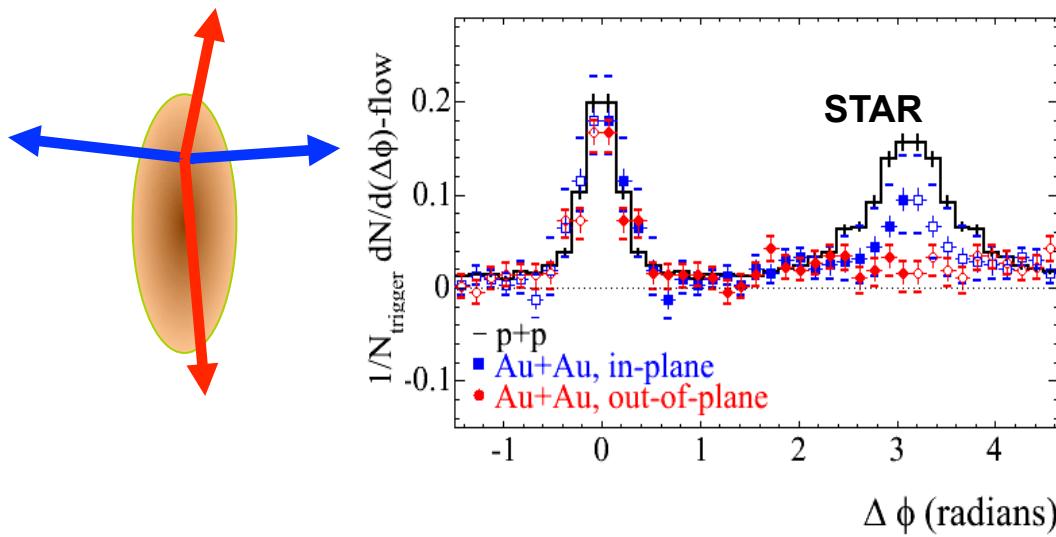
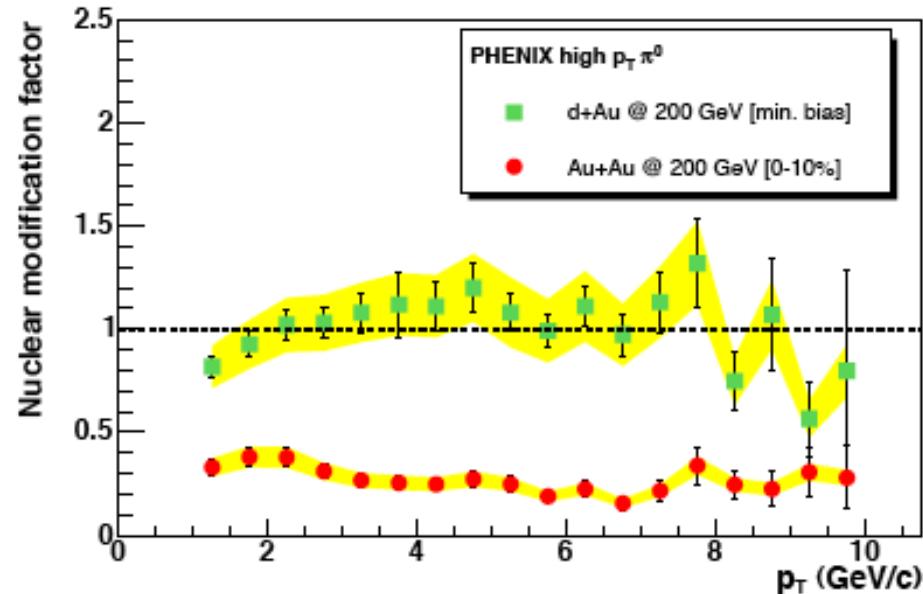


$$E \frac{d^3\sigma}{dp^3} = \sum_{abc} \int dx_a dx_b \phi_{a/A}(\zeta_a, \zeta_b, \tau_b/B) \phi_{b/B}(\zeta_a, \zeta_b, \tau_b)$$

