

# WISE survey of novae

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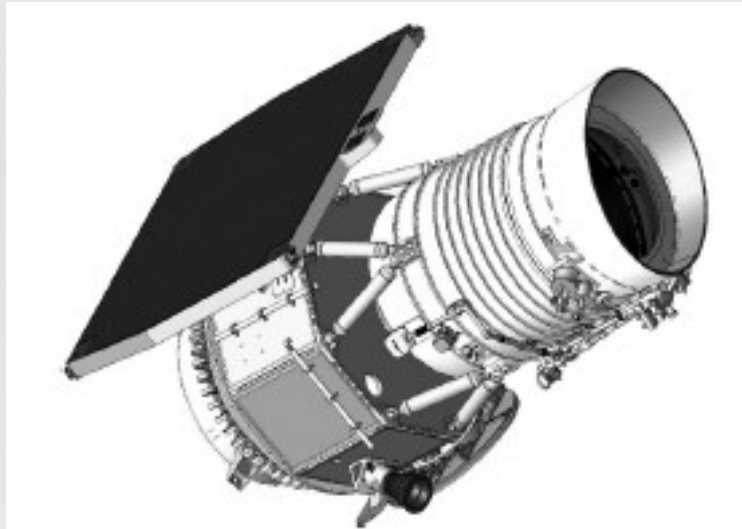
Andrew Helton (SOFIA/USRA)

# WISE survey of novae

- ▶ **Widefield Infrared Survey Explorer (WISE)**
- ▶ **All Sky Survey**
- ▶ **3.4, 4.6, 12, 22 microns**
- ▶ **2010 January 14 – 2010 November**
- ▶ **0.08, 0.11, 1, 6 mJy sensitivity (5sigma)**
- ▶ **WISE PSFs: 6.1" (Band 1), 6.4" (2), 6.5" (3) 12" (4)**
- ▶ **Astrometric precision for high S:N sources <0.15"**

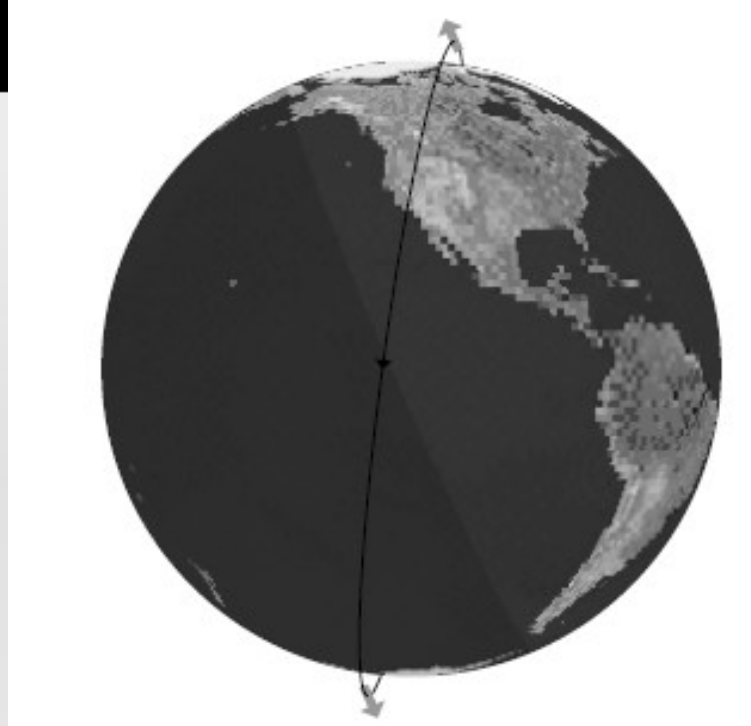
**See Wright et al., 2010, AJ, 140, 1868 for details of WISE mission**

# WISE survey of novae



- ▶ 0.4m telescope, solid H cooled
- ▶ Solar elongation @ 90degrees
- ▶ Scan N ecliptic pole – S ecliptic pole

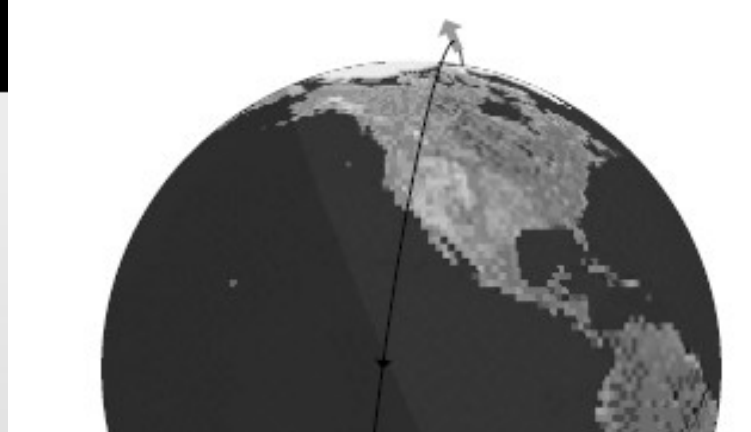
# WISE survey of novae



WISE pointing and orbit at June solstice

WISE points perpendicular to Earth-Sun line

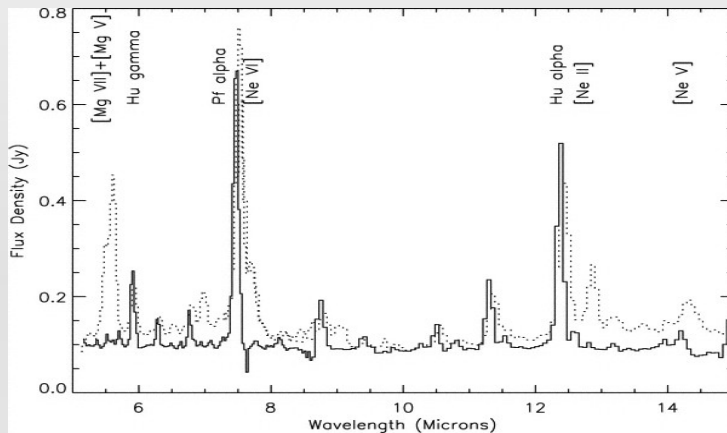
# WISE survey of novae



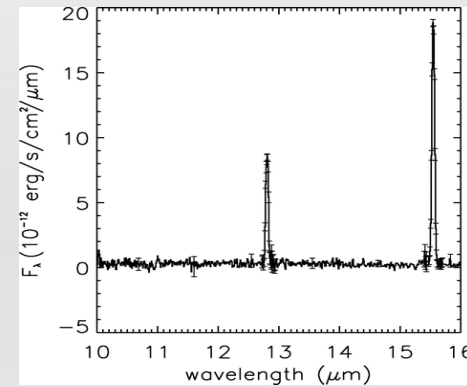
**Survey coverage for 1 orbit, 2 successive orbits,  
two orbits separated by 20 days**

