

A long-exposure photograph of the night sky, showing numerous concentric circular star trails. The trails are centered on a point in the sky, likely the celestial pole. In the foreground, the dark silhouette of the South African Astronomical Observatory is visible, including a prominent dome and a tall, thin structure. The sky is a deep blue-black, and the star trails are a mix of white and light blue. The overall scene is a classic astronomical time-lapse.

High time-domain Astrophysics with SALT

Stephen Potter

South African Astronomical Observatory

Outline

SALTICAM

Observing modes

Results

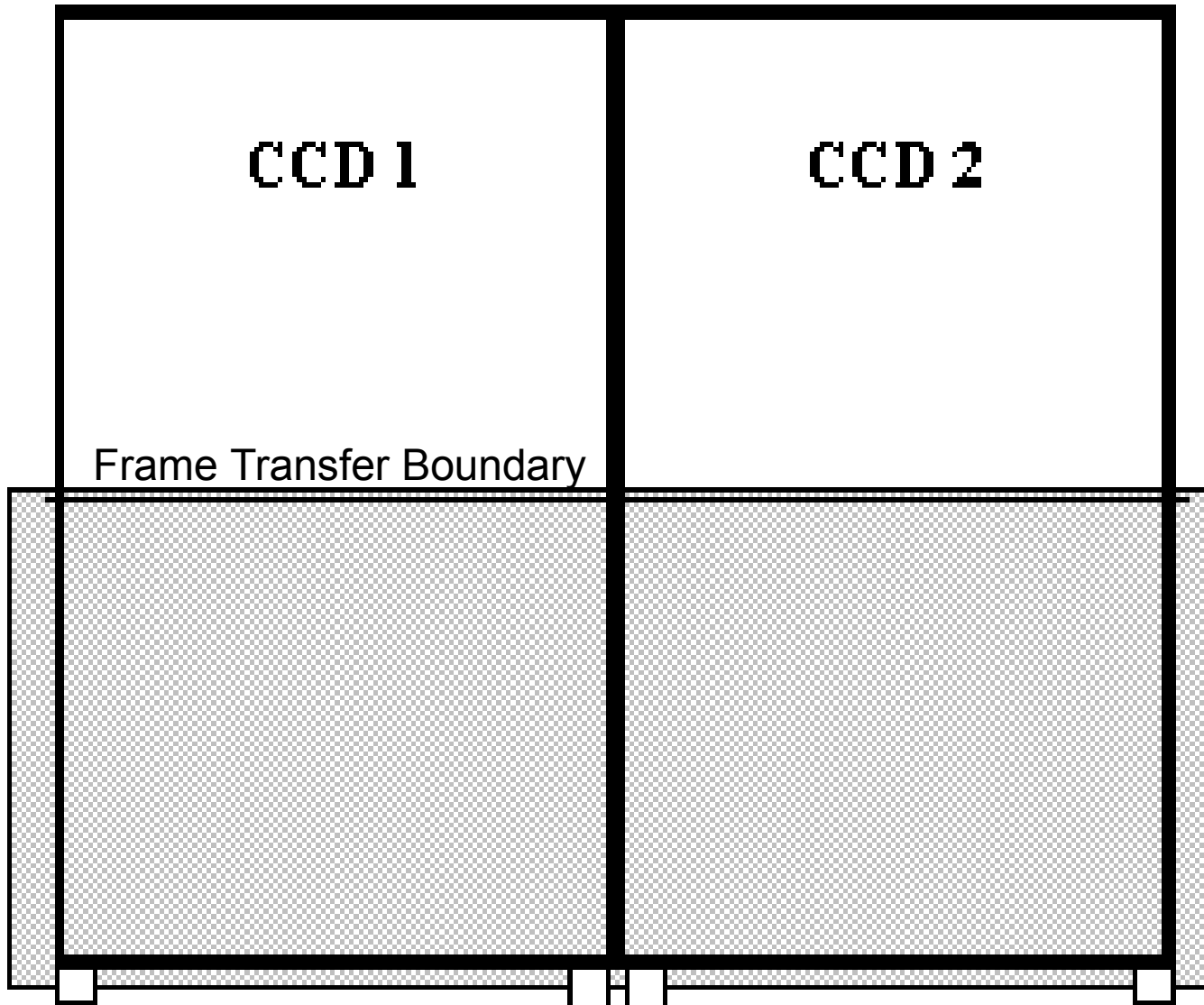
Future

RSS

Observing modes

Examples

SALTICAM



F.O.V 8 arcmins

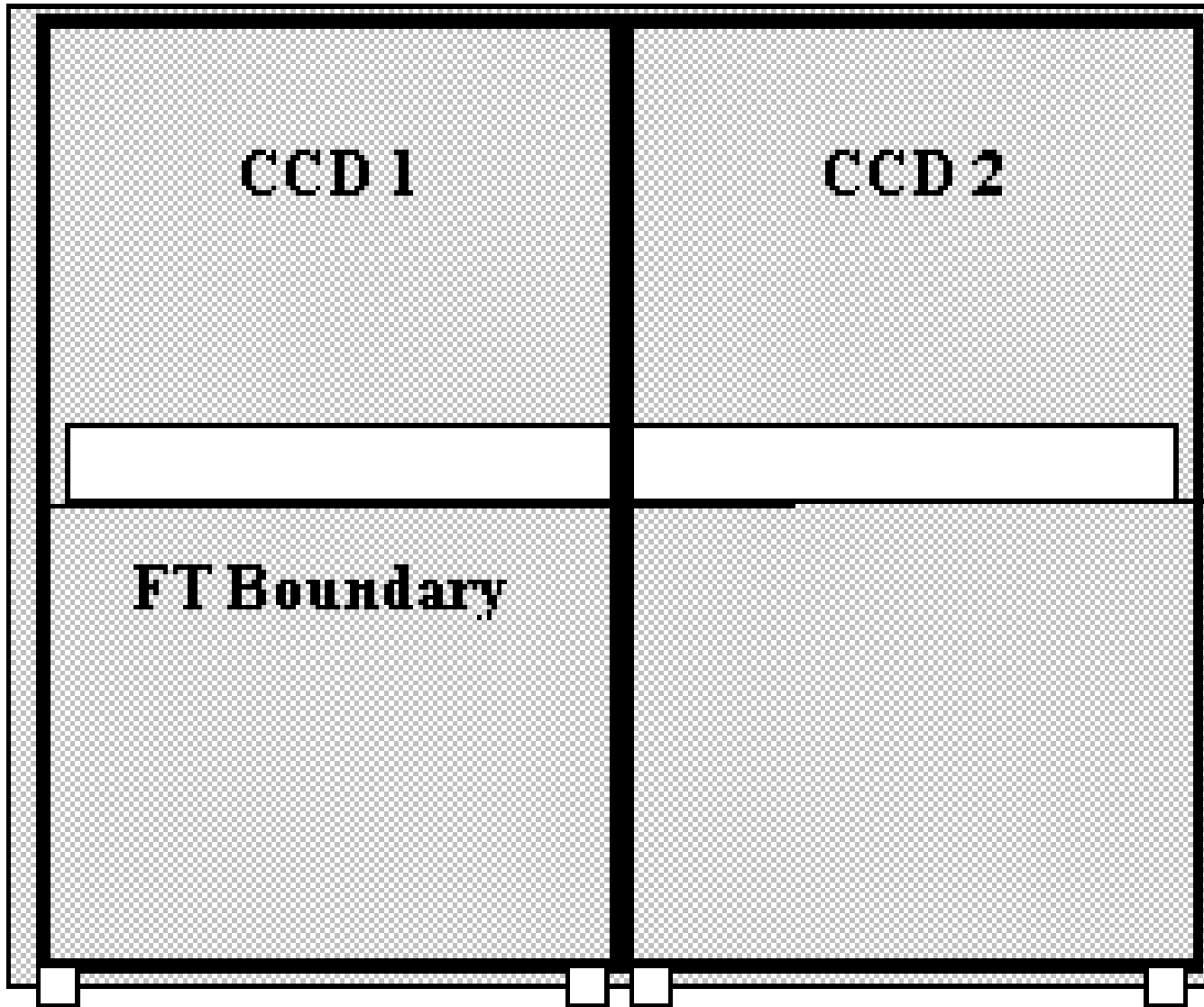
frame transfer boundary

frame transfer mask

4 read-out amplifiers

2x2, 4x4, 6x6'

SALTICAM



F.O.V 8'x20"