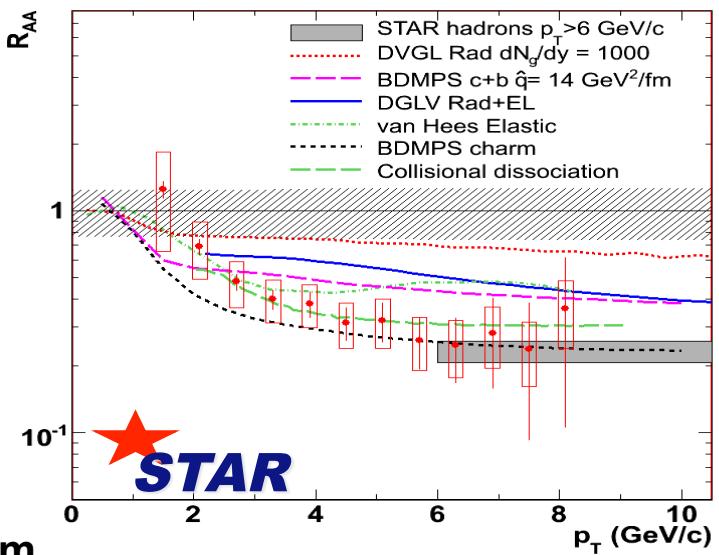
pQCD works exceptionally well!

Hard probes are modified by strongly interacting medium

$$R_{AA} = \frac{\text{Yield}_{\text{AuAu}} / \langle N_{\text{binary}} \rangle_{\text{AuAu}}}{\text{Yield}_{pp}}$$

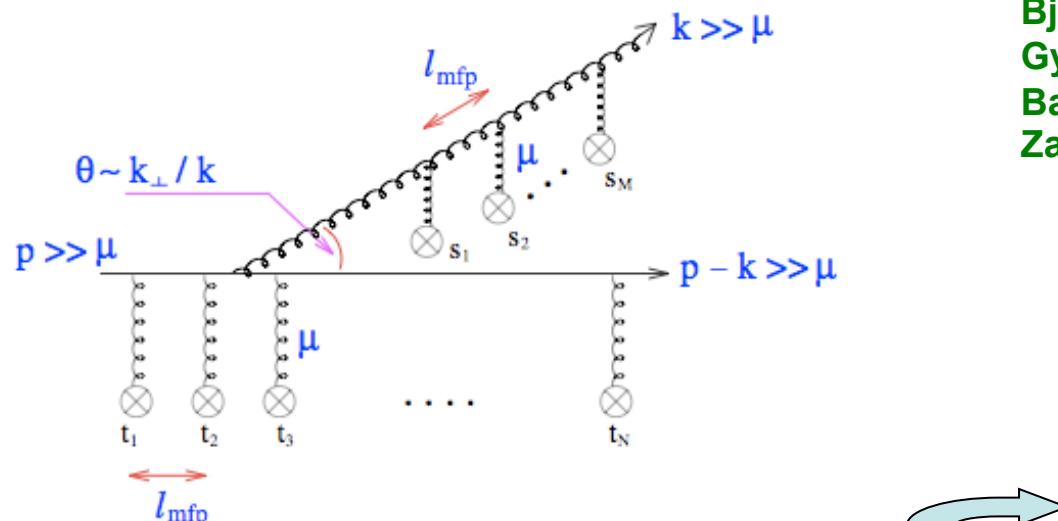


Energy loss dependent on parton path length in medium



Heavy quarks suppressed just as much as light quarks!