# WISE survey of novae

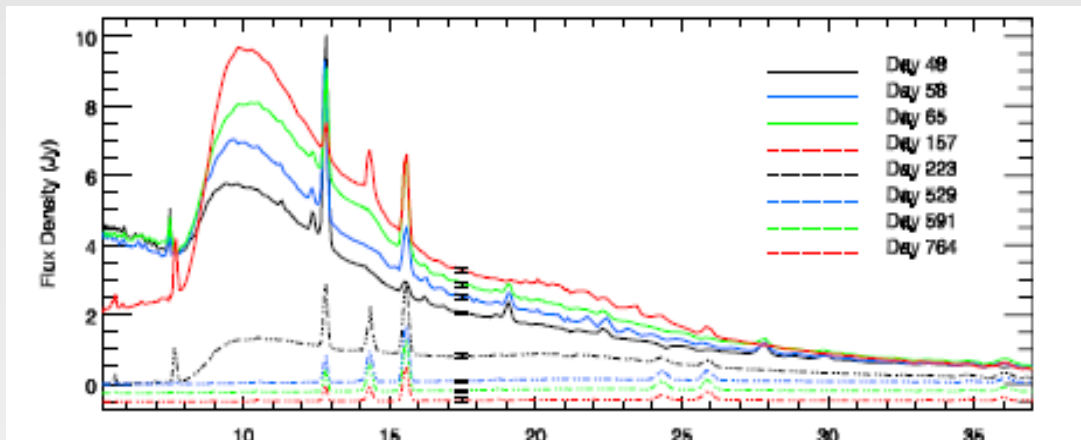
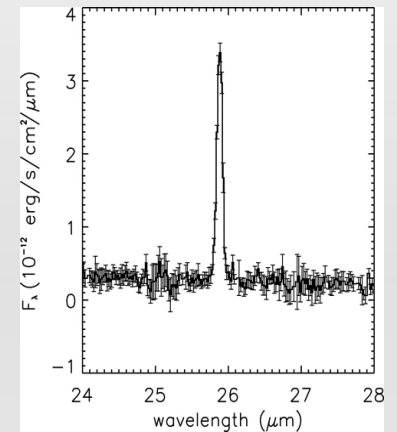
## Spitzer spectra of dusty and emission line novae



V1186/7 Sco (Schwarz et al. 2007)

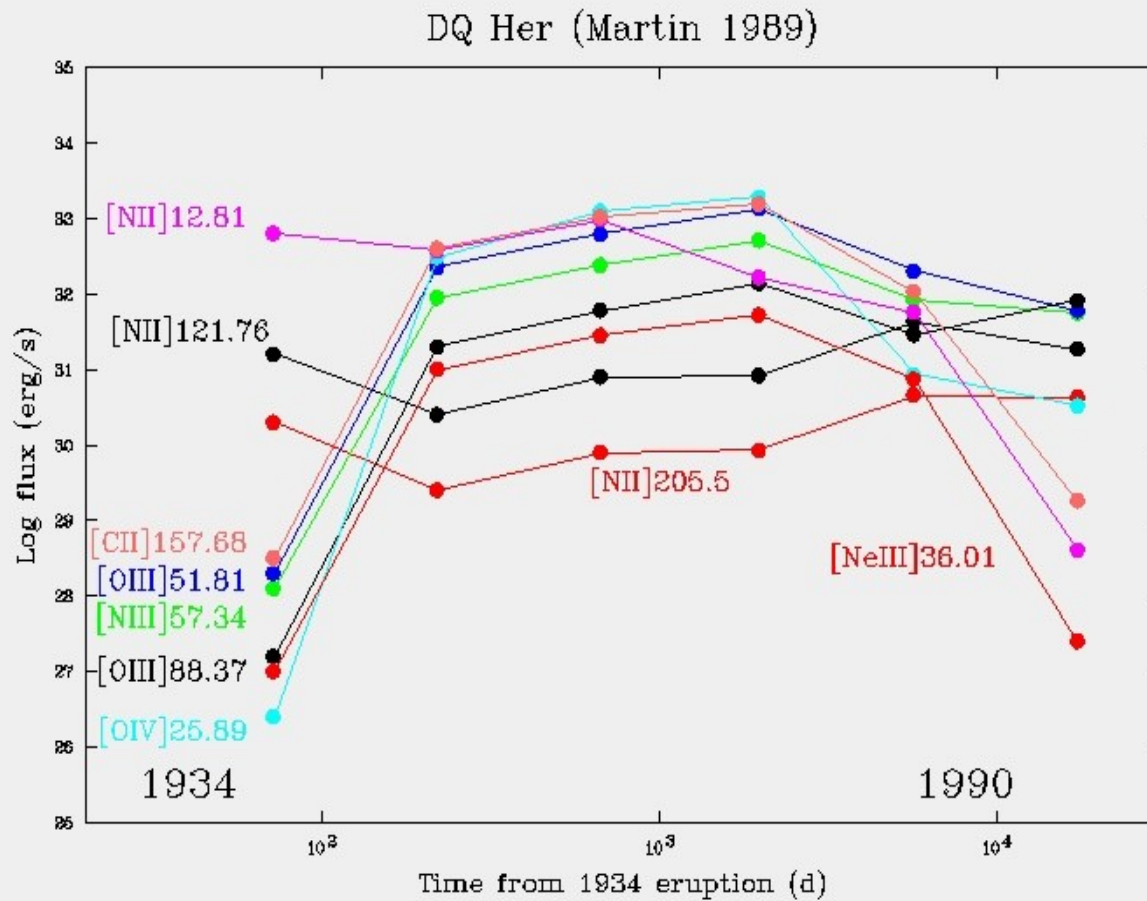


QU Vul – a “neon nova” (Gehrz et al. 2008)