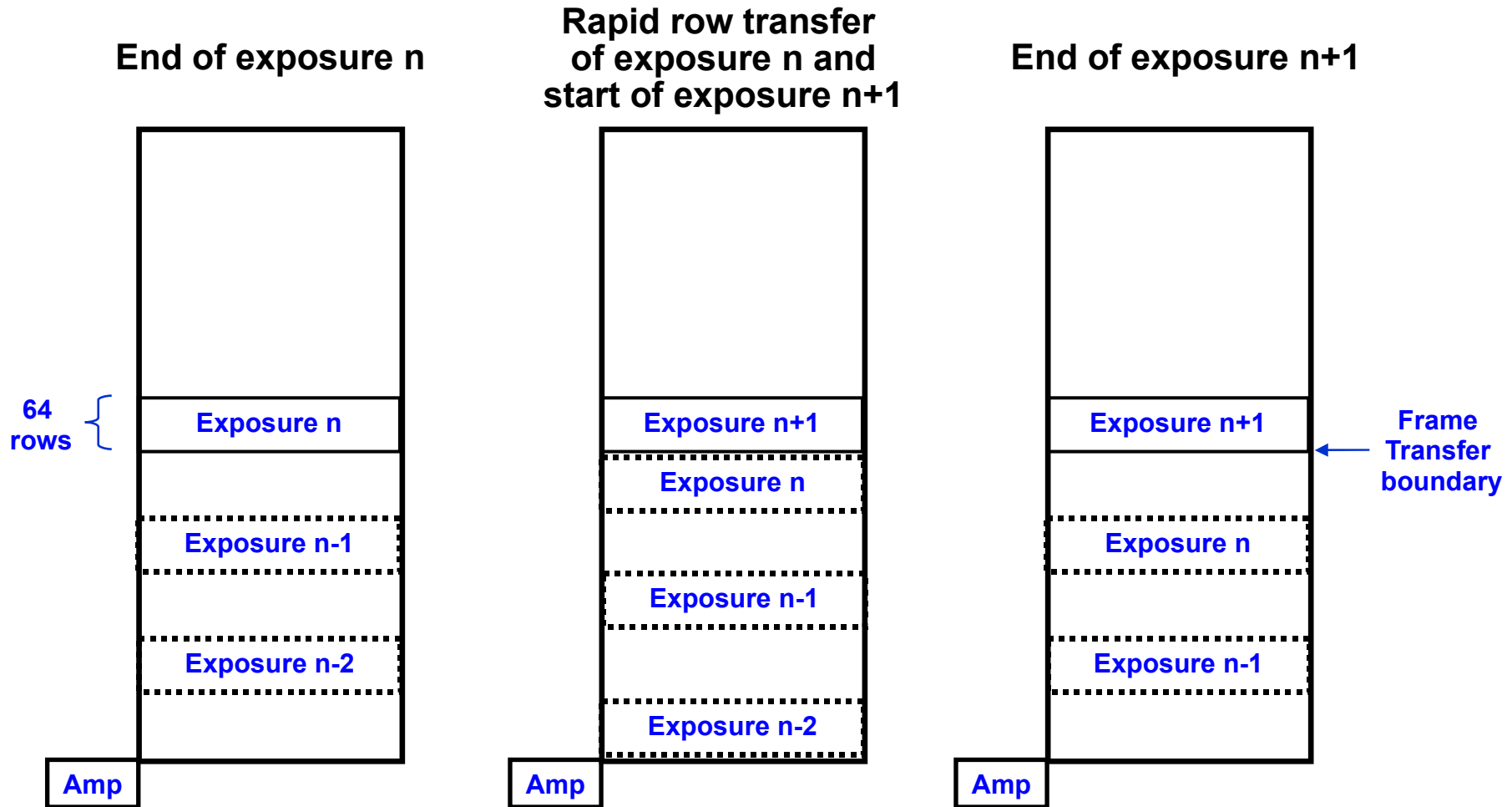
frame transfer boundary

slot mask

4 read-out amplifiers

2x2, 4x4, 6x6'... 0.08s

SALTICAM

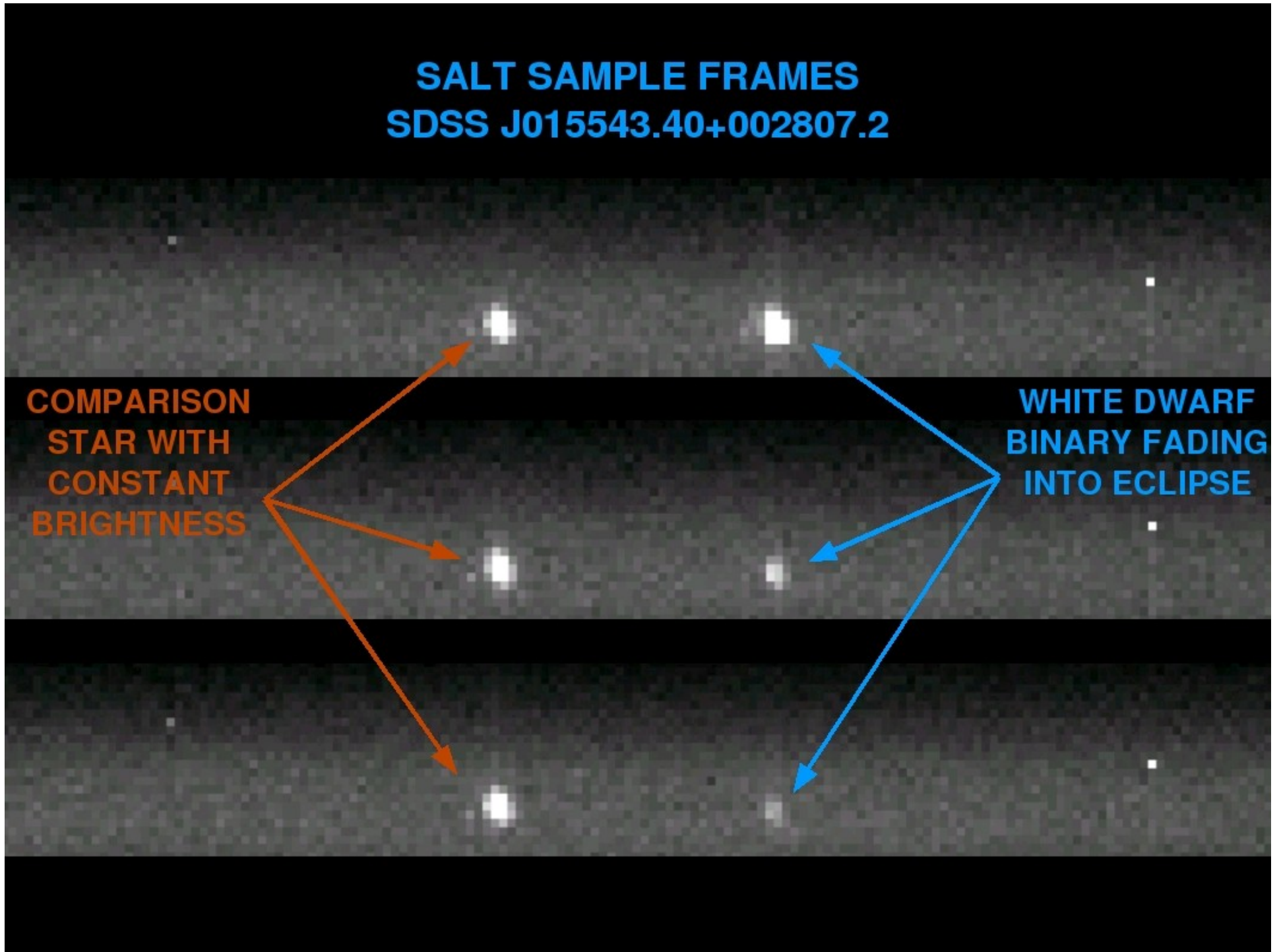


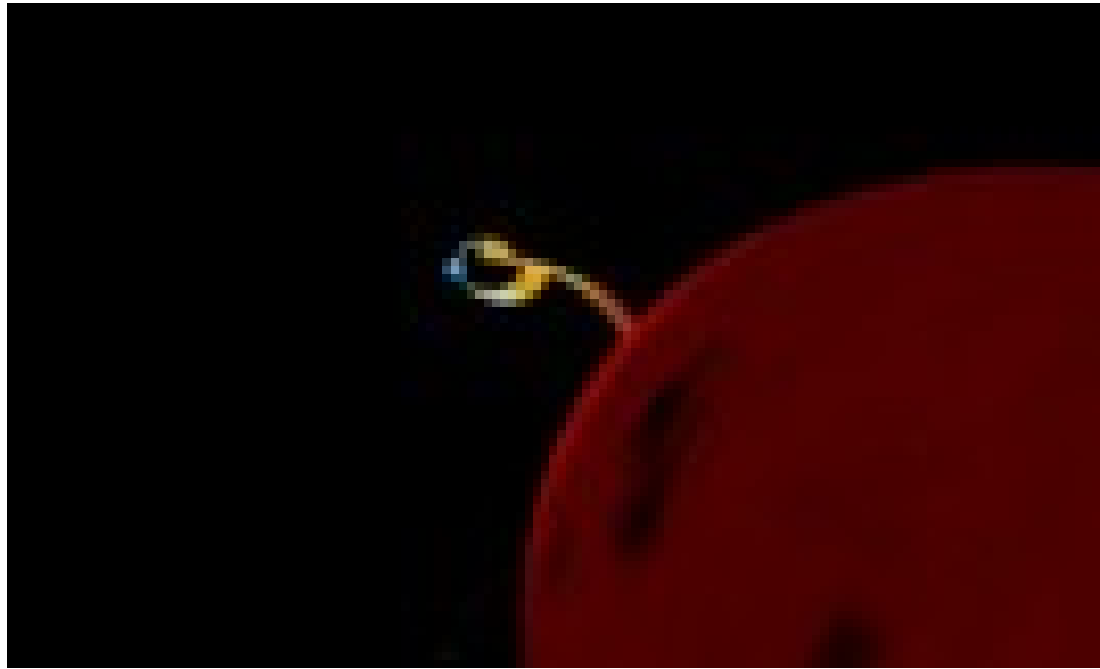
Control scheme for fastest sampling

SALT SAMPLE FRAMES
SDSS J015543.40+002807.2

**COMPARISON
STAR WITH
CONSTANT
BRIGHTNESS**

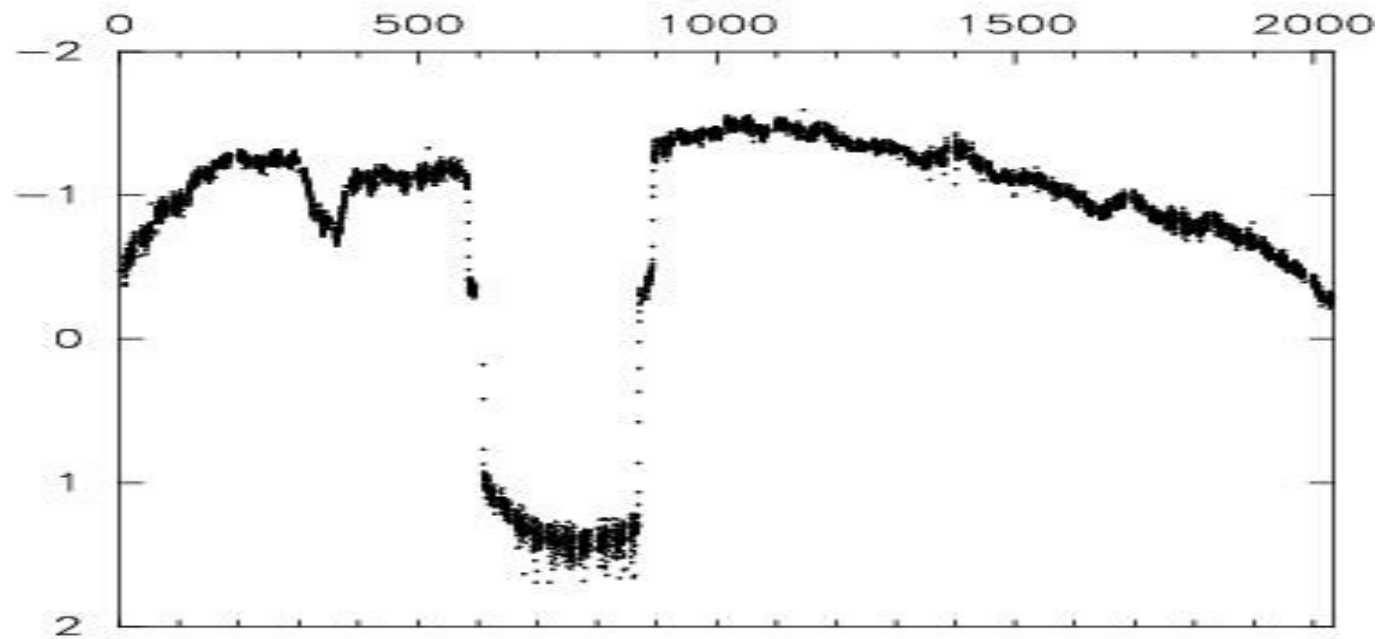
**WHITE DWARF
BINARY FADING
INTO ECLIPSE**

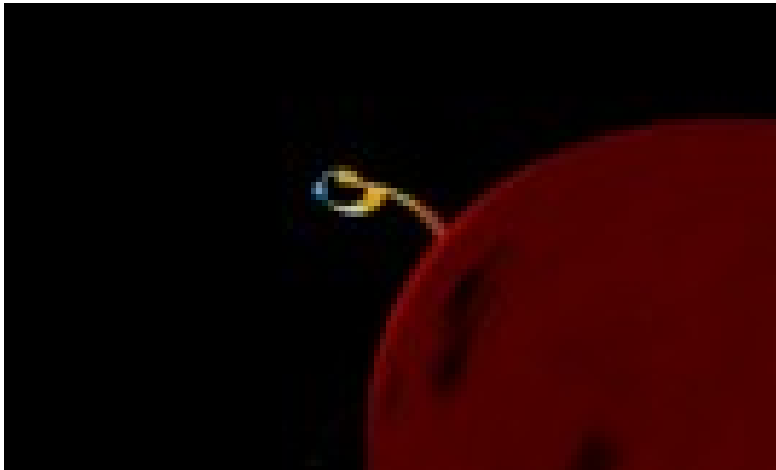




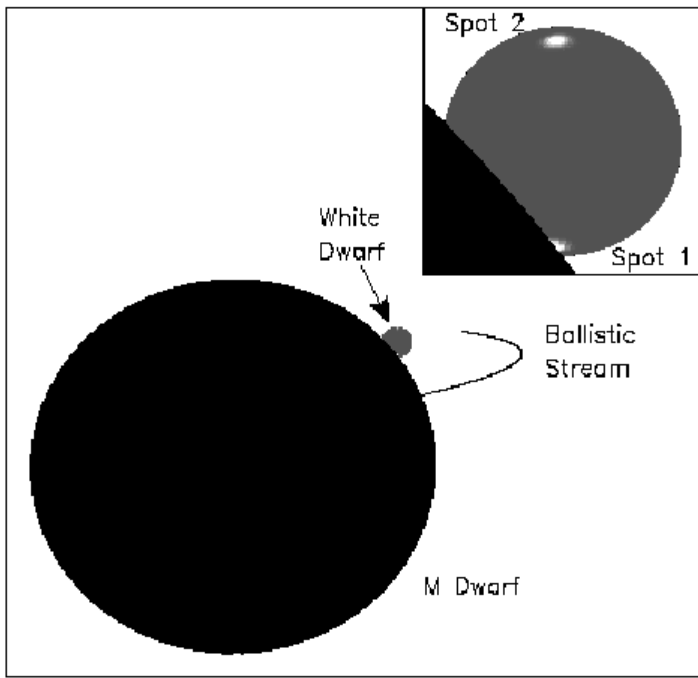
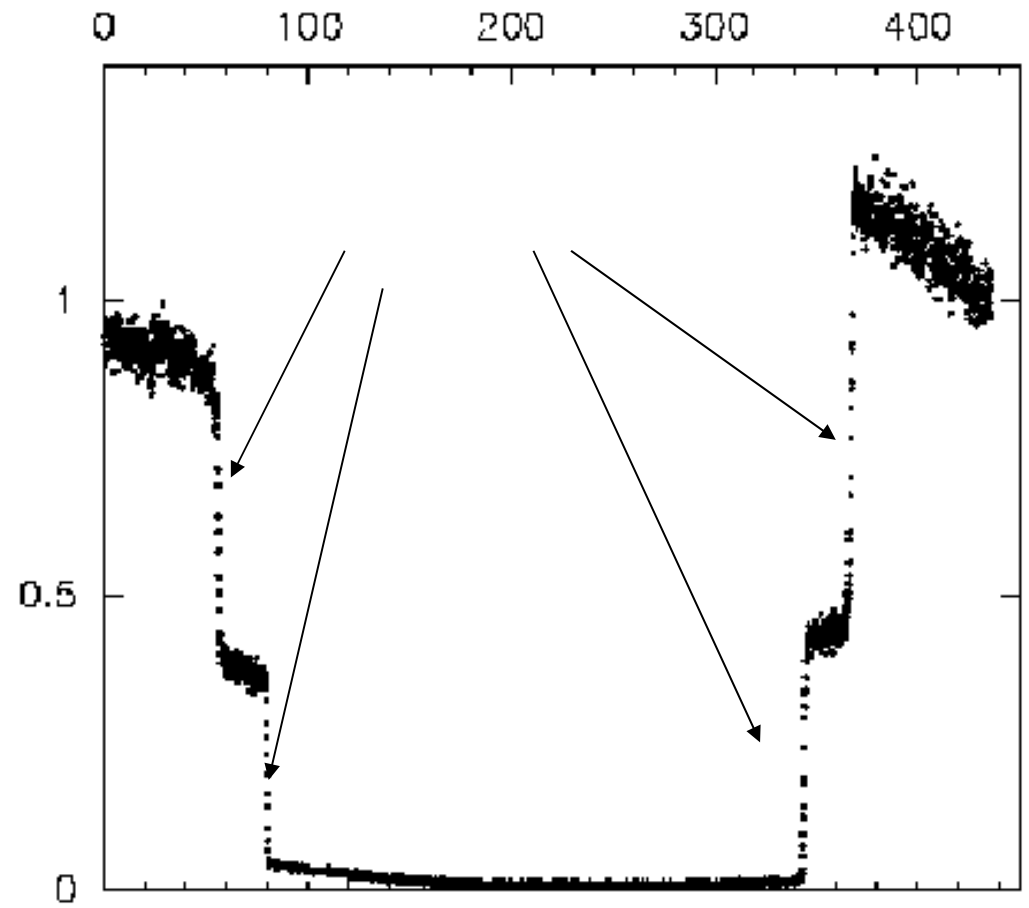
SDSS015543+002807

(Sec)