Energy loss in hot QCD matter

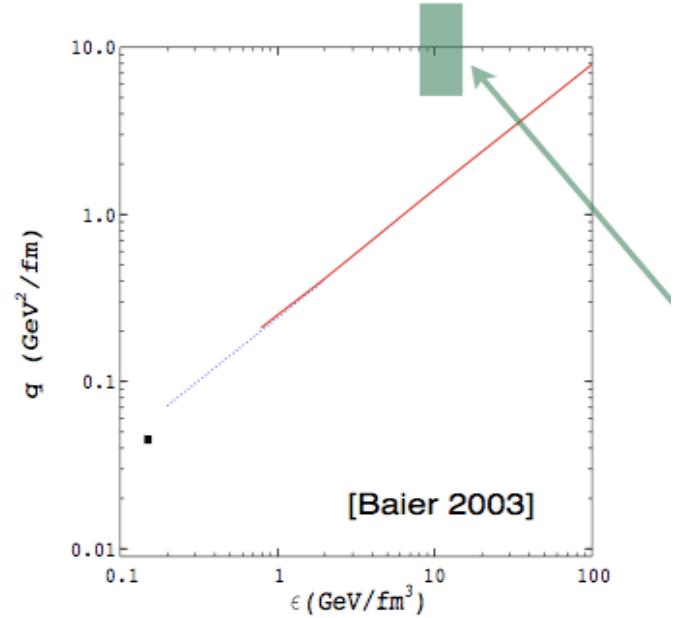


$$-\frac{dE}{dz} = \frac{\alpha_s N_c}{4}, p_{\perp}^2 ; \quad p_{\perp}^2 = \hat{q} L$$

$$\hat{q}_{\text{ideal gas}} \simeq \frac{72}{\pi} \xi(3) \alpha_s^2 T^3 \simeq 2 \epsilon^{3/4}$$

Bjorken
Gyulassy-Wang
Baier,Dokshitzer,Mueller,Peigne,Schiff
Zakharov

Characterizes properties
of hot medium

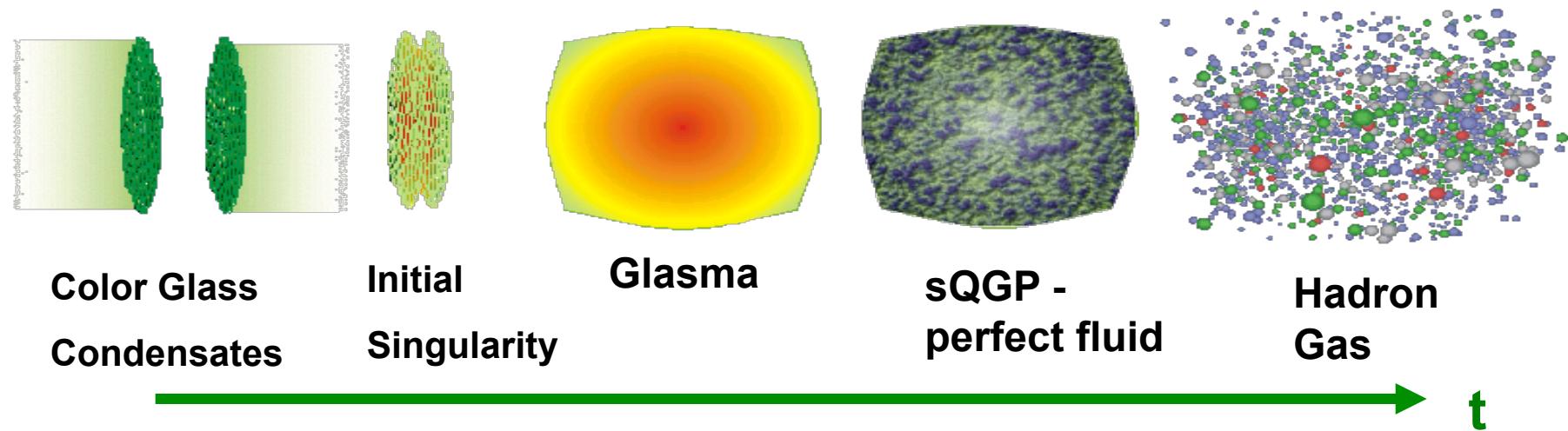


Detailed model estimates all over the place

$$\hat{q} = 1.5 - 19 \text{ GeV}^2/\text{fm}$$

Significantly higher than ideal gas!

Standard model of HI Collisions





QGP