Dusty V1065 Cen (2007)  
(Helton et al., 2010, AJ, 140, 1347)

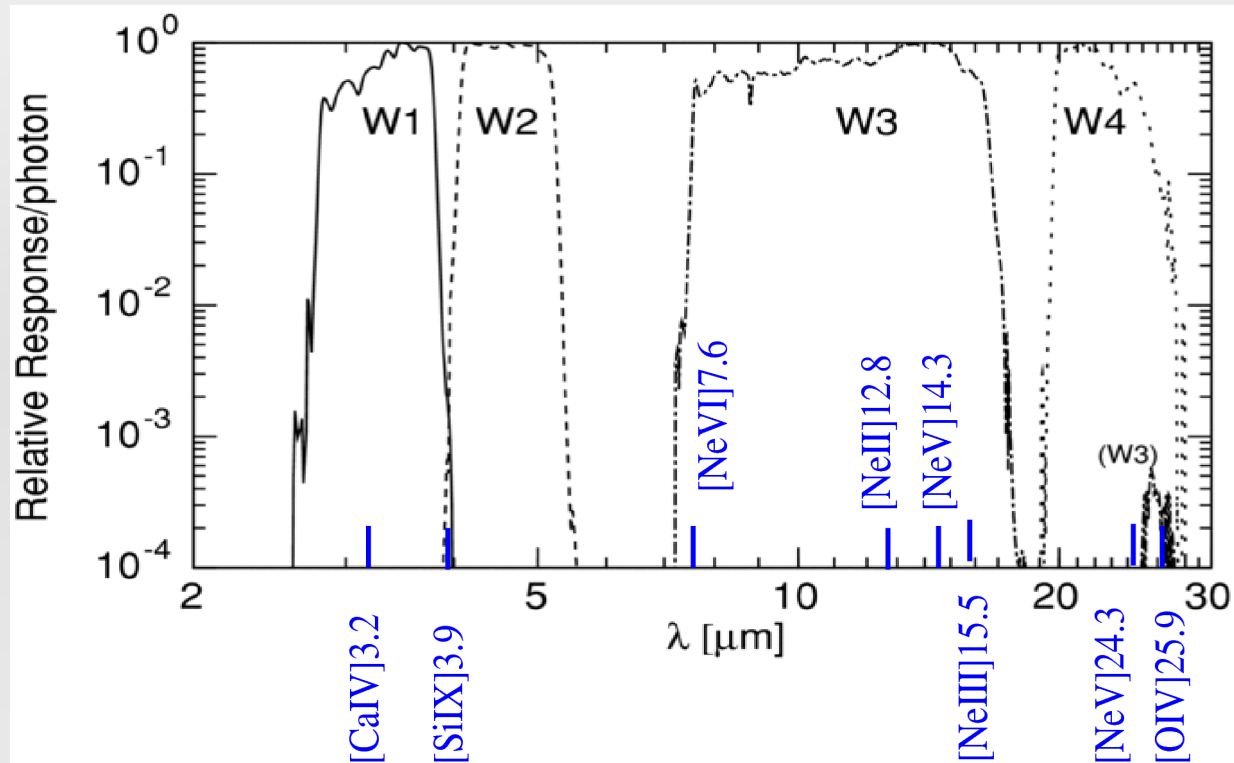
# WISE survey of novae



Schematic of long-term variation of emission line fluxes.

Based on CLOUDY modelling of DQ Her by Martin (CNI 1989)

# WISE survey of novae



WISE filter responses

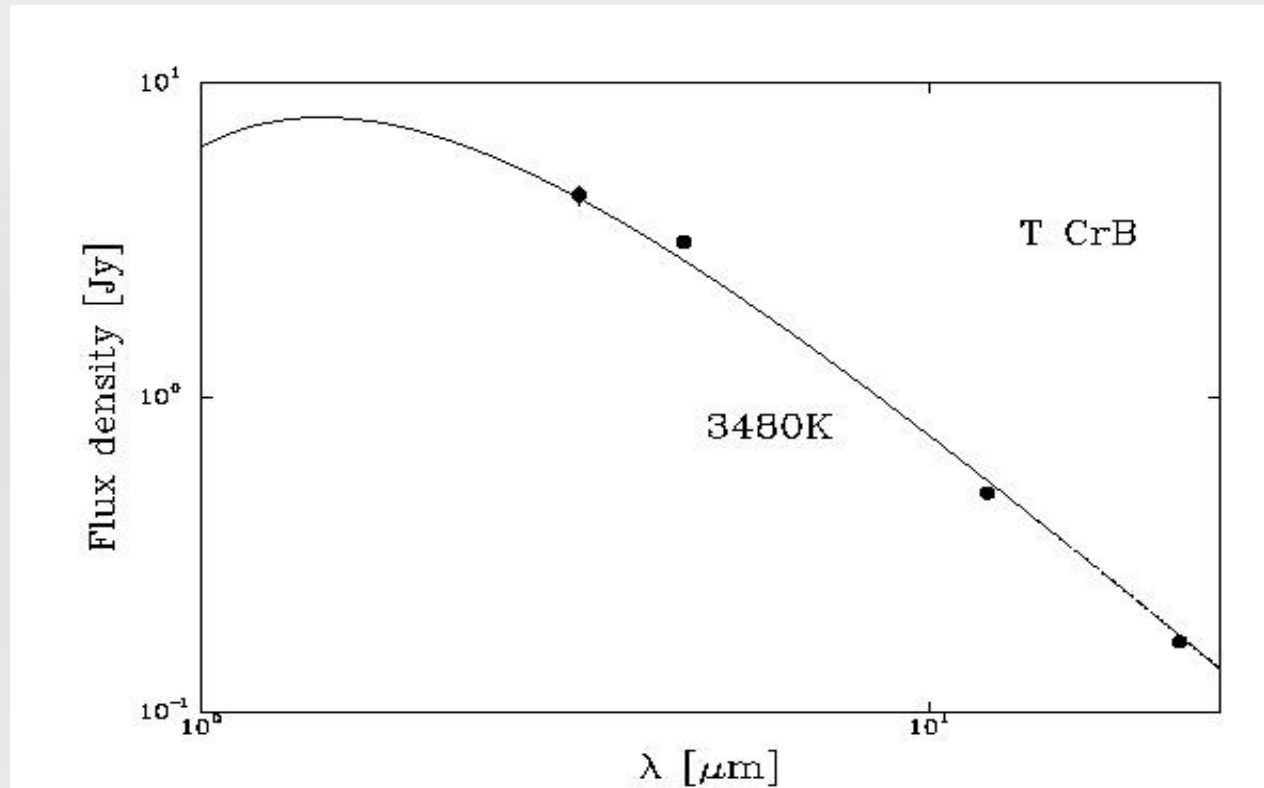
Wavelengths of prominent nova lines indicated