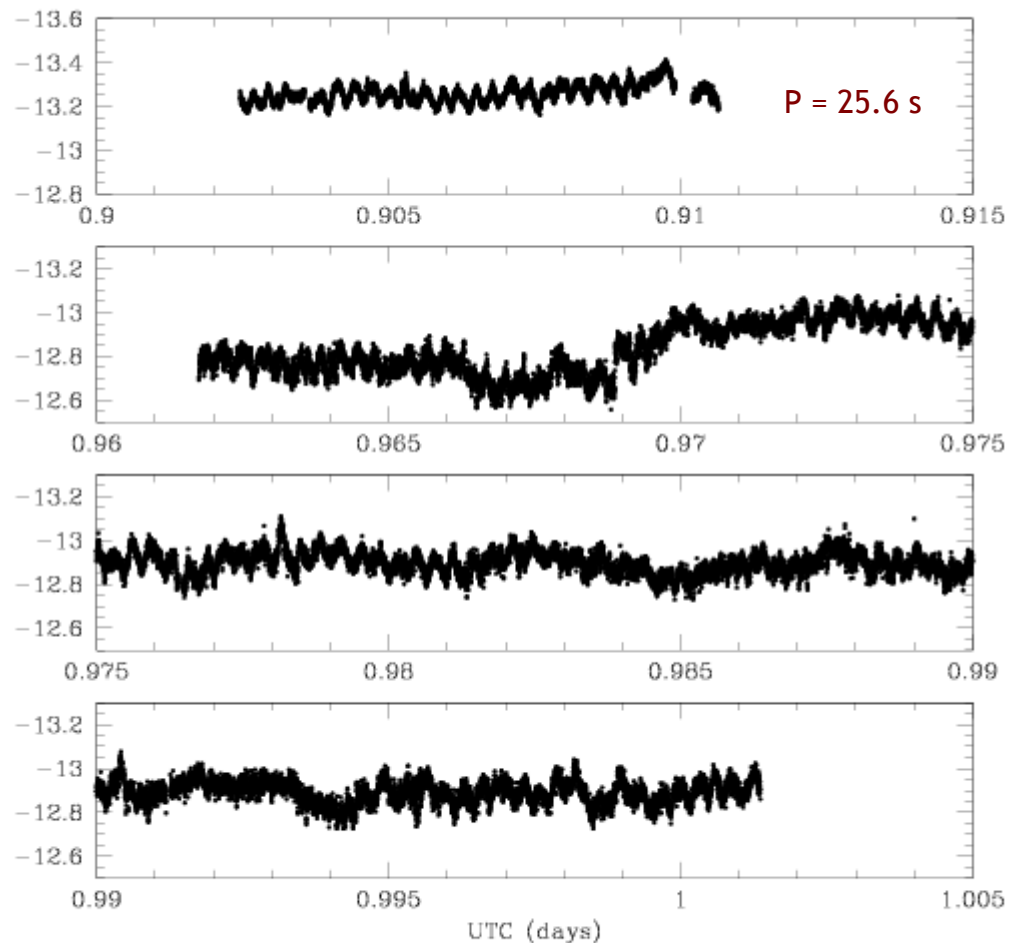


SDSS015543+002807 2005 Sep 6 Eclipse
(Sec)

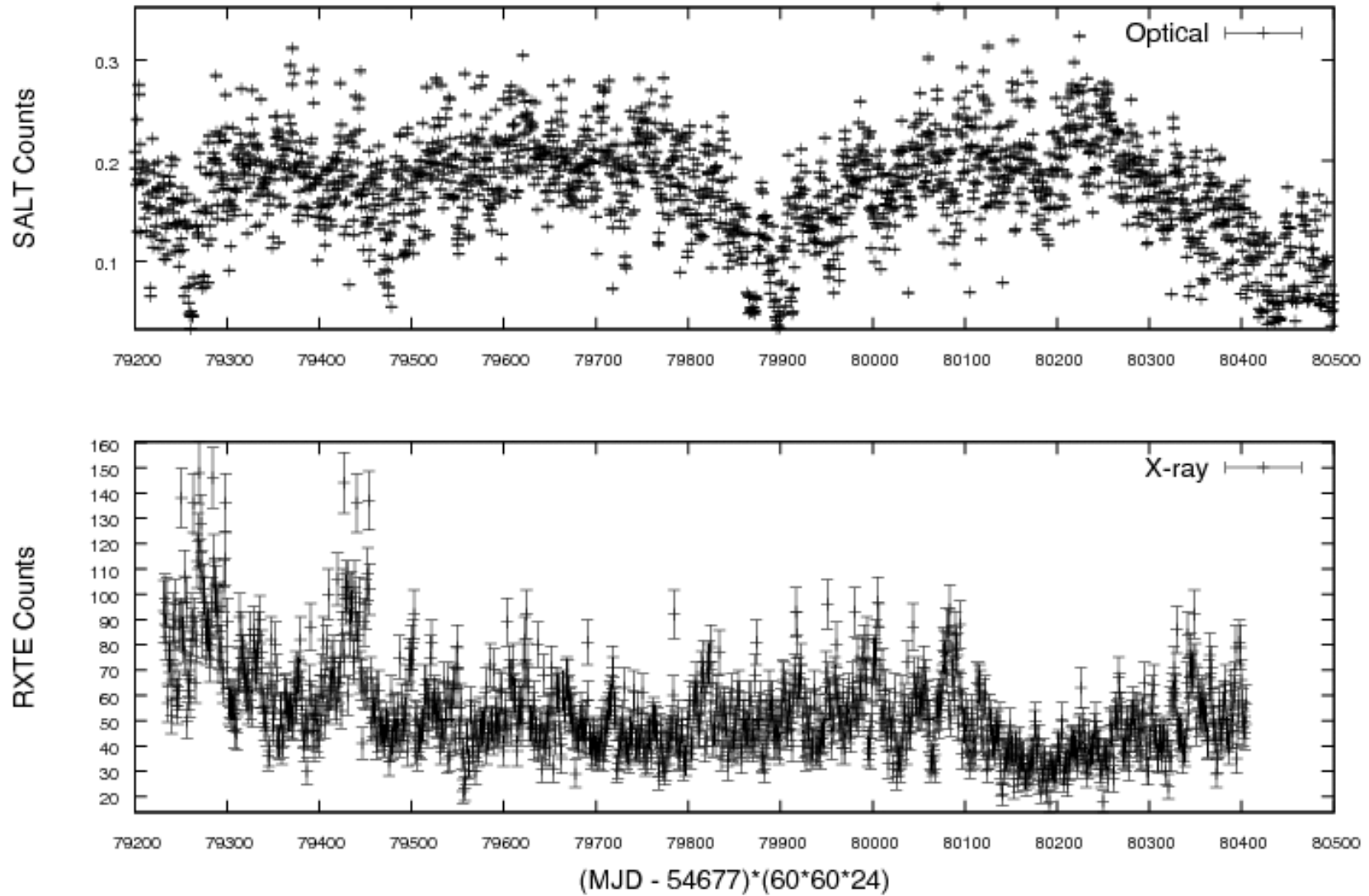


Rapid variability in cataclysmic variable stars: [VW Hydri](#) dwarf nova oscillation (DNO) / quasi-periodic oscillation (QPO)

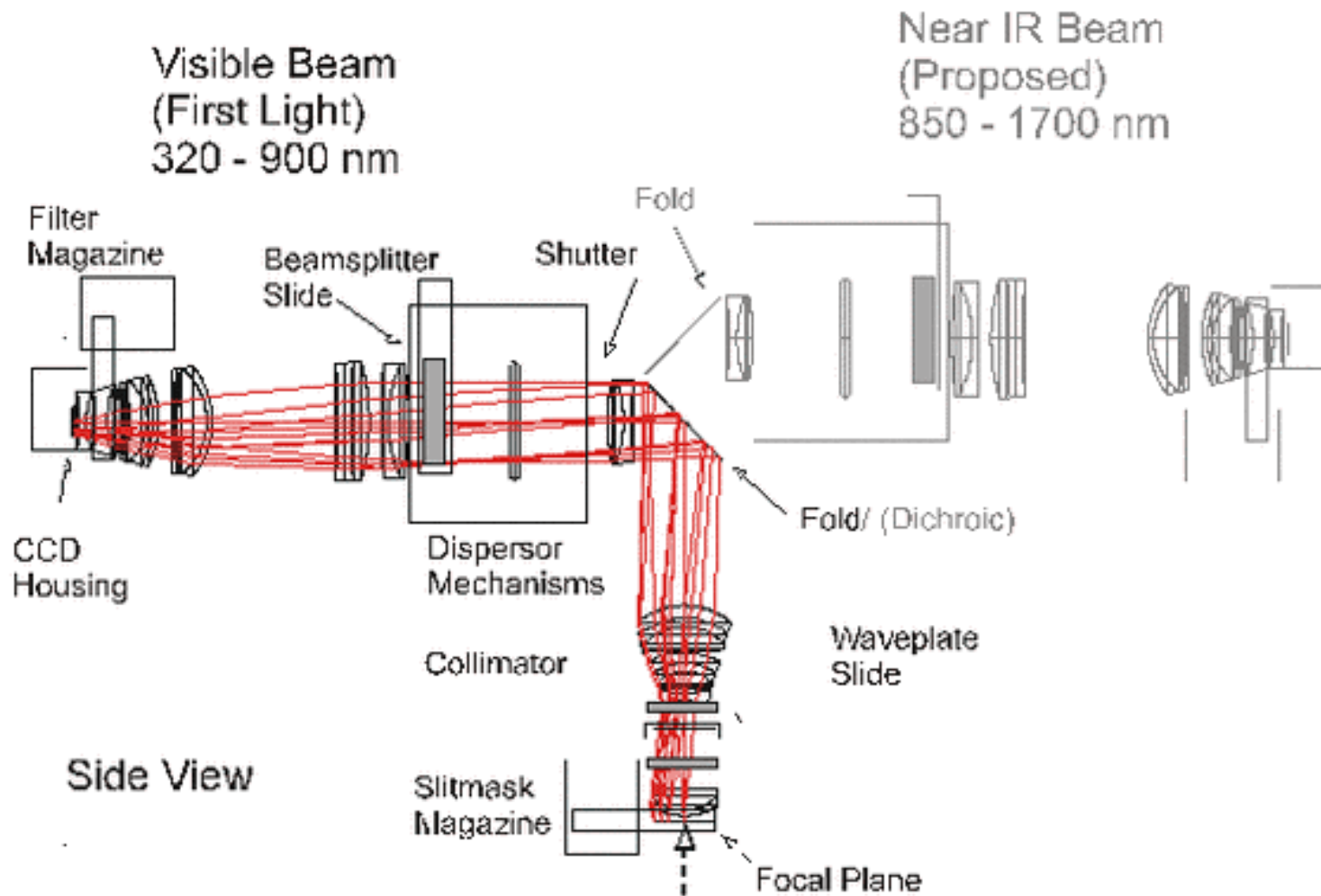
V = 13.3 mag, Filter = white light, Binning = 6x6, $T_{\text{exp}} = 80$ ms
Photometry: 2.5 pix + aperture correction (relative)



Optical (SALT) and X-ray (RXTE) Light curves of GX339-4 (20080730 , MJD 54677)



Robert Stobie Spectrograph (RSS)



Gratings
Beamsplitter
 $\frac{1}{2}$ and $\frac{1}{4}$ waveplates
Masks (e.g. slit masks)
Filters

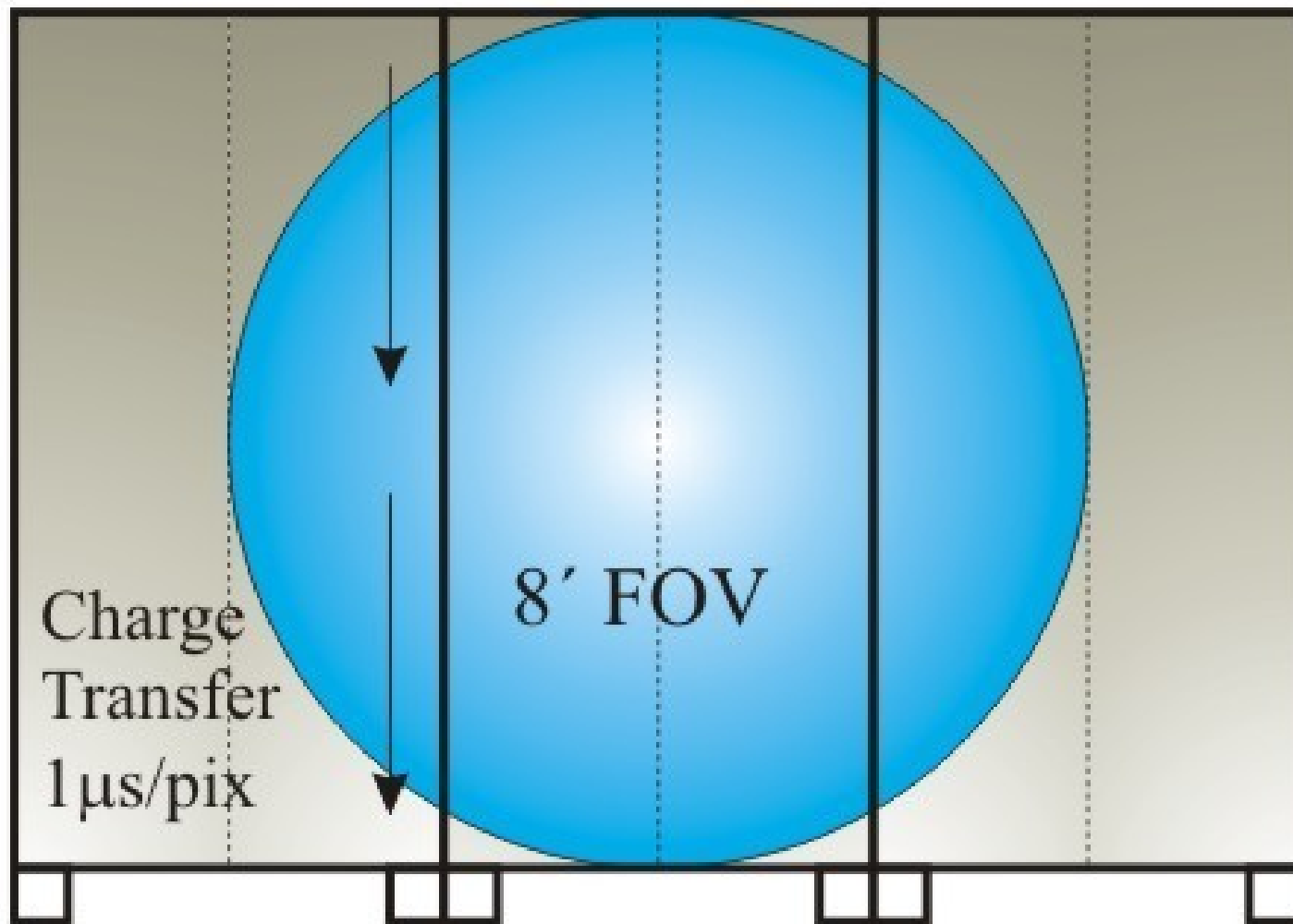
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Charge
Transfer
1 μs/pix