# WISE survey of novae

- ▶ WISE catalogue searched [so far] for 64 CNe and RNe
- ▶ 22 non-detections/spatial blends
- ▶ 38 with  $>\sim 4$  sigma detections in at least one band
- ▶ 14 with  $>\sim 4$  sigma detections in all 4 bands
- ▶ 17 post-WISE CNe/RNe (so we see progenitors) -
  - 4 with detections in at least one band
  - 3 with detections in all 4 bands

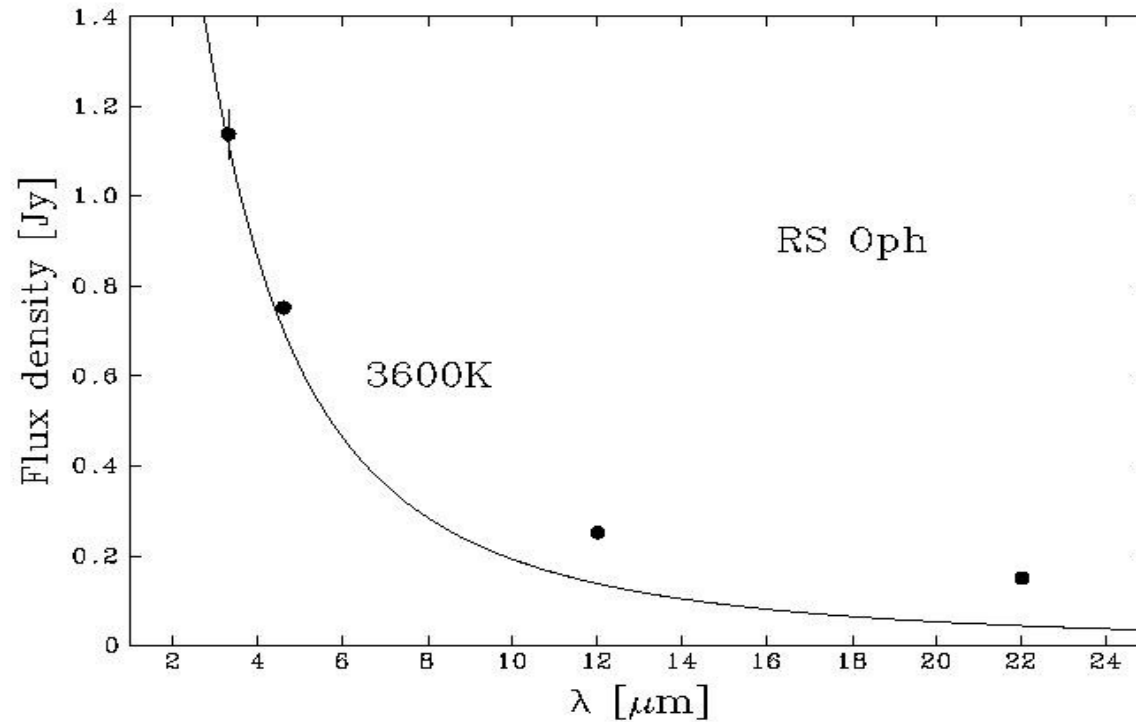
# WISE survey of novae



**SEDs of RNe**

**T CrB – showing RG only**

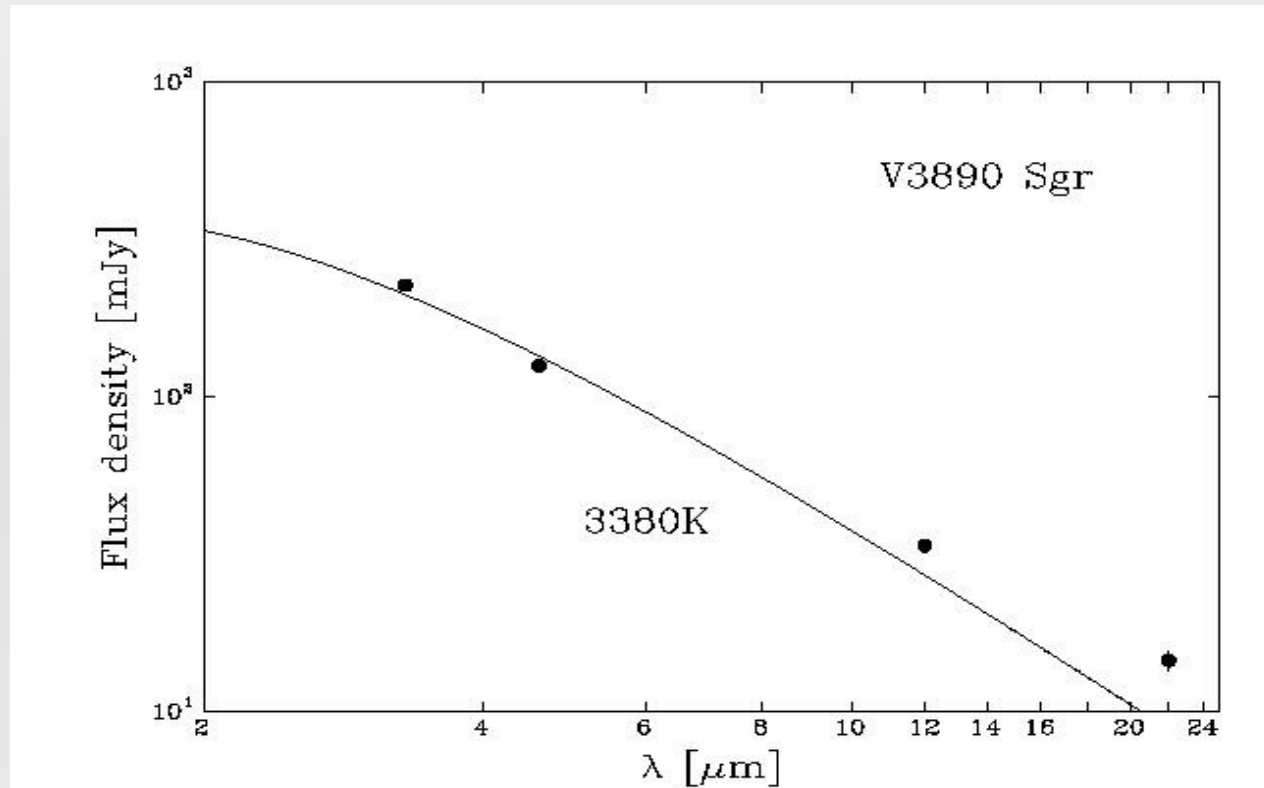
# WISE survey of novae



## SEDs of RNe

RS Oph – showing evidence of dust  
(Evans et al. ApJ, 671, L157,  
Rushton et al., MNRAS, 401, 99)