8' FOV

Amplifier

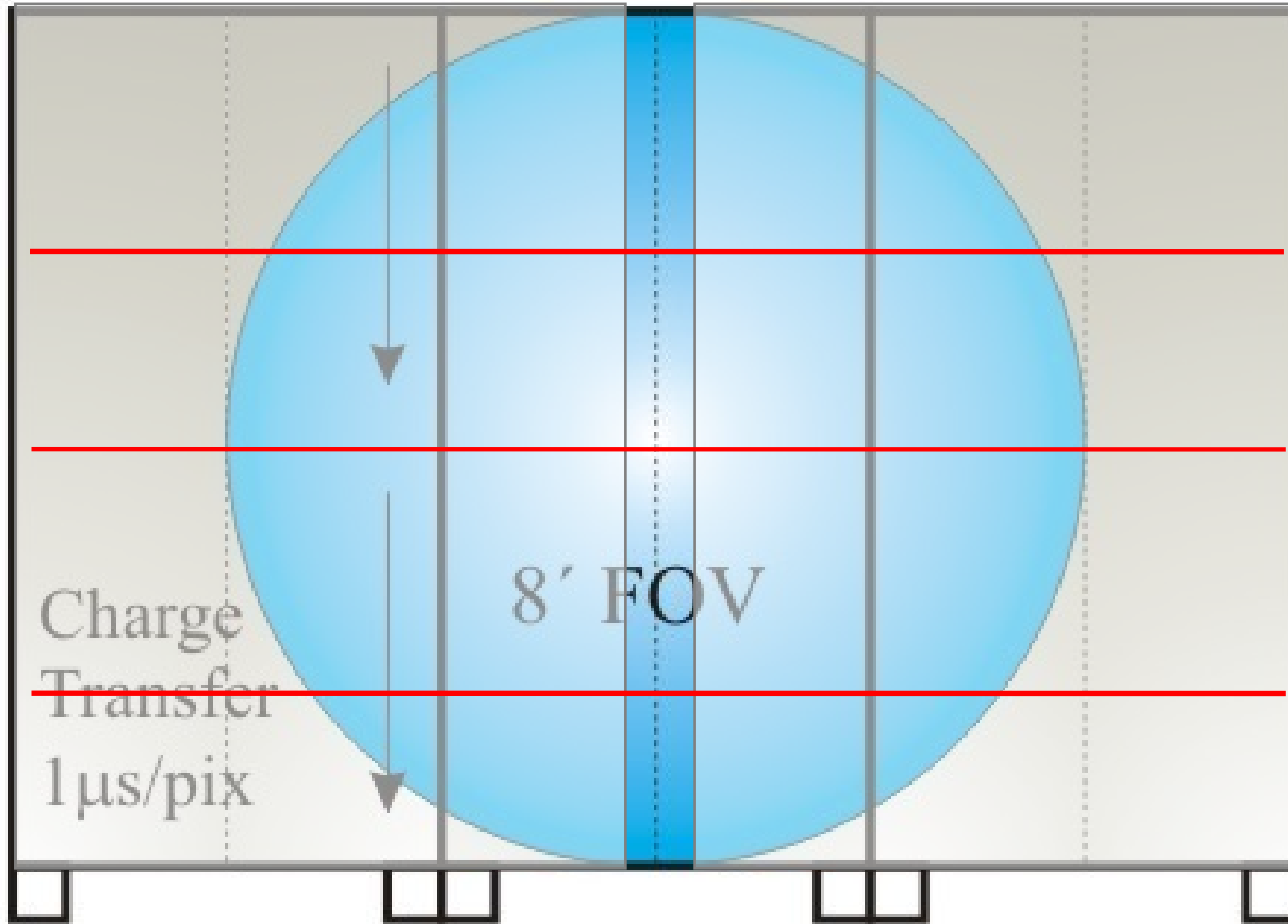
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Long slit spectroscopy

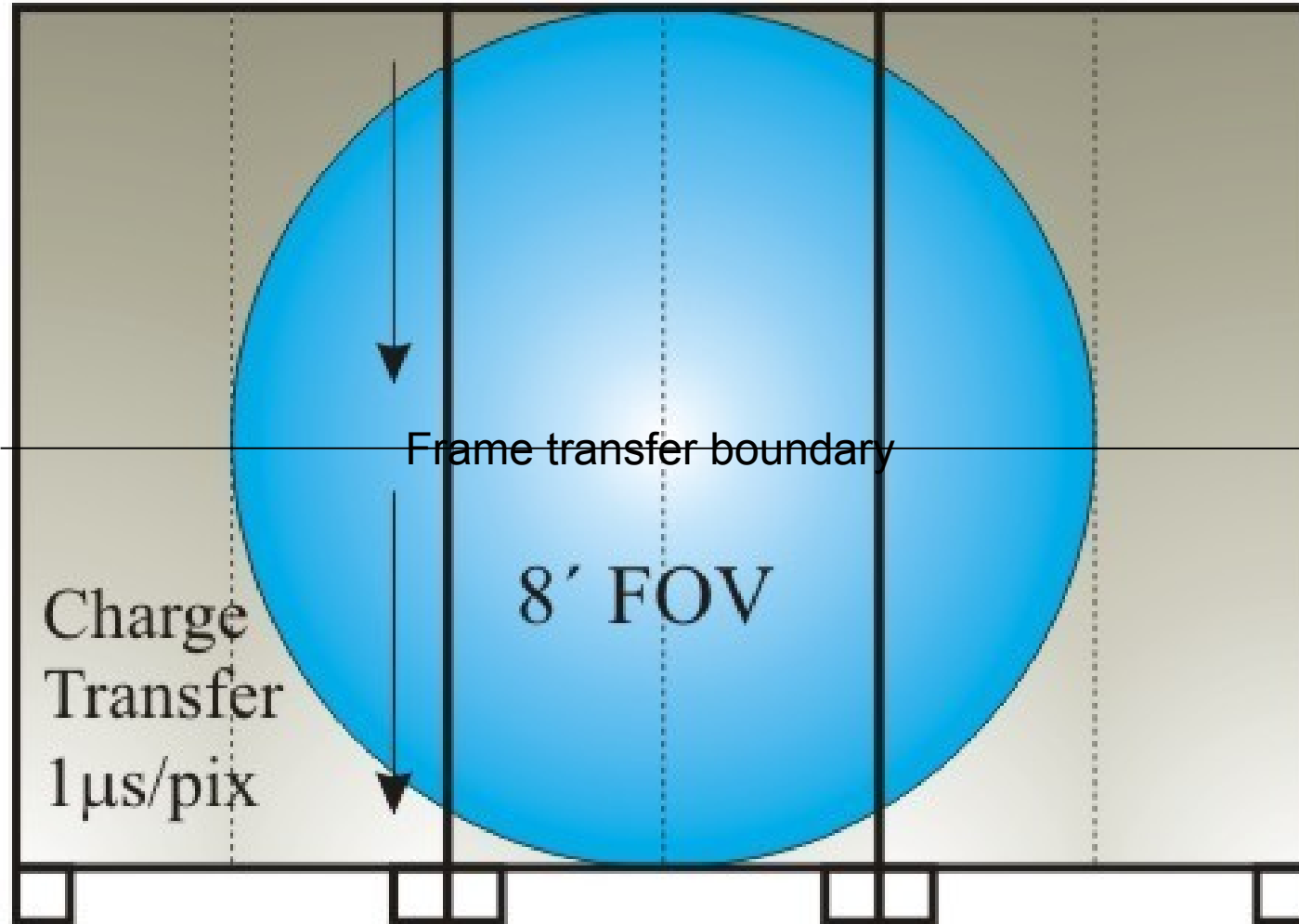
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



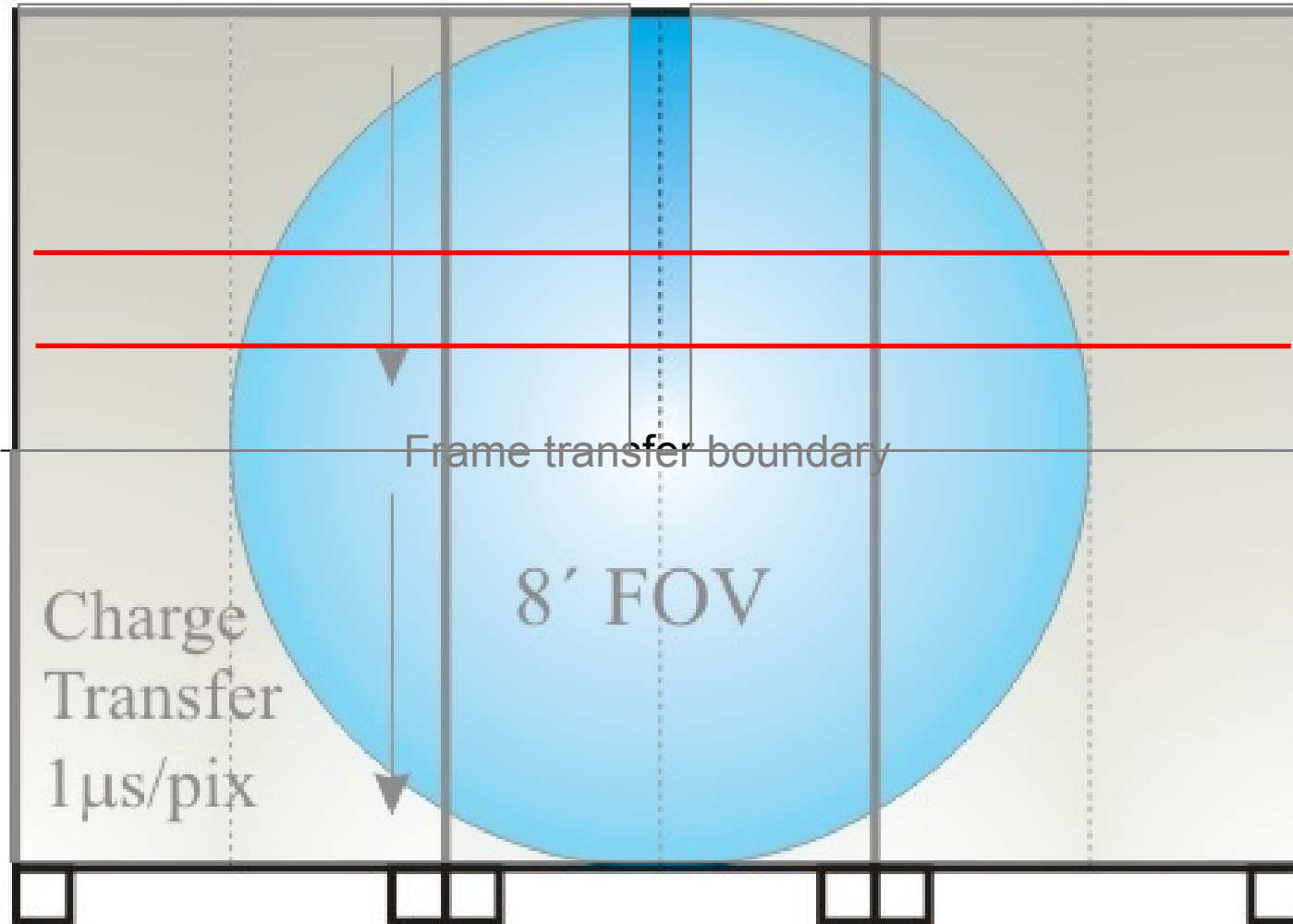
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Amplifier