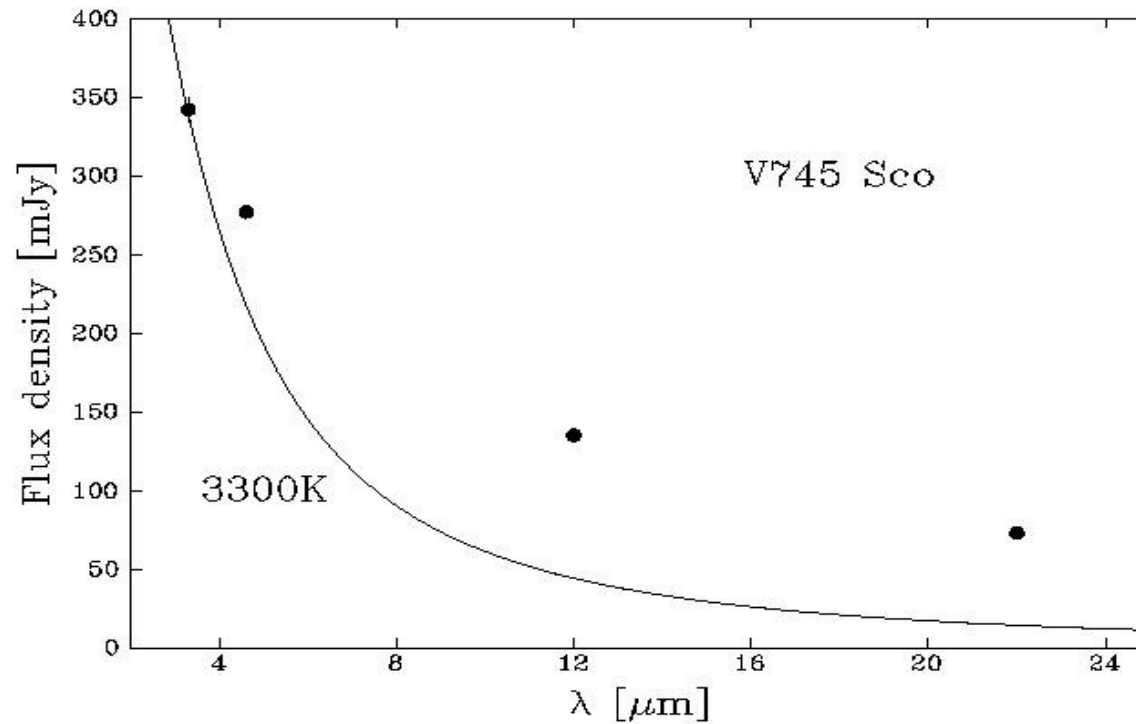
# WISE survey of novae



## SEDs of RNe

V3890 Sgr – showing evidence of an excess at 12 and 22 microns

# WISE survey of novae

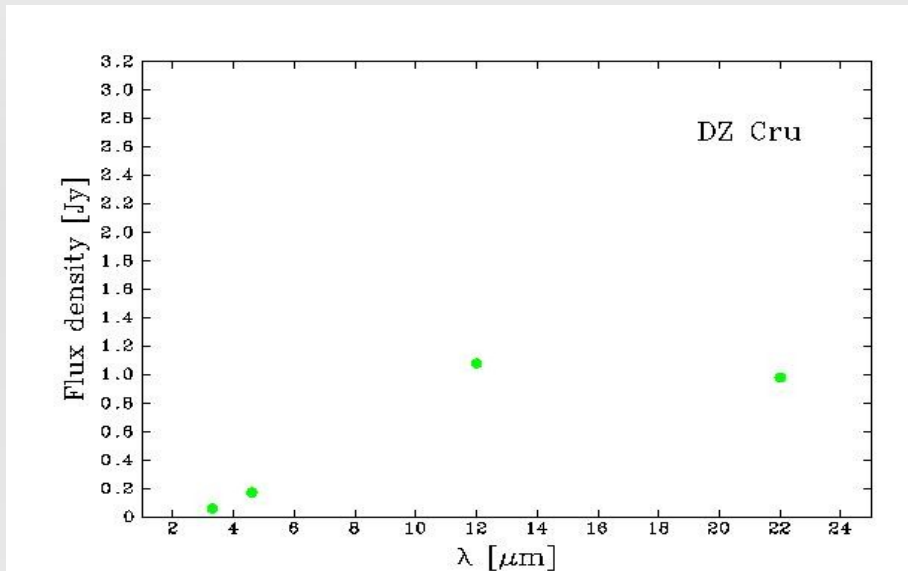


**SEDs of RNe**

**V745 Sco – showing evidence for an excess at 4.6, 12 and 22 microns**

# WISE survey of novae

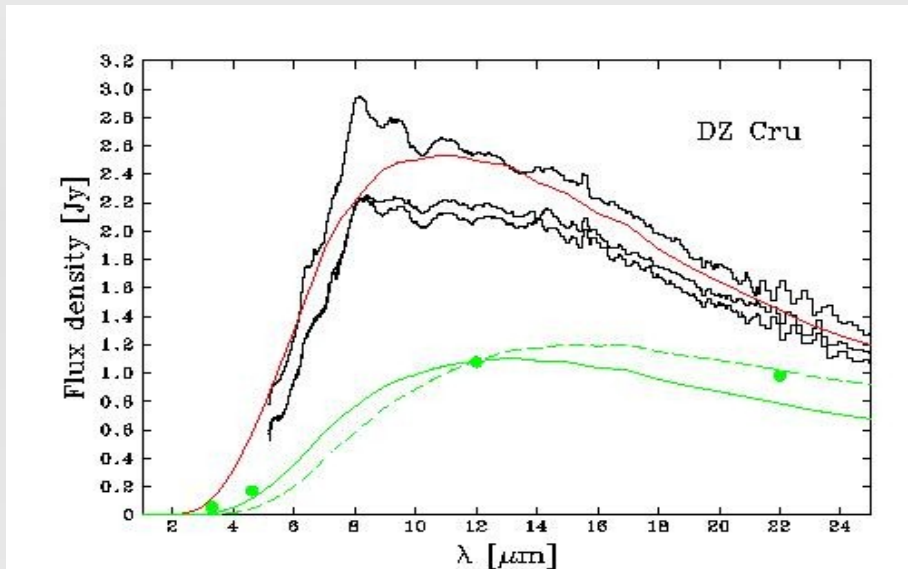
- Dust SEDs



**WISE SED of DZ Cru**

# WISE survey of novae

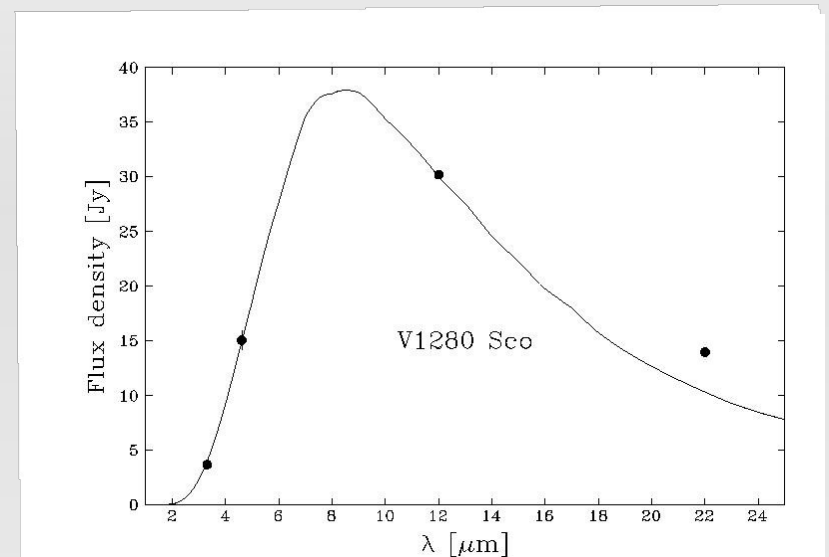
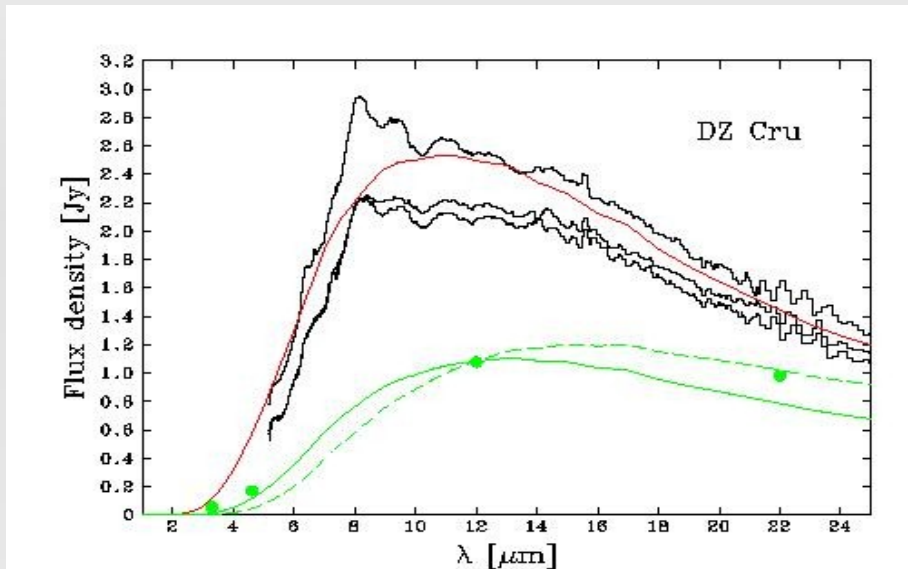
## Dust SEDs



WISE SED compared with *Spitzer*  
(Evans et al., MNRAS, 406, L85)