Frame transfer long slit spectroscopy

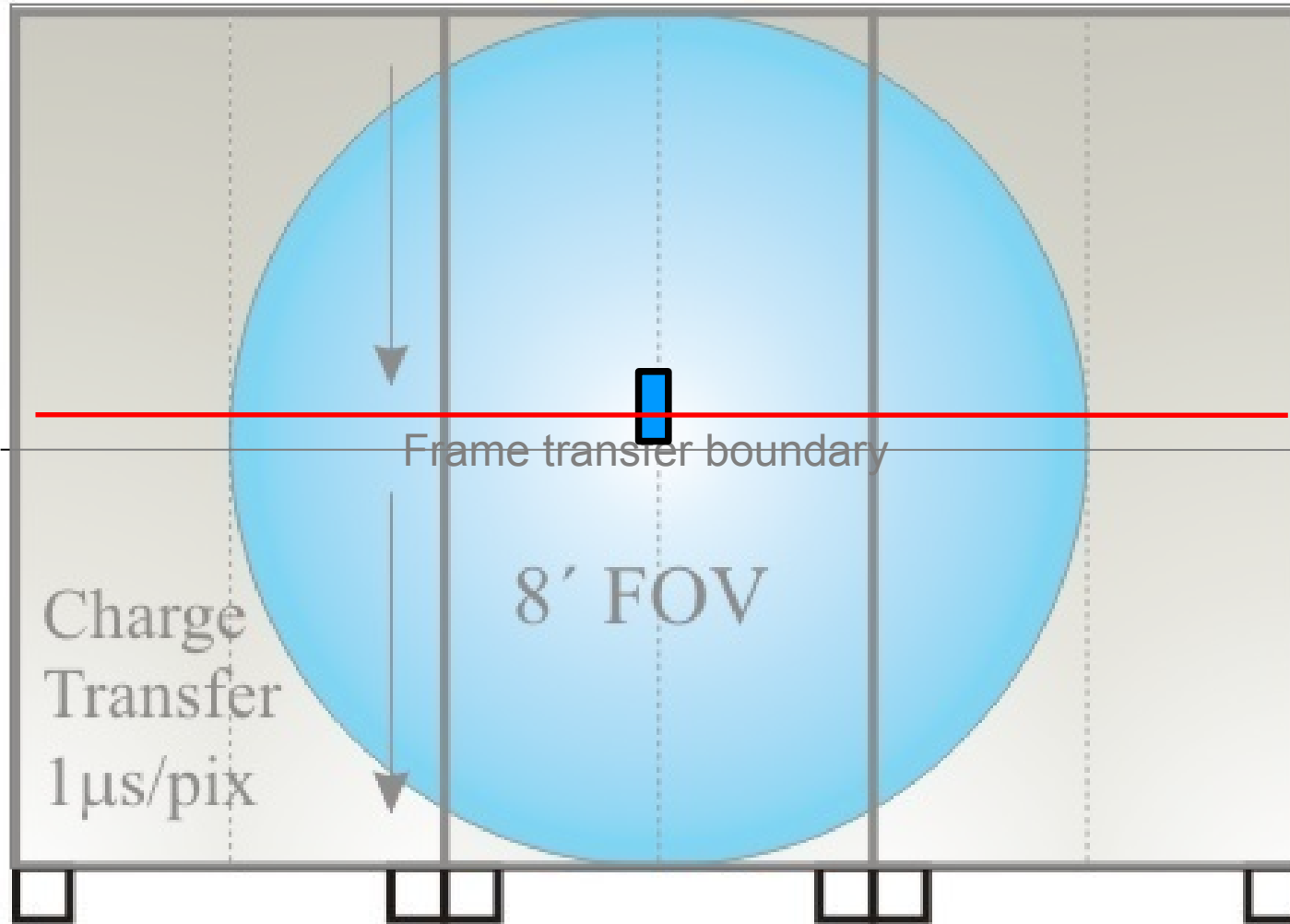
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Slot mode spectroscopy

Amplifier

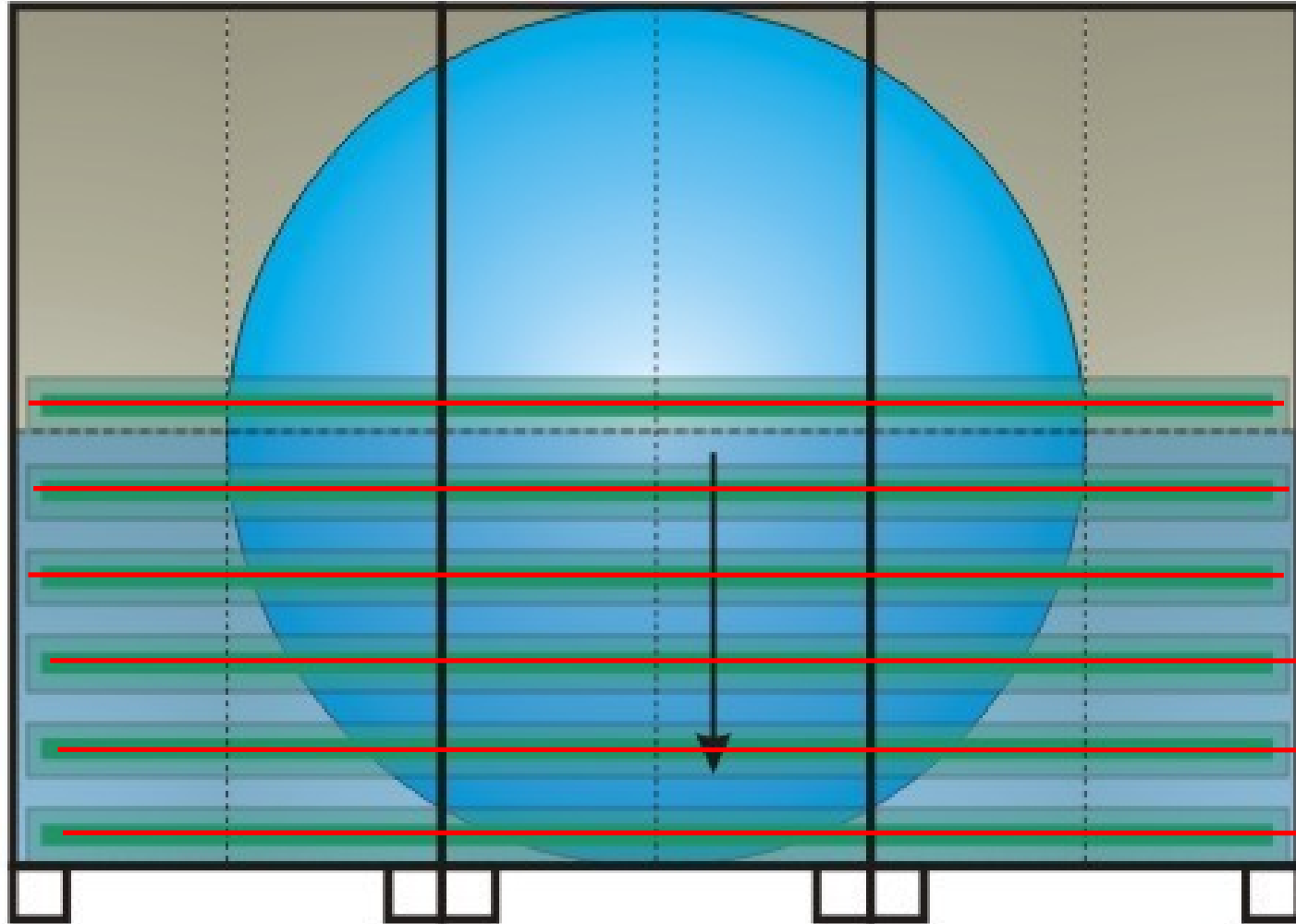
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Slot mode spectroscopy

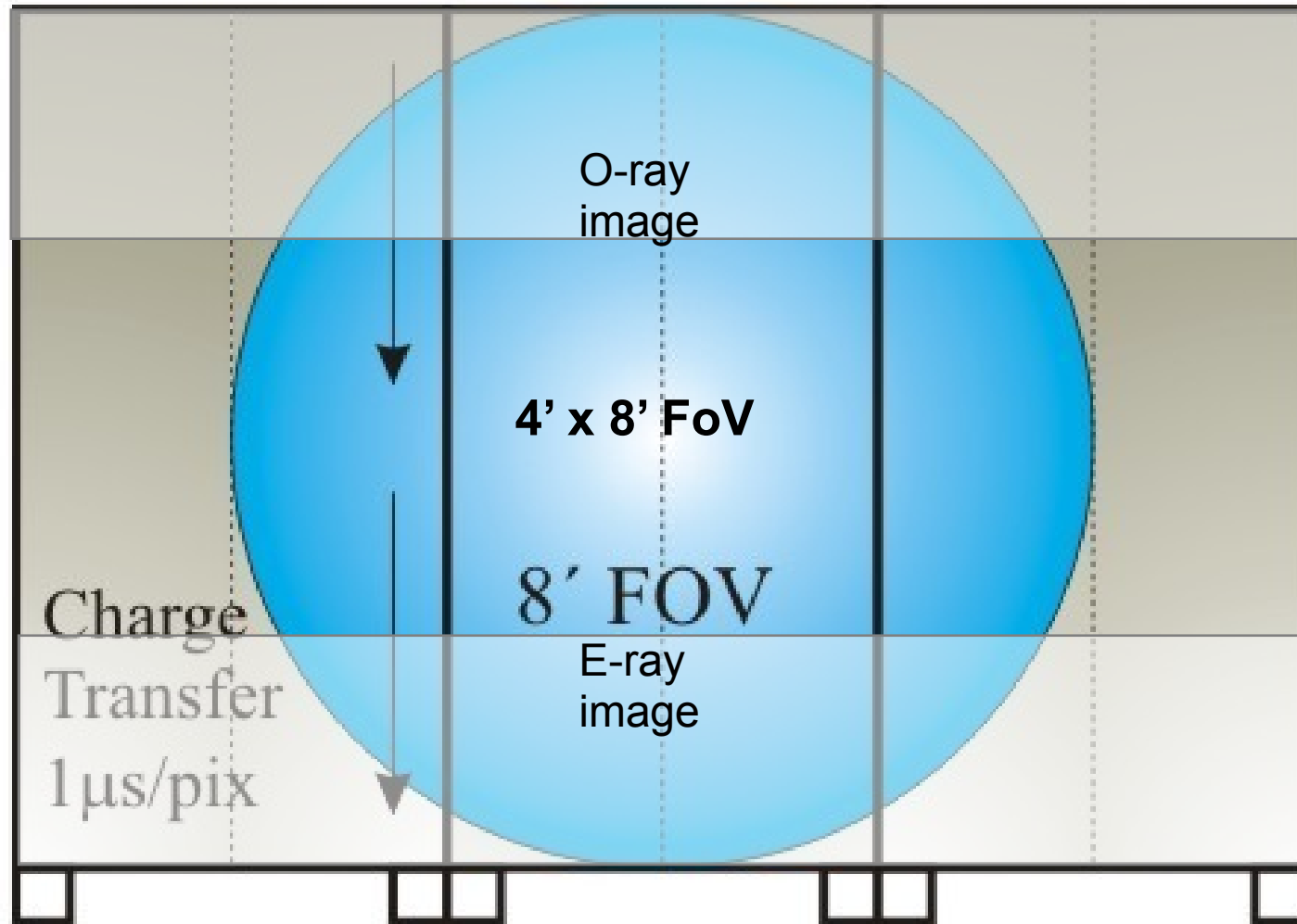
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



Imaging polarimetry
Beamsplitter has slight dispersion

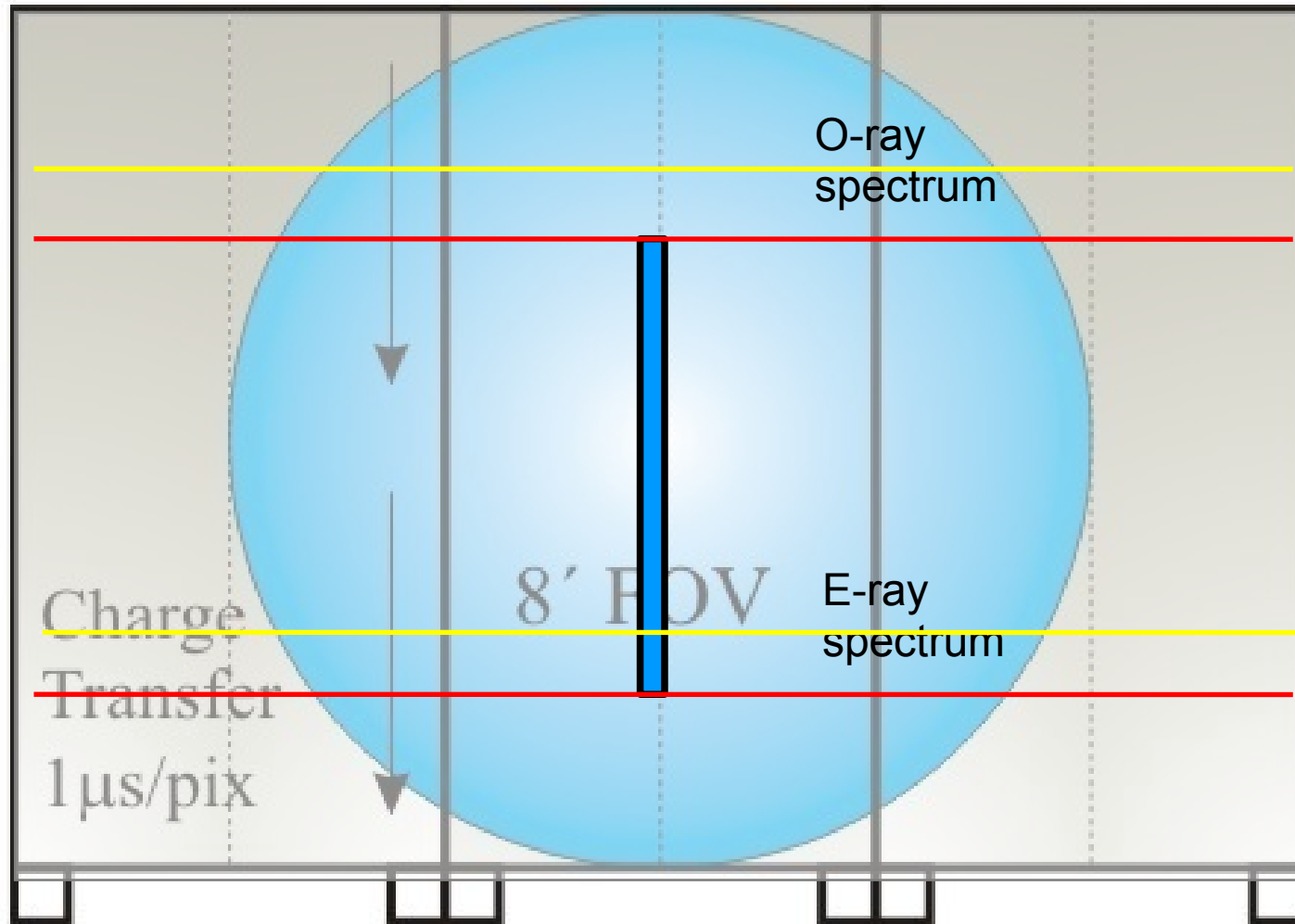
Robert Stobie Spectrograph (RSS)

2048

2048

2048

4098 spatial rows = 8'



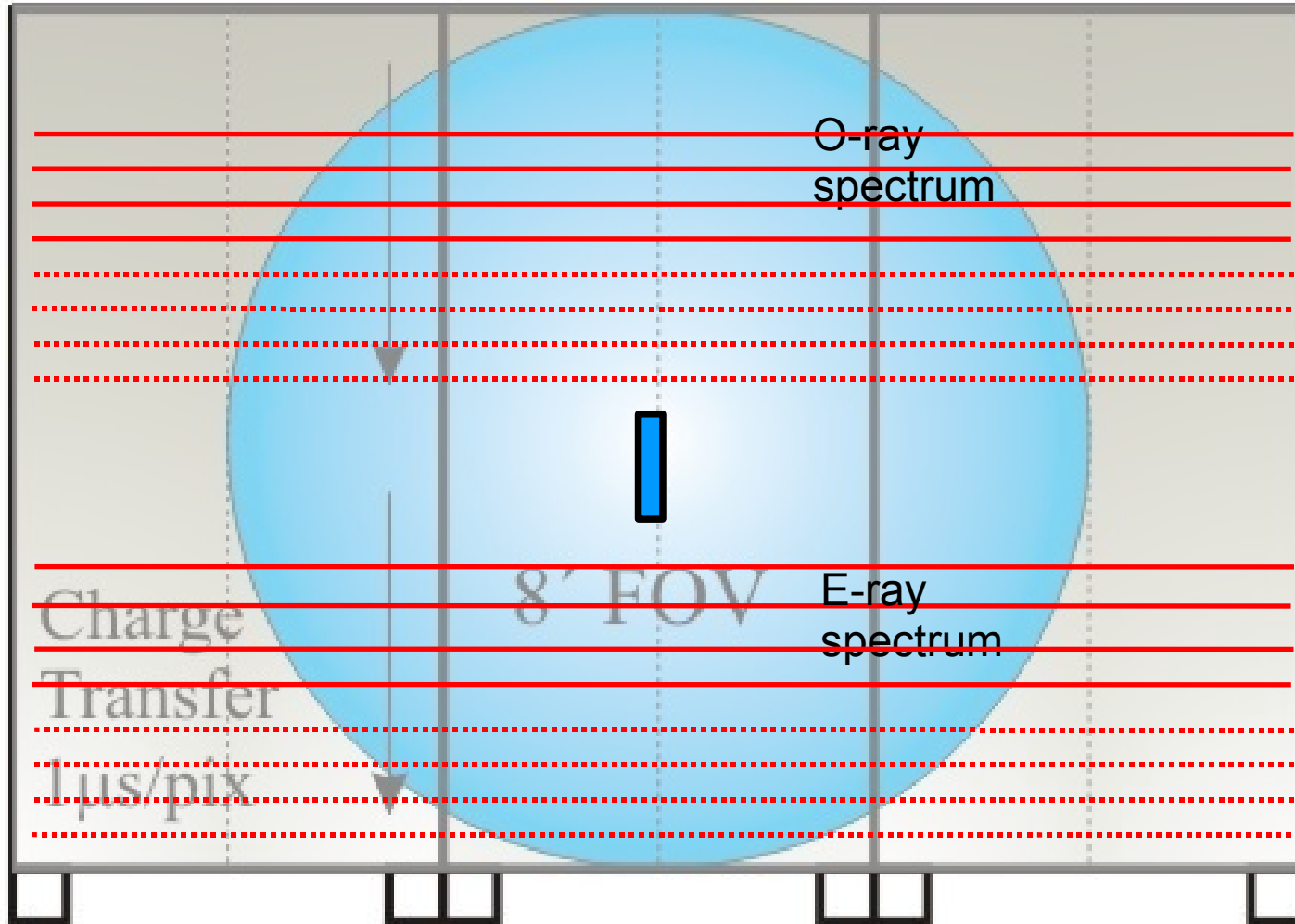
Amplifier

Spectro-polarimetry

Robert Stobie Spectrograph (RSS)

2048 2048 2048

4098 spatial rows = 8'



Rotate waveplates & Shuffle CCD rows down or backwards/ forwards

Waveplate angles

Spec	$\frac{1}{2}$	$\frac{1}{4}$
1	0	0
2	22.5	33.75
3	45	67.5
4	67.5	101.25
5	90	135
6	112.5	168.75
7	135	202.5
8	157.5	236.25

Charge shuffling Spectro-polarimetry

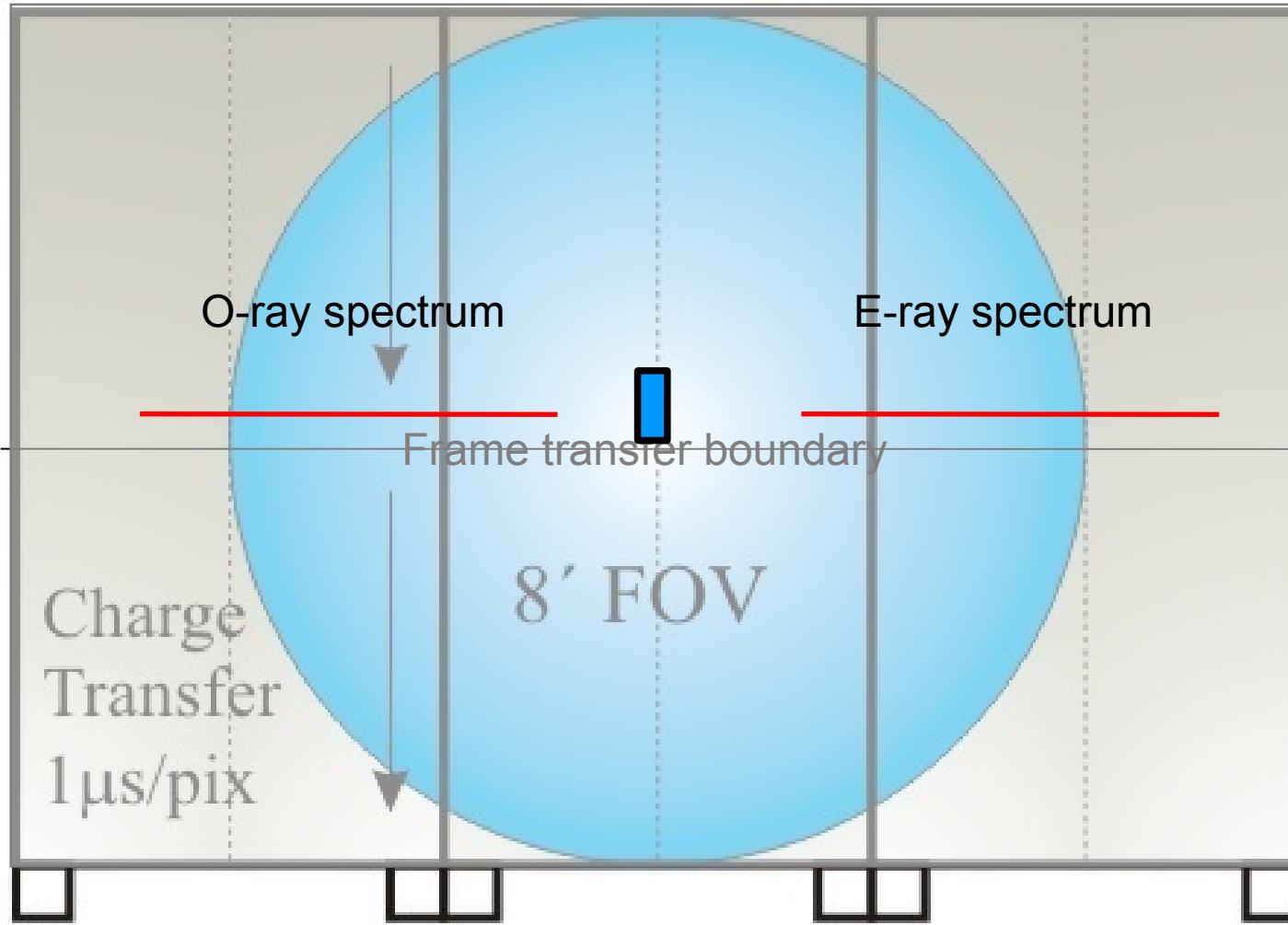
Robert Stobie Spectrograph (RSS)

2048

2048

2048

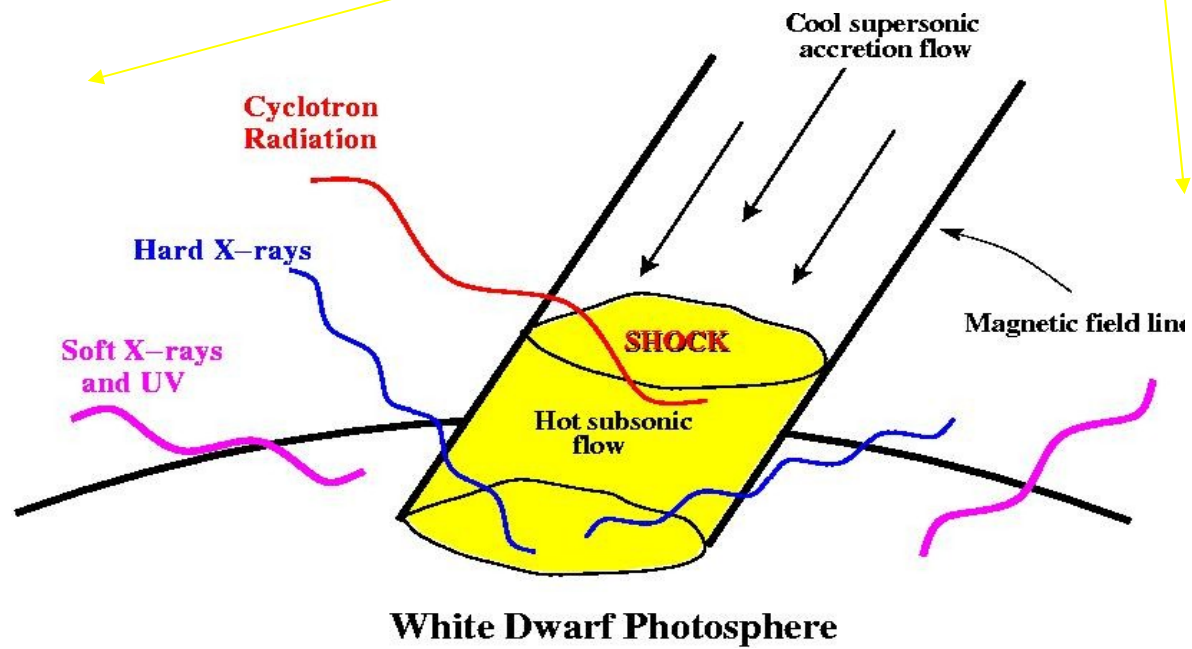
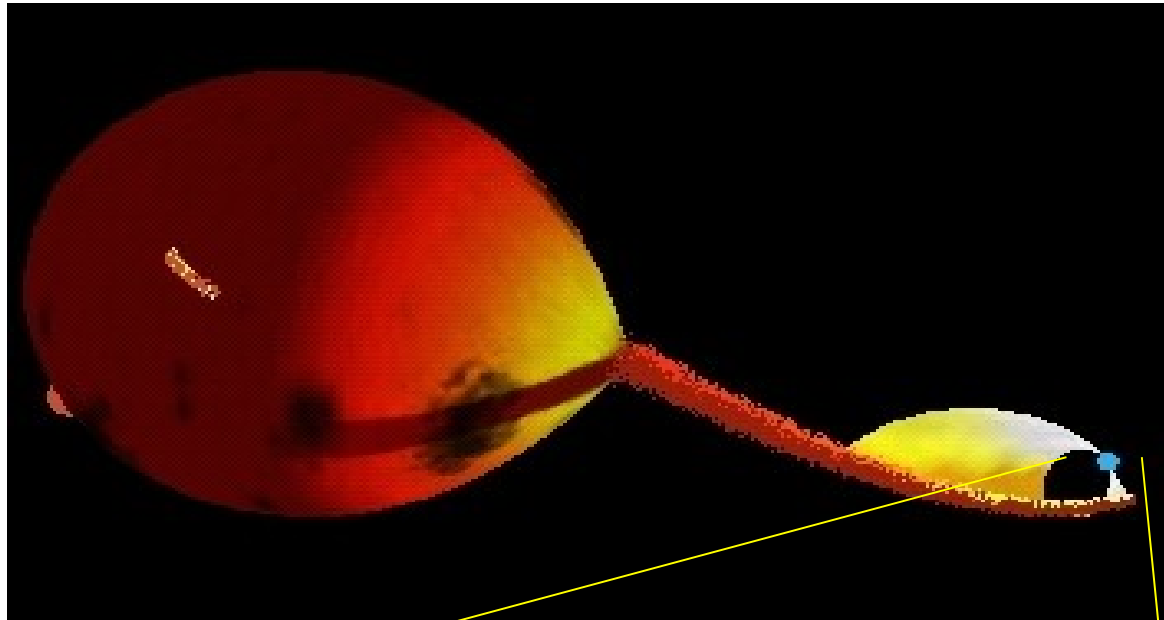
4098 spatial rows = 8'