# WISE survey of novae

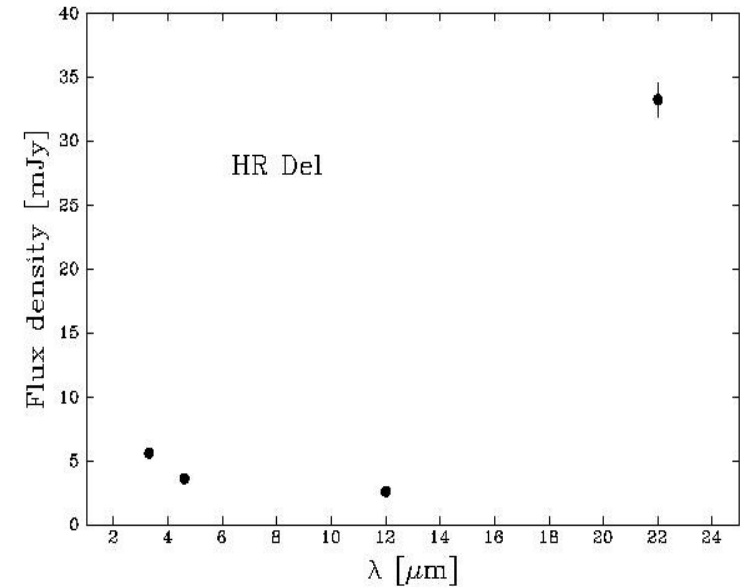
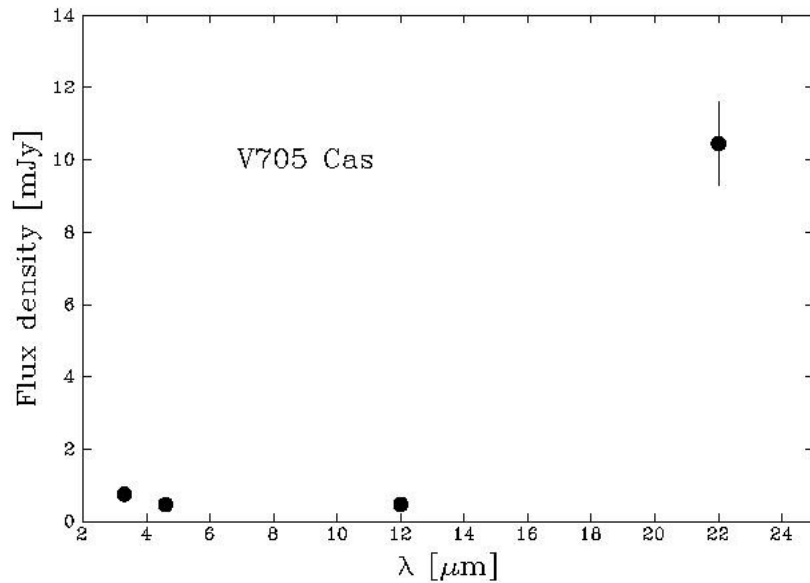
## Dust SEDs