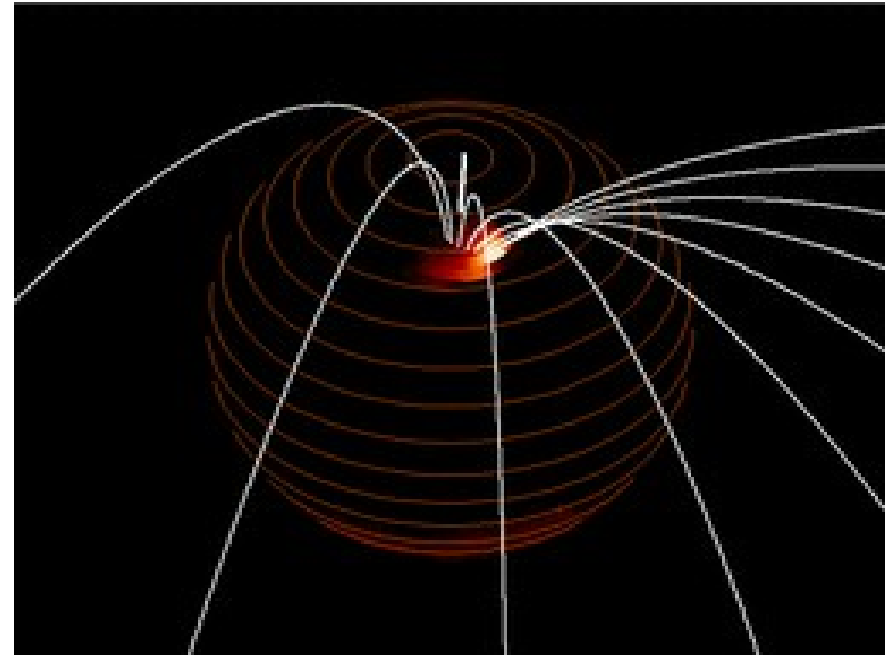
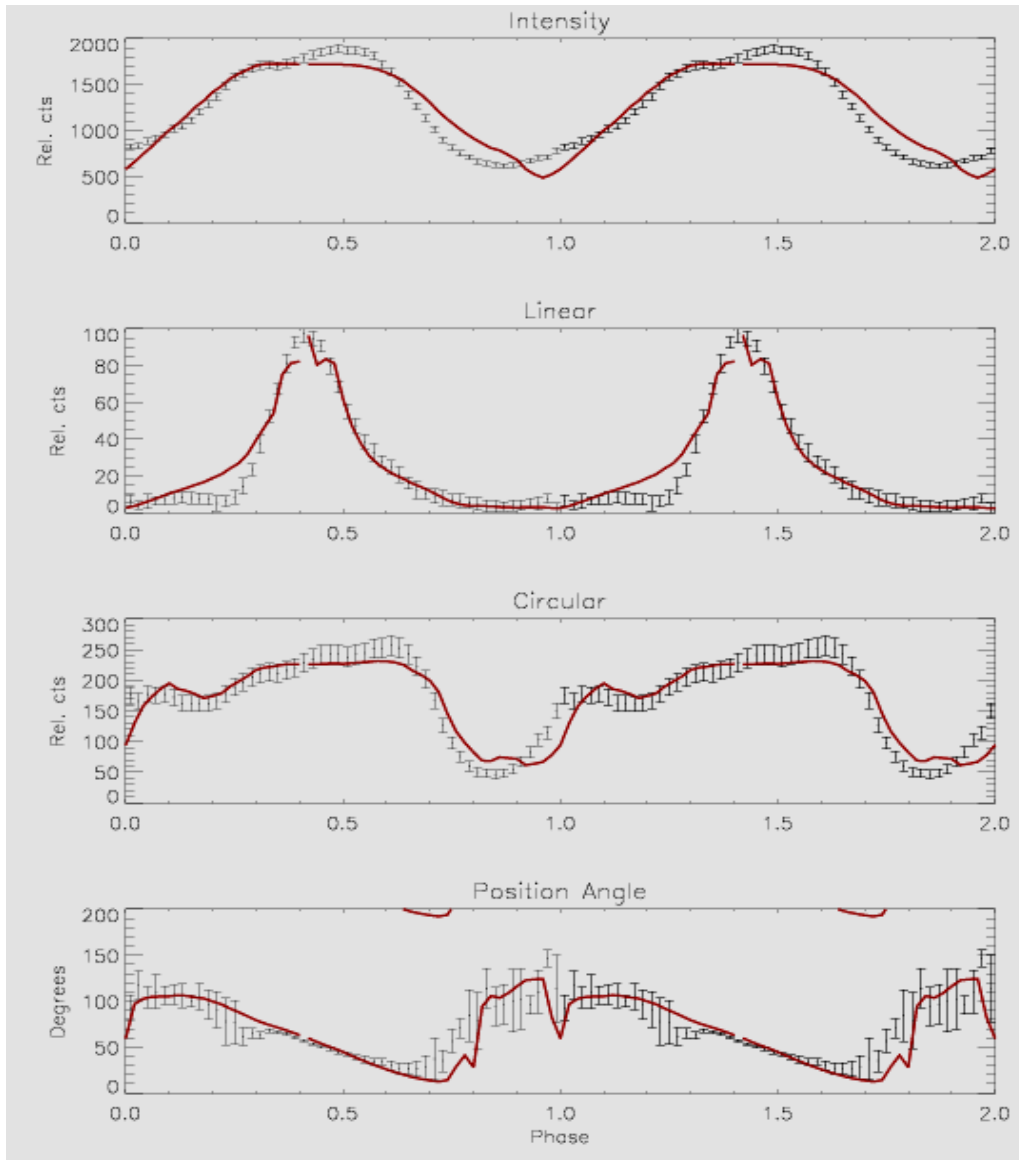
Amplifier

Rotated beamsplitter slot mode spectro-polarimetry

High Speed Polarimetry of Polars



High Speed Polarimetry of Polars



Stokes imaging technique (Potter):

Fitting model to data using a genetic algorithm.

Single pole-system with cyclotron emission region always in view

Spectro-Polarimetry of Polars

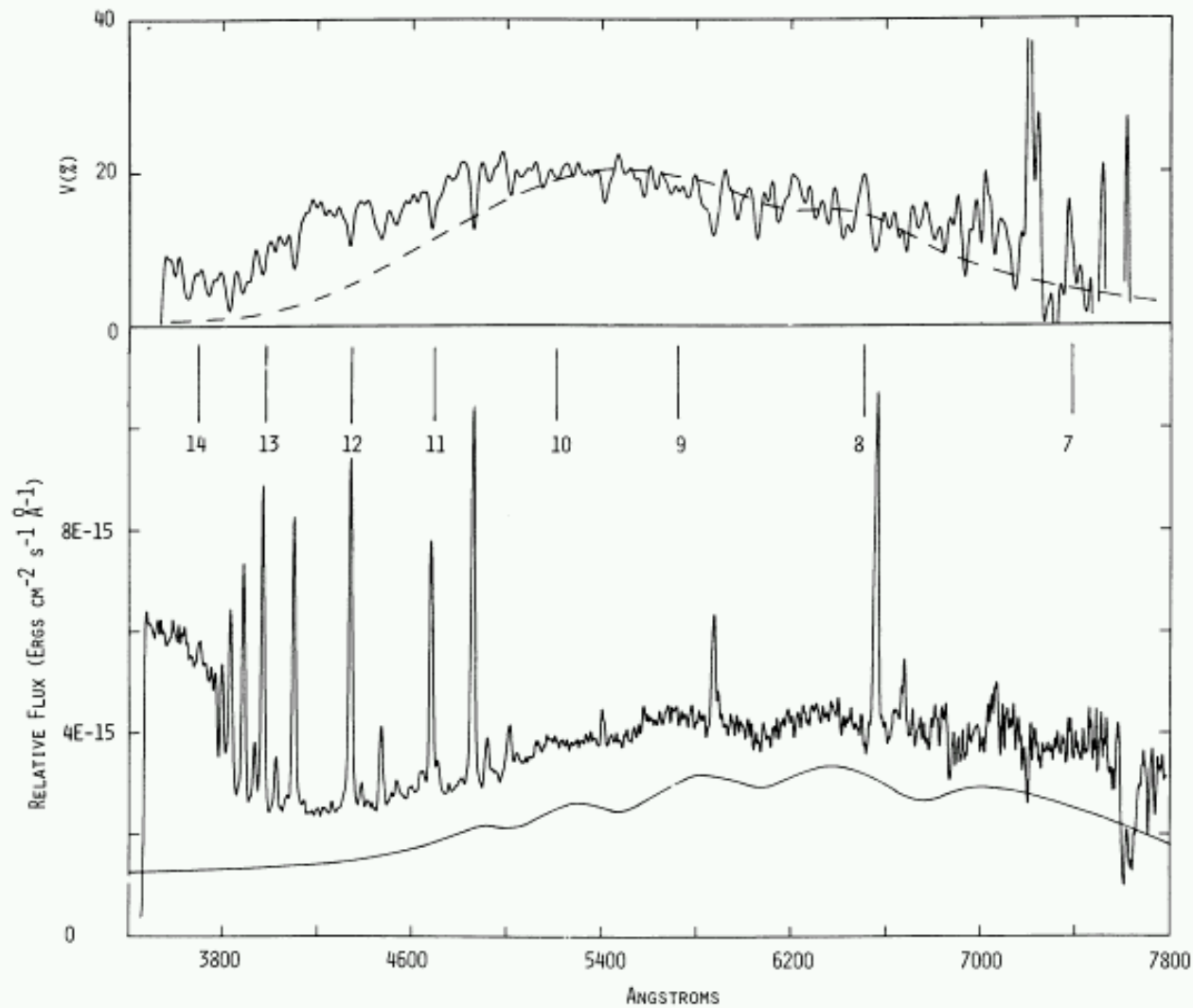


FIG. 4.—A constant Λ cyclotron model compared with observed intensity and polarization at $\phi = 0.5$ (bin 2). The model has parameters $B_p = 20.5$ MG, $T_e = 8$ keV, $N_e = 10^{16}$ cm $^{-3}$, $\Lambda = 10^7$, and $\theta = 85^\circ$. The theoretical intensity curve has been scaled to match the observed intensity at 6400 Å but is displaced vertically for clarity. Cyclotron harmonic numbers between 7–14 are indicated.

Spectro-Polarimetry of Polars

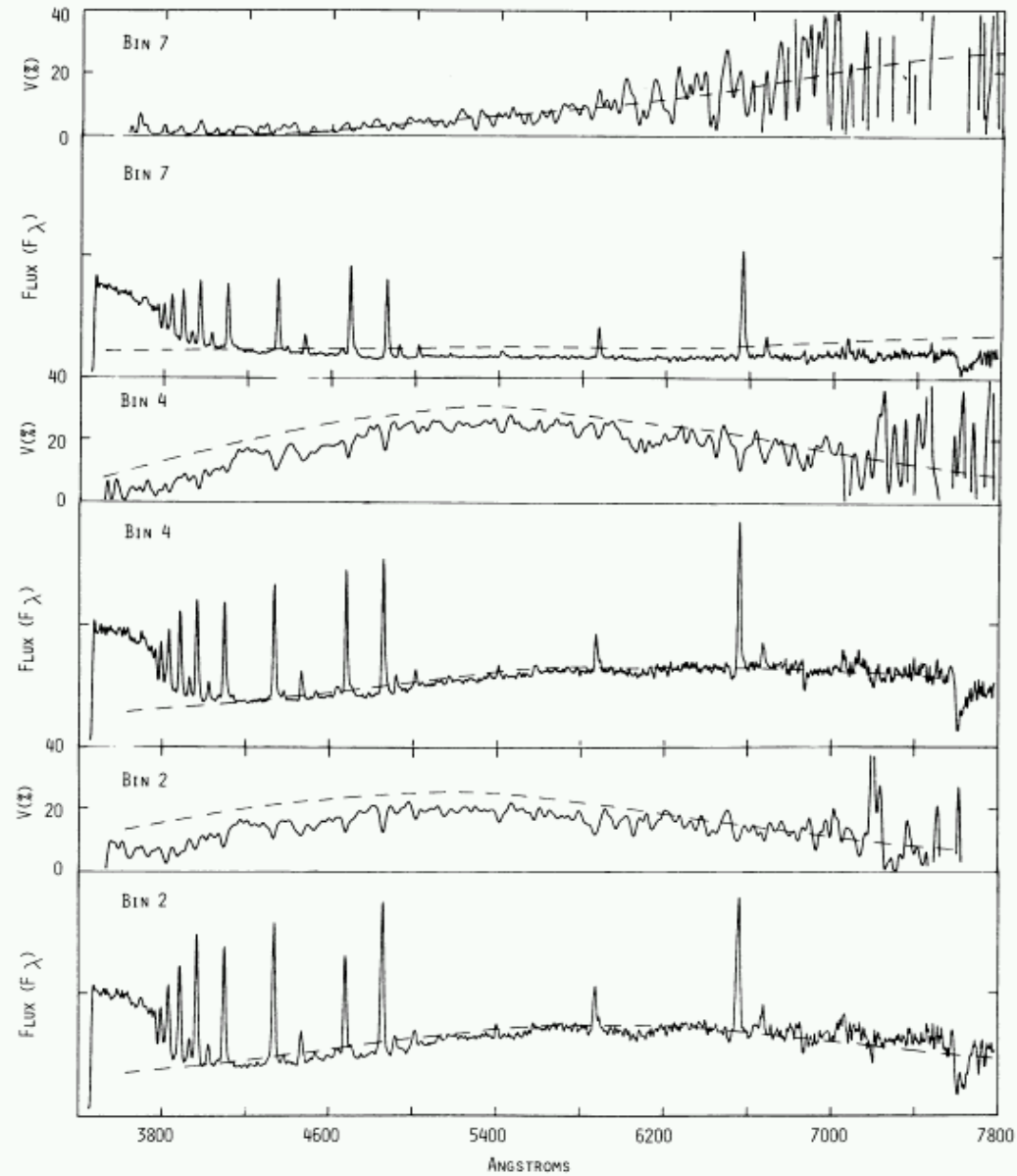
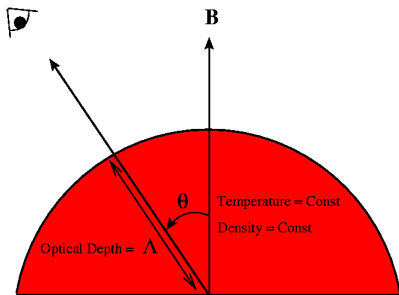
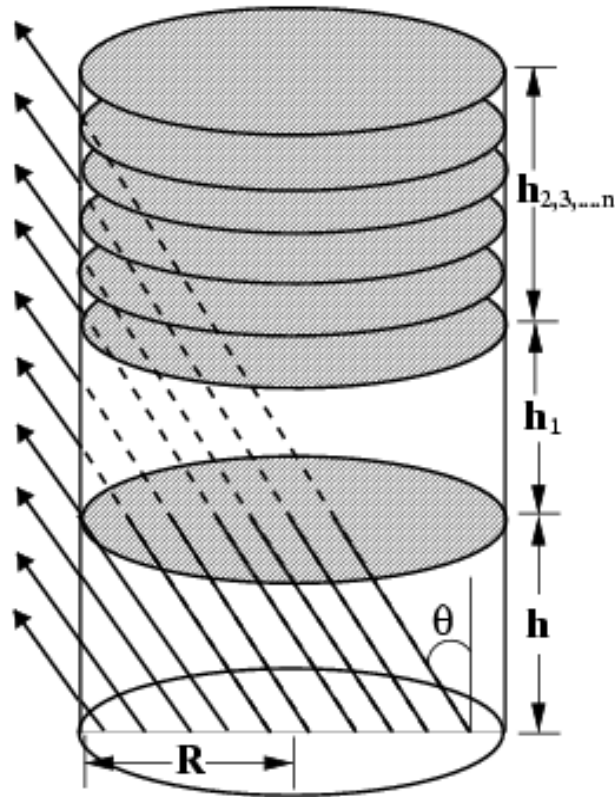
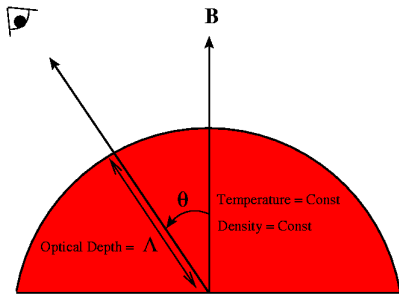
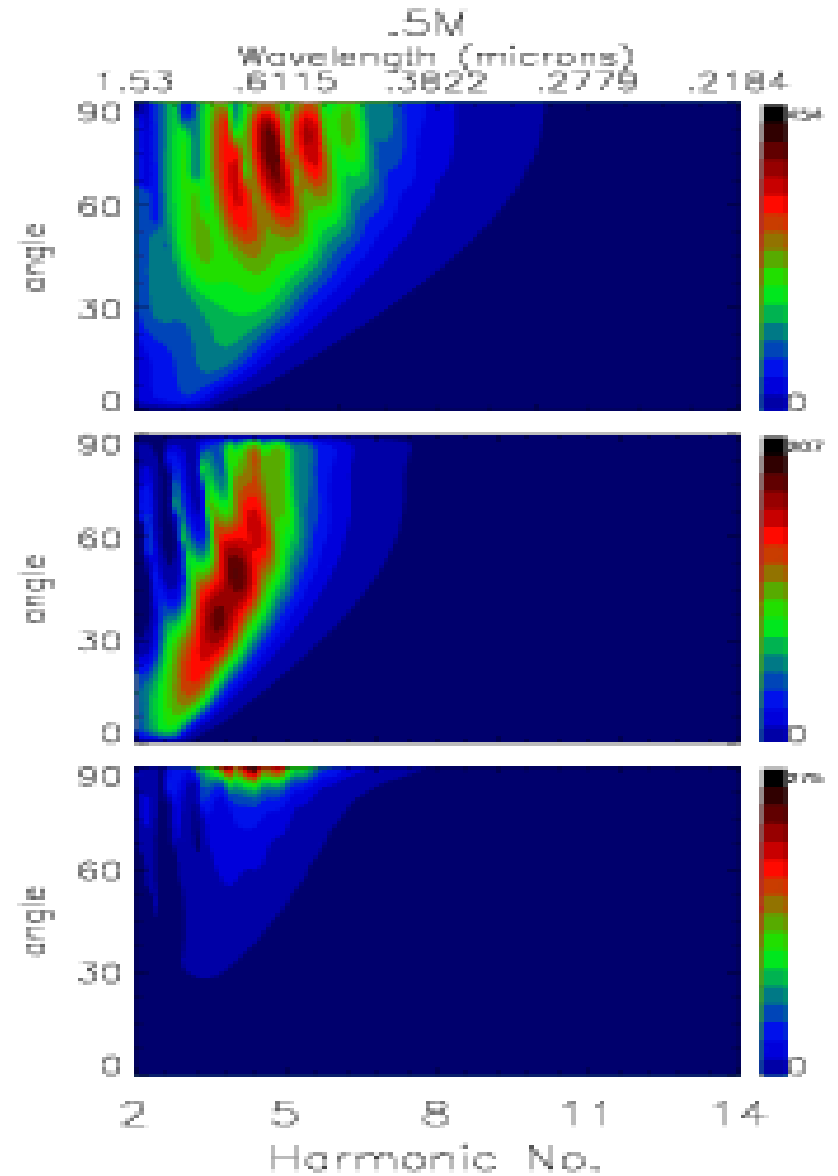


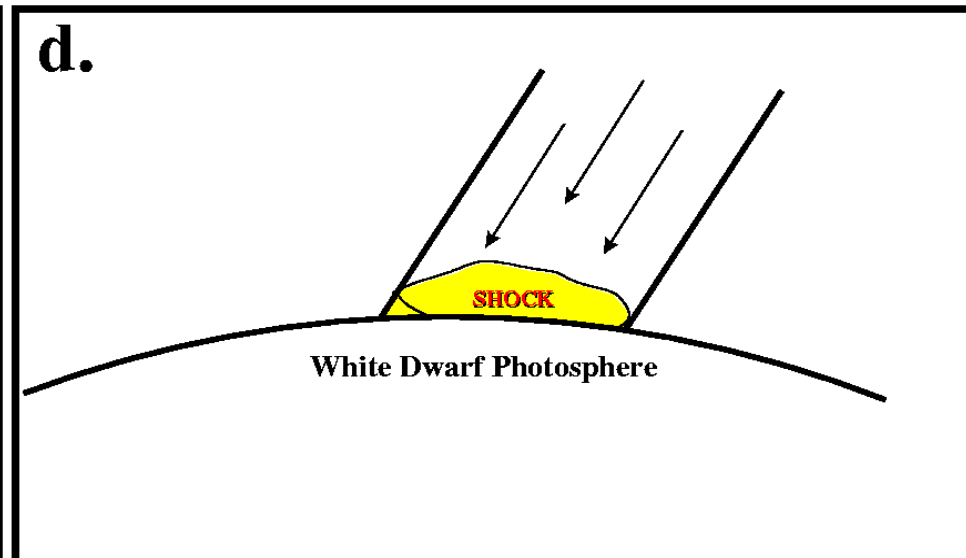
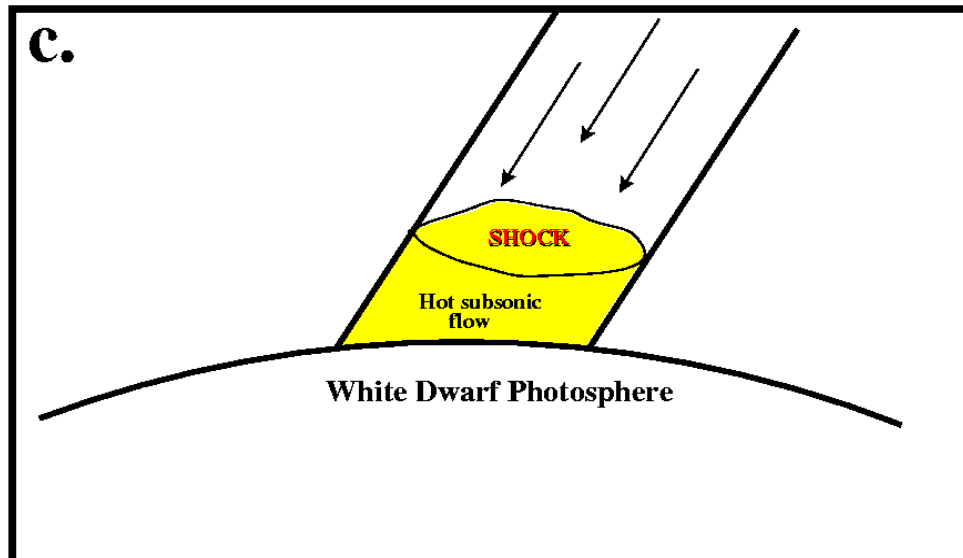
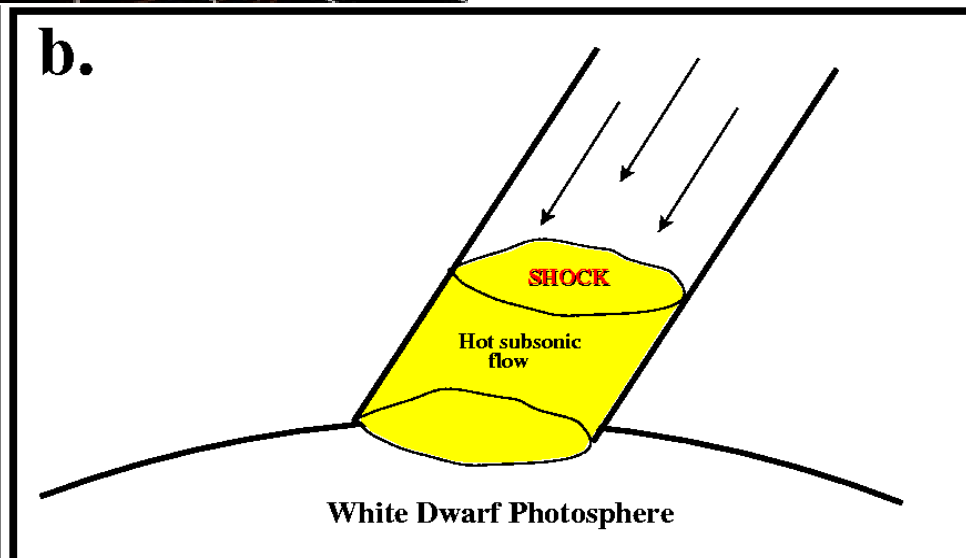
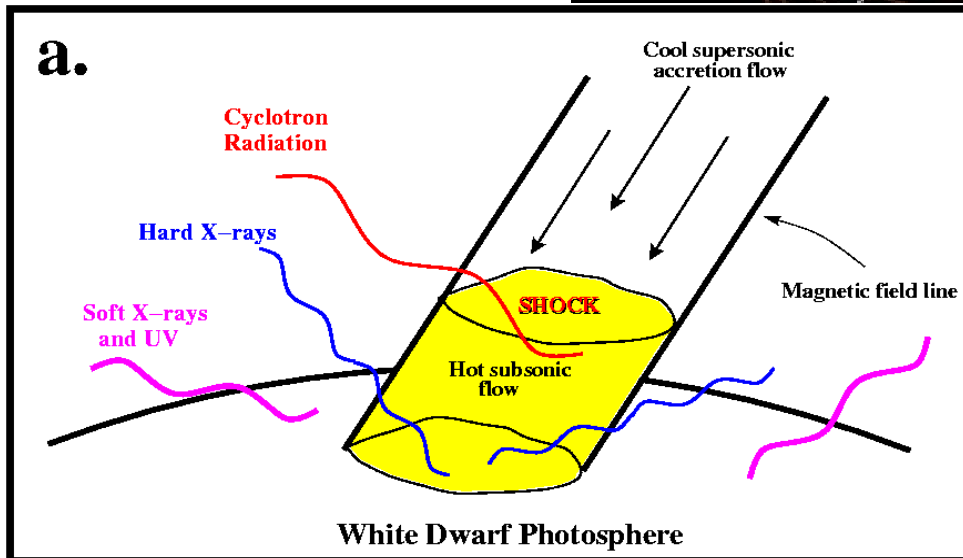
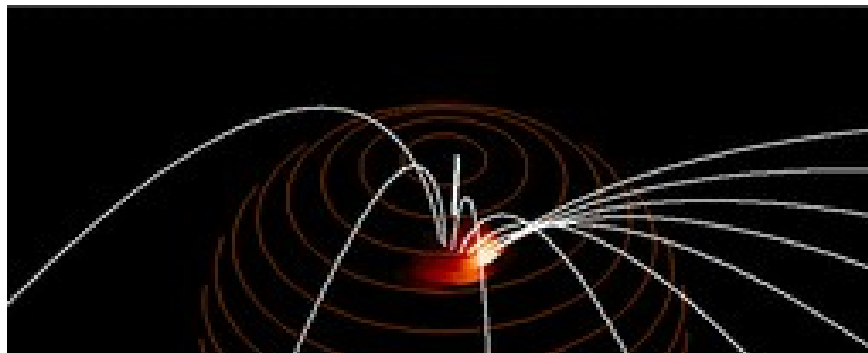
FIG. 7.—Cyclotron emission spectra (intensity and polarization) are compared with observations at three different phases corresponding to $\phi = 0.0$ (bin 7), $\phi = 0.7$ (bin 4), and $\phi = 0.5$ (bin 2). Model parameters are $B_p = 25$ MG, $T_e = 20$ keV, $N_e = 10^{16}$ cm $^{-3}$, $\Lambda_h = \Lambda_d = 4.5 \times 10^4$.

Stratified accretion shocks



Allow testing of more realistic shock models with stratified temperature and density profiles dependent on parameters like:
 White dwarf mass, accretion rate
 magnetic field strength..





The Other Telescopes

Telescopes include:

1.9m, 1.0m, 0.75m 0.5m, IRSF 1.4m, MONET 1.2m

Instruments include

UCTCCD

CASS Spectrograph

HIPPO

SIRIUS

Why?

Simultaneous: absolute photometry for SALTICAM and RSS
photo-polarimetry

Monitoring: Outbursts
High/low states
Novae etc