## WISE SED of V1280 Sco



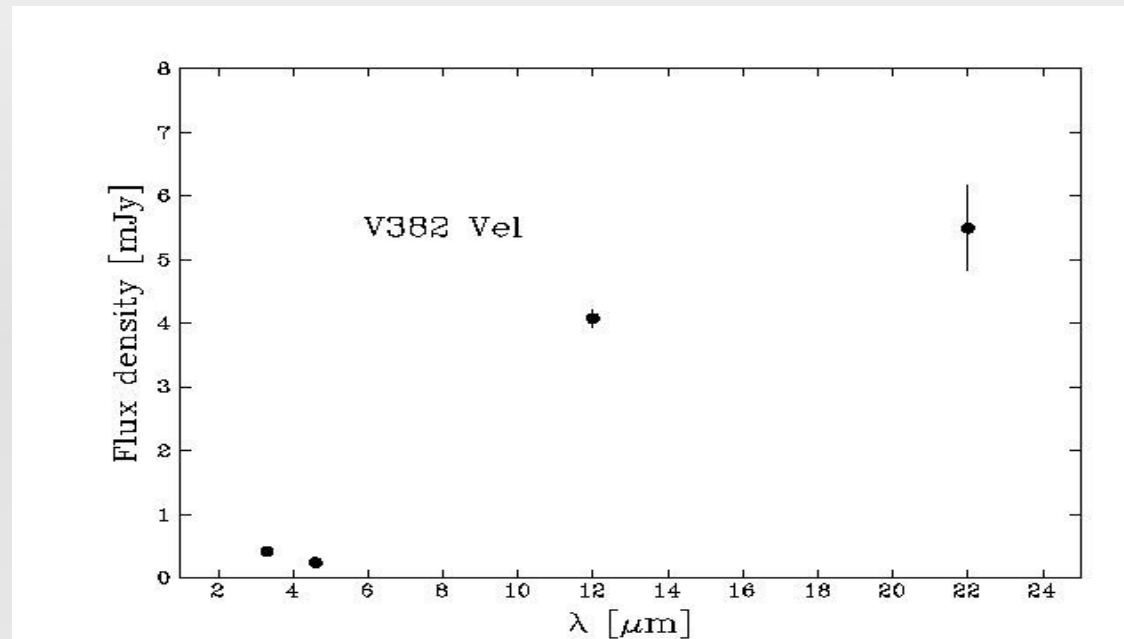
# WISE survey of novae



V705 Cas and HR Del – [OIV]25.9 micron emission

“Emission Line” SEDs

# WISE survey of novae

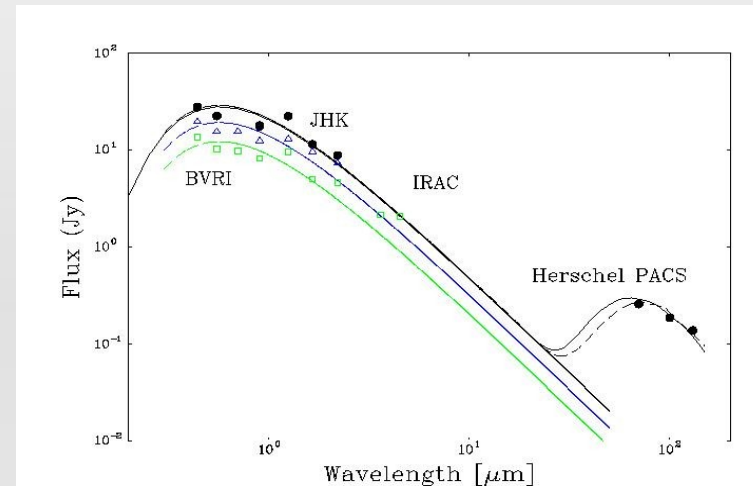
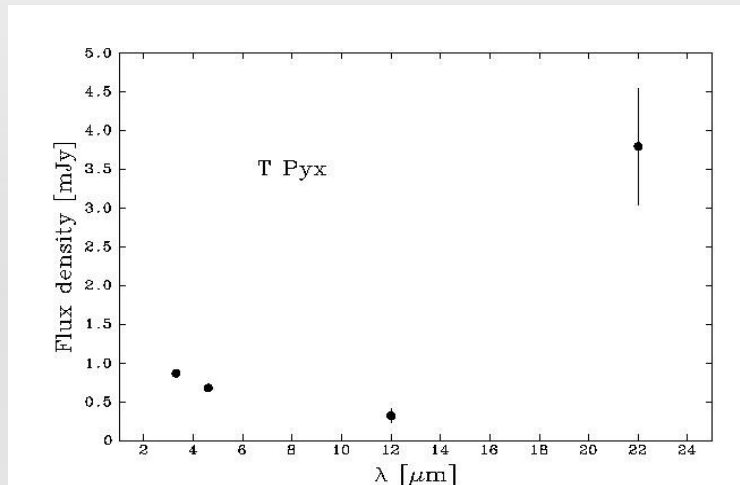


**V382 Vel – a neon nova.**

**Likely [NeII]12.8, [NeIII]15.5, [OIV]25.9**

**“Emission Line” SEDs**

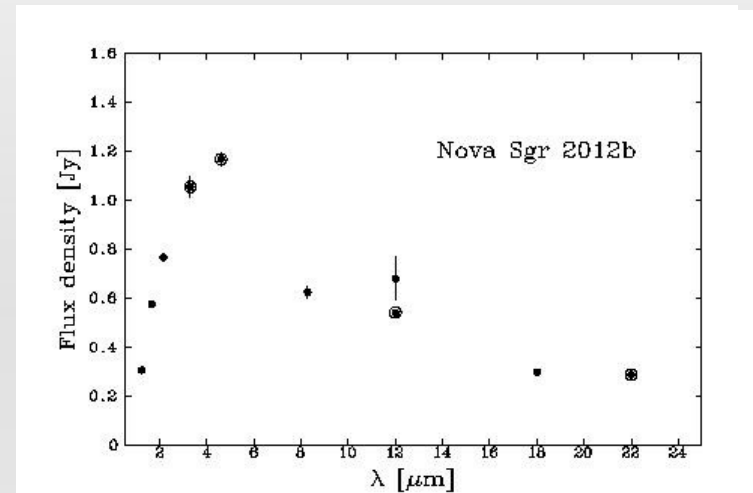
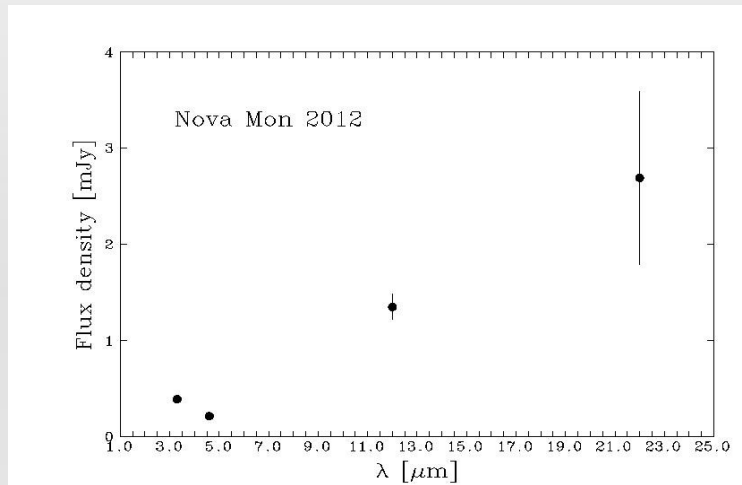
# WISE survey of novae



Left: WISE view of T Pyx  
Right: SMARTS/MtAbu/Spitzer/Herschel  
view of the 2011 eruption

Progenitors of post-WISE novae

# WISE survey of novae



Left: Nova Mon 2012

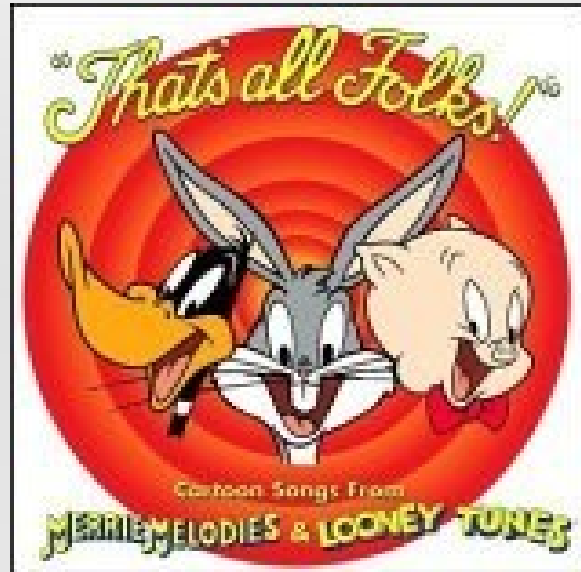
Right: Nova Sgr 2012b 2MASS/MSX/IRAS/WISE

Progenitors of post-WISE novae

# WISE survey of novae

- ▶ The WISE survey will (as was the *IRAS* survey 30 years ago – see Harrison & Gehrz, *AJ*, 96, 1001; 101, 587, 103, 243, 108, 1899) be a valuable resource for investigating the circumstellar environments of mature novae
- ▶ WISE does not have the long wavelength coverage of *IRAS* but clearly beats *IRAS* for sensitivity and spatial resolution
- ▶ It has the potential to provide unprecedented information about nova progenitors

# WISE survey of novae



With thanks to Alex d'Angelo for doing an early trawl